

Baryogenesis from a Majorana Fermion Coupled to Quarks

Shrihari Gopalakrishna



IMSc, Chennai, and HBNI

[2211.12115; 2311.14636; 2411.13231][PRD]
(with Rakesh Tibrewala, LNMIIT, Jaipur)

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Talk Outline

- Baryogenesis from dim-6 Vector-Vector (VV) eff. op: $\frac{1}{\Lambda^2} (D\Gamma^\mu D) (\chi\gamma_\mu U)$
 - ▶ χ Majorana mass breaks B
- Baryon asymmetry in \mathcal{X} decay and scattering
 - ▶ UV completion details
- Solve Boltzmann Equations
 - ▶ match to observed Baryon Asymmetry of the Universe (BAU)
- Terrestrial probes
 - ▶ neutron-antineutron ($n - \bar{n}$) oscillation

Baryogenesis basics

- Today we only have matter in the (observable) Universe

$$\eta_B \equiv \frac{n_B}{n_\gamma} = \frac{(n_b - n_{\bar{b}})}{n_\gamma} = 6 \times 10^{-10} \frac{\Omega_b h^2}{0.0222} ; \quad Y_B \equiv \frac{n_B}{s} = 0.85 \times 10^{-10}$$

- ▶ What happened to all the anti-matter?
- ▶ Particle physics explanation is most compelling
- Sakharov conditions for Baryogenesis
 - ▶ B violation
 - ▶ C and CP violation
 - ▶ Departure from Thermal Equilibrium
- When did baryogenesis happen?
 - ▶ EW, Leptogenesis, GUT, ...

Our study: baryogenesis in \mathcal{X} Decay and Scattering

$$\mathcal{L}_{Dirac M} \supset -M_{\mathcal{X}} \bar{\mathcal{X}}\mathcal{X} - M_U \bar{U}U - M_D \bar{D}D$$

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$$\mathcal{L}_{Dirac M} \supset -M_\chi \bar{\chi}\chi - M_U \bar{U}U - M_D \bar{D}D$$

$$\mathcal{L}_{VV} \supset \frac{g_{L,R}}{\Lambda^2} [\bar{D}^c \gamma_\mu D] [\bar{\chi} \gamma^\mu P_{L,R} U]$$

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VV interaction: χ with $Q = (U, D)$

$$B(Q) = 1/3, B(\chi) = +1$$

Dirac mass: M_χ

Majorana mass: $\tilde{M}_{L,R}$

Decay and Scattering processes:

- $\mathcal{X} \leftrightarrow UDD$ vs. $\mathcal{X} \leftrightarrow U^c D^c D^c$
- $\mathcal{X} Q^c \leftrightarrow QQ$ vs. $\mathcal{X} Q \leftrightarrow Q^c Q^c$
- $UDD \leftrightarrow U^c D^c D^c$ vs. $U^c D^c D^c \leftrightarrow UDD$

Our study: baryogenesis in \mathcal{X} Decay and Scattering

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- Sakharov conditions

▶ B violation: **Majorana mass** $\tilde{M}_{L,R}$

▶ C and CP violation: complex $g_{L,R}, \tilde{M}_{L,R}$:

▶ Departure from Thermal Equilibrium: Hubble expansion

- When did Baryogenesis happen?

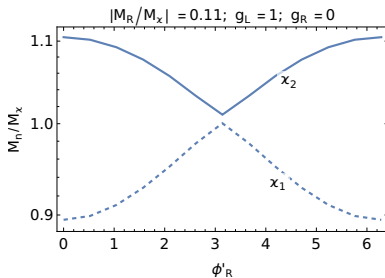
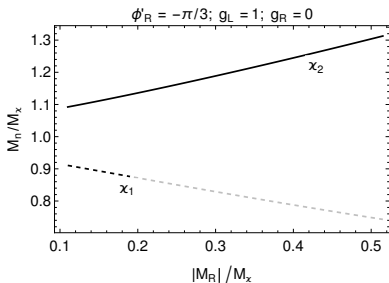
▶ At scale M_χ

