

Zero Temperature Dissipation & Holography

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National Strings Meeting - 2015



PB & B. Sathiapalan (on going)

Outline

- 1 Introduction
- 2 Langevin Dynamics From Holography
- 3 Dissipation
- 4 Conclusions

Introduction

Motivation

- ▶ AdS/CFT has been serving as a theoretician's best tool in studying strongly coupled systems analytically.
- ▶ Its predictions are mostly qualitative in nature, but they can be quantitative too (e.g, $\frac{\eta}{s} = \frac{1}{4\pi}$).
- ▶ The duality has glued many phenomena appearing in apparently different branches of physics together.
- ▶ Studying Brownian motion of a heavy particle using classical gravity technique is one such example.

Introduction

Langevin Dynamics

- The Langevin equation

$$M \frac{d^2 x}{dt^2} + \gamma \frac{dx}{dt} = \xi(t) \quad (1)$$

$$\text{with } \langle \xi(t) \xi(t') \rangle = \Gamma \delta(t - t') \quad (2)$$

Introduction

Langevin Dynamics

► The Langevin equation

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The **generalized Langevin equation** for a heavy particle under noise ξ

$$M_0 \frac{d^2 x(t)}{dt^2} + \int_{-\infty}^t dt' \, G_R(t, t') \dot{x}(t') = \xi(t) \quad \langle \xi(t) \xi(t') \rangle = i G_{\text{sym}}(t, t') \quad (3)$$

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Langevin Dynamics

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The **generalized Langevin equation** for a heavy particle under noise ξ

$$\left[-M_0 \omega^2 + G_R(\omega) \right] x(\omega) = \xi(\omega) \quad \langle \xi(-\omega) \xi(\omega) \rangle = i G_{\text{sym}}(\omega) \quad (3)$$

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- Expanding $G_R(\omega)$ for small frequencies

$$G_R(\omega) = -\Delta M \omega^2 - i\gamma\omega + \dots$$

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$$G_R(\omega) = -\Delta M \omega^2 - i\gamma\omega + \dots$$

- **Fluctuation-Dissipation** relation

$$i G_{\text{sym}}(\omega) = -(1 + 2n_B) \text{Im } G_R(\omega) \quad (4)$$

Introduction

Retarded Green's Function

- For small frequency

$$G_R(\omega) = -i\gamma \omega - \Delta M \omega^2 - i\rho \omega^3 + \dots$$

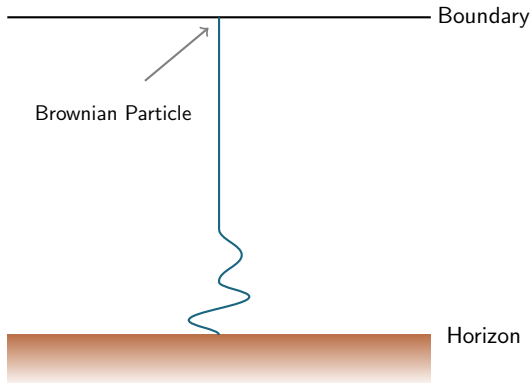
- Dimensional analysis

$$G_R(\omega) \sim [M]^3 \Rightarrow \begin{cases} \gamma \sim T^2 \\ \Delta M \sim T \\ \rho \sim T^0 \end{cases}$$

- At zero temperature : $G_R(\omega) = -i\rho \omega^3 \iff$ Dissipative!

Langevin Dynamics From Holography

Idea & Set-up



J. de Boer et al; Son & Teaney (2009)

Langevin Dynamics From Holography

Retarded Green's function : Holographic prescription

- **Strongly** coupled field theory \Leftrightarrow **Weakly** coupled gravity.

$$\left\langle \exp \left(\int_{S^d} \phi_0^i \mathcal{O}_i \right) \right\rangle_{\text{CFT}} = Z_{\text{QG}}(\phi_0^i) \quad (5)$$

Langevin Dynamics From Holography

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- Real time **retarded** Green's function for **scalar field** theory can be obtained by choosing **ingoing** boundary condition at the horizon.

$$G_R(k) = K \sqrt{-g} g^{rr} f_{-k}(r) \partial_r f_k(r) \Big|_{\text{Boundary}} \quad (6)$$

Son & Starinets (2002)

Langevin Dynamics From Holography

Retarded Green's function : Holographic prescription

- **Strongly** coupled field theory \Leftrightarrow **Weakly** coupled gravity.

$$\left\langle \exp \left(\int_{S^d} \phi_0^i \mathcal{O}_i \right) \right\rangle_{\text{CFT}} = Z_{\text{QG}}