Majorana mass \implies Dirac χ to Majorana pair \mathcal{X}_n

Majorana mass \tilde{M} splits Dirac χ into a **pair** of Majorana fermions

$\mathcal{X}_n = (\mathcal{X}_1, \mathcal{X}_2)$ with **indefinite baryon number**, mass eigenvalues M_n

$$\chi = (U_{1n} P_L + U_{2n}^* P_R) \mathcal{X}_n$$



$$\mathcal{L}_{\text{int}}^{\text{VV}} = \frac{\epsilon^{abc}}{2\Lambda^2} \left\{ \left[\overline{D_b^c} \gamma^\mu D_a \right] \left[\tilde{\mathcal{X}}_n G_{V\mu}^n U_c \right] \right\}$$

$$G_{V\mu}^n \equiv \gamma_\mu (g_L U_{1n}^* P_L + g_R U_{2n} P_R)$$

Baryon asymmetry in Decay $\mathcal{X}_n \rightarrow UDD$

Amplitude (tree+loop): Process $\mathcal{A} = \mathcal{A}_0 + \mathcal{A}_1$; Conjugate process $\mathcal{A}^c = \mathcal{A}_0^c + \mathcal{A}_1^c$

Decay rate: $\Gamma = \frac{1}{2M_n} \int d\Pi_3 |\mathcal{A}|^2$; $\Gamma^c = \frac{1}{2M_n} \int d\Pi_3 |\mathcal{A}^c|^2$

$$\mathcal{A}_B \equiv \frac{\Gamma - \Gamma^c}{\Gamma + \Gamma^c}$$

CP violation needs nonzero **weak phase** (ϕ) and **strong phase** ($\delta = \pi/2$)

$$\mathcal{A}_B^n \propto \left| \begin{array}{c} \begin{array}{c} (q_2, b) \quad \bar{D}^c \\ \nearrow \\ \chi_n \\ \leftarrow p_n \quad \rightarrow (q_U, c) \\ \text{U} \end{array} \\ A_0 \end{array} \right. + \left. \begin{array}{c} \chi_n \\ \text{Loop} \\ i\hat{A}_1 \end{array} \right|^2 - \left| \begin{array}{c} \begin{array}{c} \bar{D} \\ \nearrow \\ \chi_n \\ \leftarrow U^c \\ \text{D}^c \end{array} \\ A_0^c \end{array} \right. + \left. \begin{array}{c} \chi_n \\ \text{Loop} \\ i\hat{A}_1^c \end{array} \right|^2$$

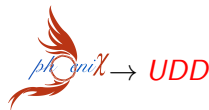
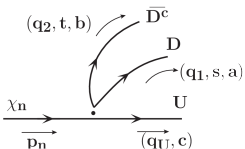
Interference term $\hat{A}_{01} \equiv \hat{A}_1 \mathcal{A}_0^*$, $\hat{A}_{01}^c \equiv \hat{A}_1^c \mathcal{A}_0^{c*}$

$$\Delta \hat{\Gamma}_{01} = \frac{1}{2M_n} \int d\Pi_3 \text{Im}(\hat{A}_{01} - \hat{A}_{01}^c); \quad \mathcal{A}_B \approx -\frac{\Delta \hat{\Gamma}_{01}}{\Gamma_0}$$

[S.Gopalakrishna, R. Tibrewala: 2211.12115]

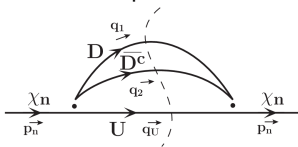
χ_n Decay at tree-level

At tree level: $\mathcal{A}_0 \propto$



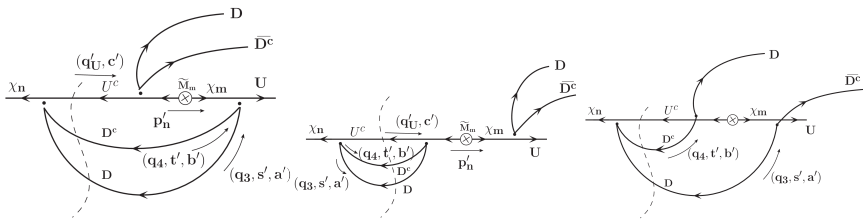
Tree-level decay width as a loop:

$$\Gamma_0 \propto |\mathcal{A}_0|^2 =$$

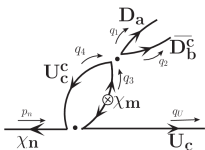


$$\text{Cutkosky rule: } 1/(k^2 - m^2 + i\epsilon) \rightarrow -2\pi i \delta(k^2 - m^2)$$

χ_n Decay at loop level



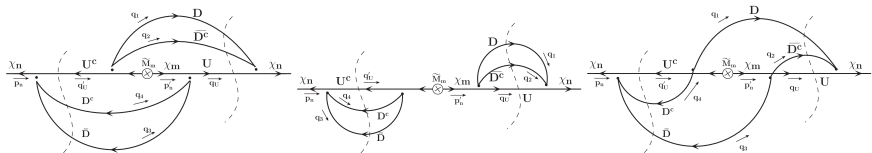
Single Operator Contributions



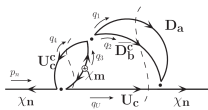
Multiple Operator Contribution

Arrow clash \leftrightarrow Majorana mass \leftrightarrow B violation

χ_n Decay tree-loop interference term



Single Operator Contributions



Multiple Operator Contribution

Compute \hat{A}_{01} :

- Dirac traces and the matrix element
- Loop integral
- Fold in phase-space (as loop integral) and integrate

Benchmark points: BP-A, BP-B, BP-C

Table 1: The Benchmark Points (BP) parameters, and the $g, M_\chi/\Lambda$ scaling factors.

	\hat{M}_Q	$\hat{\Gamma}_0$ ($\times g ^2 M_\chi^4/\Lambda^4$)	$\Delta \hat{\Gamma}_{01}$ ($\times \text{Im}(g^4) M_\chi^6/\Lambda^6$)	$\langle \hat{\sigma}_0 v \rangle$ ($\times g ^2 M_\chi^4/\Lambda^4$)	$\langle \Delta \hat{\sigma}_{01} v \rangle$ ($\times \text{Im}(g^4) M_\chi^6/\Lambda^6$)	$\langle \hat{\Gamma}'_0 \rangle$ ($\times g ^4 M_\chi^8/\Lambda^8$)
BP-A	1/4	1.7×10^{-6}	$-1.4 \times 10^{-9} M_\chi^2/\Lambda^2$	12.7	$-0.04 M_\chi^2/\Lambda^2$	0.4
BP-B	1/20	2×10^{-5}	-6×10^{-11}	3.5	-0.3	6.4
BP-C	1/20	10^{-4}	-10^{-7}	12.7	-10^{-5}	0.4

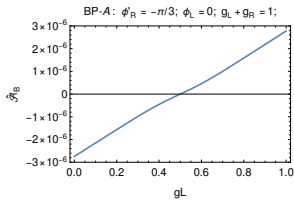
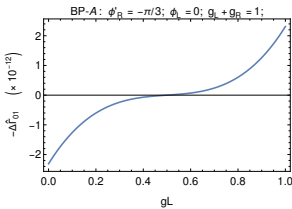
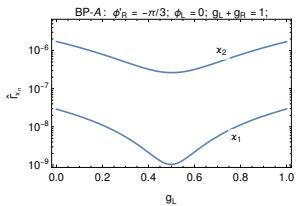
	BP-1A	BP-1B	BP-1C	BP-2A	BP-2B	BP-2C	BP-3A	BP-3B	BP-3C
M_χ (GeV)	10^{12}			10^9			10^6		
M_χ/Λ	1/10	1/10	1/10	1/10	1/10	1/10	1/10	1/50	1/10
g	0.25	0.008	0.048	0.065	0.009	0.065	0.024	0.022	0.087
Modified parameters							$(2/5) \langle \hat{\sigma}_0 v \rangle$ $(5/2) \langle \Delta \hat{\sigma}_{01} v \rangle$		$(1/30) \hat{\Gamma}_0$

$M_\chi = 1, \tilde{M}_L = 0.1, |\tilde{M}_R| = 0.11, \phi'_R = -\pi/3$ (all masses scaled with M_χ)

Compute (numerically) the integrals, baryon asymmetry, solve Boltzmann Eqn.

(*FeynCalc, Mathematica*)

Decay \mathcal{A}_B : Single Operator contrib: BP-A

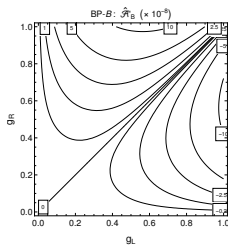
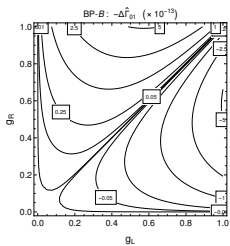
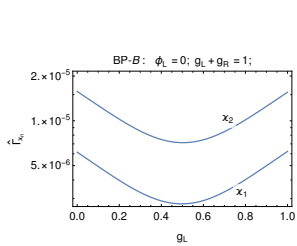


Scaling factors (M_χ^5/Λ^4 , M_χ^9/Λ^8 , M_χ^4/Λ^4)

\mathcal{A}_B is phenomenologically interesting

[S.Gopalakrishna, R. Tibrewala: 2311.14636]

Decay \mathcal{A}_B : Multiple operator contrib: BP-B



Scaling factors ($M_\chi^5/\Lambda^4, M_\chi^7/\Lambda^6, M_\chi^2/\Lambda^2$)

\mathcal{A}_B is phenomenologically interesting

UV completion example

In loops or phase-space, if $p \gtrsim \Lambda$, UV completion relevant.

- Particularly relevant in scattering

Example: introduce color triplet, vector ξ_μ^a with $Q(\xi) = -2/3$:

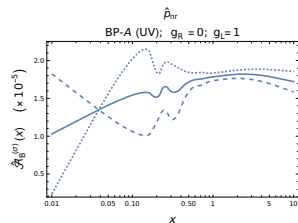
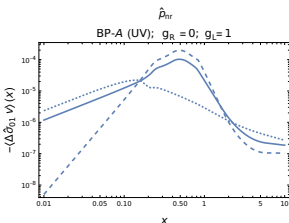
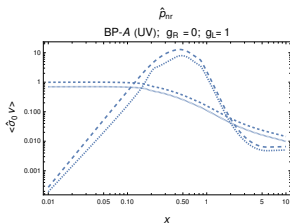
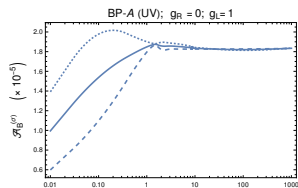
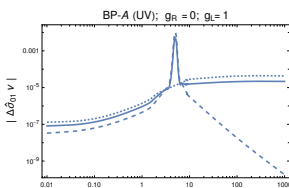
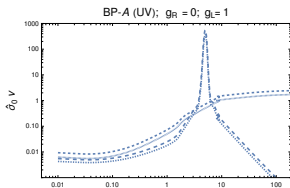
$$\mathcal{L}_{UV}^{(A)} \supset -\frac{1}{2} \epsilon^{abc} [\overline{D_b^c} \tilde{g} \gamma^\mu D_a] \xi_\mu^{c*} - [\tilde{\chi} \gamma^\mu (g_L P_L + g_R P_R) U_c] \xi_\mu^c + \text{h.c.}$$

Scattering B Asym ($\mathcal{X}\bar{Q} \rightarrow QQ$ vs. $\mathcal{X}Q \rightarrow \bar{Q}\bar{Q}$)

$$\text{SC-1: } \mathcal{X}_n(p_n)\bar{D}(k_i) \rightarrow D(q_1)U(q_2)$$

$$\text{SC-2: } \mathcal{X}_n(p_n)\bar{U}(k_i) \rightarrow D(q_1)\bar{D}^c(q_2)$$

$$\hat{\mathcal{A}}_B^{(\sigma)} \equiv \frac{\langle \sigma v \rangle - \langle \sigma^c v \rangle}{\langle \sigma v \rangle + \langle \sigma^c v \rangle} \approx -\frac{\langle \Delta \hat{\sigma}_{01} v \rangle}{\langle \sigma_0 v \rangle}$$



Scaling factors ($M_\chi^2/\Lambda^4, M_\chi^6/\Lambda^8, M_\chi^4/\Lambda^4$) ($x = M/T$)

Boltzmann Equations: expanding Universe at T

Early Universe at temperature T : time $t \approx (0.301/\sqrt{g_*})M_{Pl}/T^2$

Hubble expansion rate $H(T) = 1.66\sqrt{g_*}T^2/M_{Pl}$

Track number densities: $n_\chi(t), n_Q(t), n_{\bar{Q}}(t)$; $n_B(t) = (n_Q - n_{\bar{Q}})/3$

- $n_\chi(t), n_B(t)$: $n_B(t_b) = 0, n_B(t_e) \sim BAU$

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Boltzmann Equations

$$\frac{d}{dt}n_\chi + 3Hn_\chi = -\langle\Gamma\rangle(n_\chi - n_\chi^{(0)}) + \dots$$

$$\frac{d}{dt}n_B + 3Hn_B = \langle\Delta\Gamma\rangle(n_\chi - n_\chi^{(0)}) - \langle\Gamma\rangle n_\chi^{(0)} n_B/n_Q^{(0)} \dots$$

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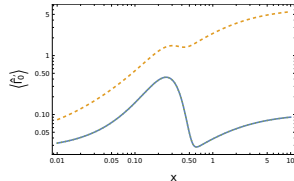
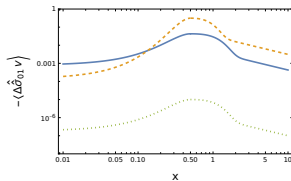
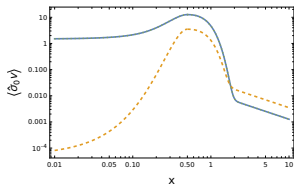
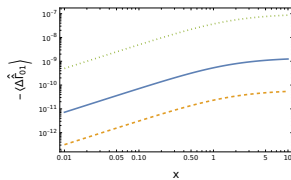
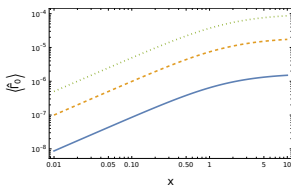
$$\frac{d}{dt}n_B + 3Hn_B = \langle\Delta\Gamma\rangle(n_\chi - n_\chi^{(0)}) - \langle\Gamma\rangle n_\chi^{(0)} n_B/n_Q^{(0)} \dots$$

Equivalently solve BE for $Y_\chi = n_\chi/s$, $Y_B = n_B/s$

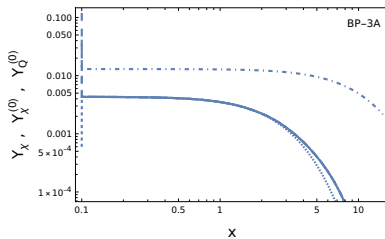
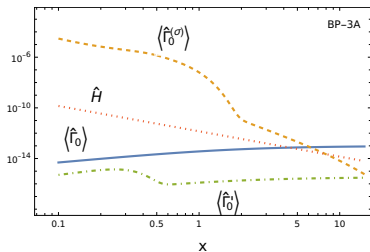
vs. $t \rightarrow$ temperature $T \rightarrow x = M_\chi/T$

- $Y_\chi(x), Y_B(x)$; $Y_B(x_b) = 0, Y_B(x_e) \sim Y_B^{obs}$

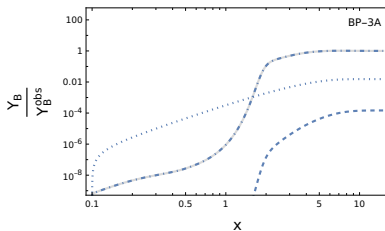
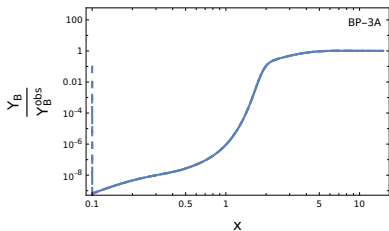
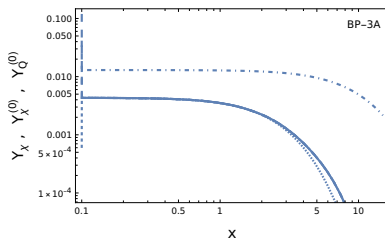
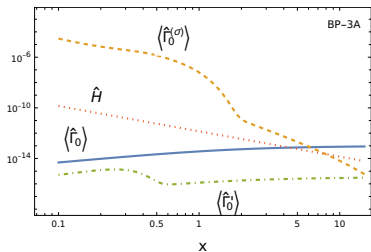
Thermally averaged rates



Boltzmann Equation Solution: BAU in BP-3A

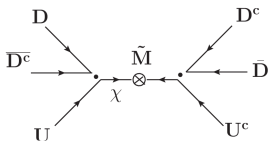


Boltzmann Equation Solution: BAU in BP-3A



Terrestrial probe: neutron-antineutron ($n - \bar{n}$) oscillation

$\Delta B = 2$ process



$$\tau_{n-\bar{n}} \geq 4.7 \times 10^8 \text{ s at 90 \% C.L. [SuperK, 2021]} \implies (\Delta m_{n-\bar{n}}) \leq 10^{-34} \text{ GeV}$$

$$g^2 s_{\text{eff}}^2 \frac{\Lambda_{\text{QCD}}^6 \langle \bar{n} | \hat{Q}_i | n \rangle}{\Lambda^4 M_\chi} \lesssim 10^{-34} \text{ GeV} \implies \Lambda \gtrsim s_{\text{eff}}^{2/5} 10^3 \text{ TeV (for } g \sim \mathcal{O}(1))$$

$$Q \equiv \epsilon^{abc} \epsilon^{a'b'c'} (u_{c'}^{\bar{c}} \gamma^\nu \gamma^\mu P_{L,R} u_c) (d_b^{\bar{c}} \gamma_\mu d_a) (d_{b'}^{\bar{c}} \gamma_\nu d_{a'})$$

Lattice computation of the matrix element not available;

for now use scalar operator in the Fierz rearrangement

$$10^{\kappa_{n\bar{n}}} \equiv \frac{s_{\text{eff}}^2}{e_{n\bar{n}}} = 2.5 \times 10^{-31} \left(\frac{13}{\langle \hat{Q} \rangle} \right) \left(\frac{M_\chi}{1 \text{ GeV}} \right)^5 \left[\frac{1}{g (M_\chi / \Lambda)} \right]^2$$

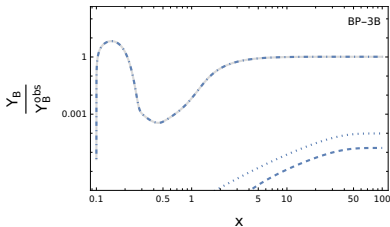
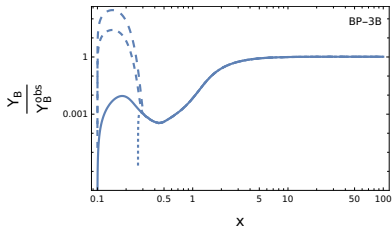
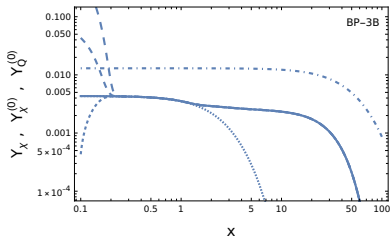
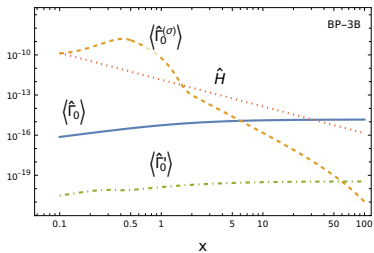
Conclusions

- Majorana \mathcal{X} with dim-6 VV interaction to UDD studied
 - ▶ Sakharov conditions satisfied
- Baryon Asymmetry due to decay and scattering computed
- Boltzmann Equations solved to match to observed BAU
 - ▶ $M_{\mathcal{X}} \in (10^4, 10^{16}) \text{ GeV}$
- Terrestrial probes
 - ▶ $n - \bar{n}$ oscillation
 - ▶ Flavor probes?

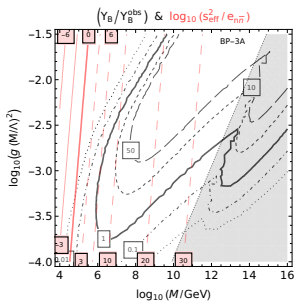
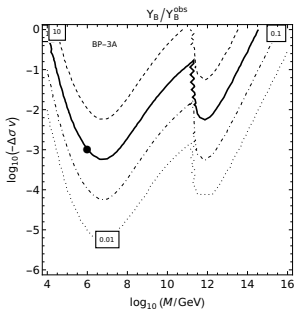
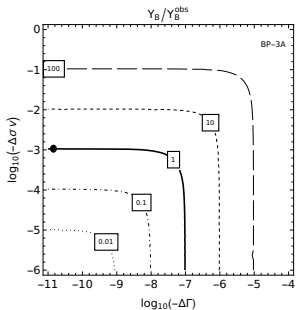
Backup Slides

BACKUP SLIDES

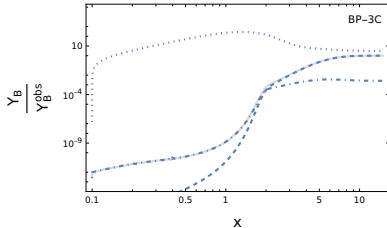
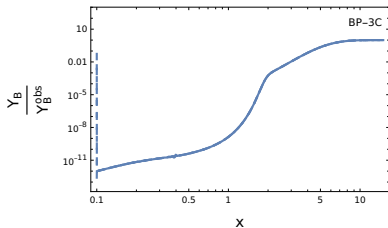
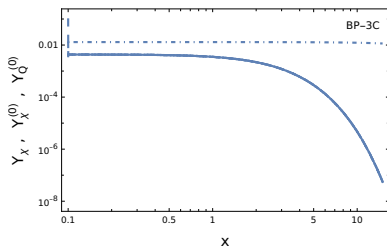
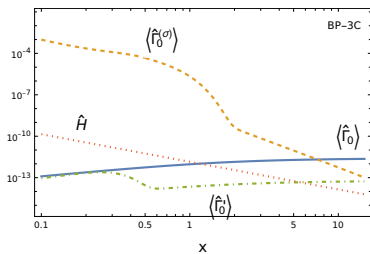
BAU in BP-3B



BAU in BP-3A



BAU in BP-3C



BAU in BP-3C

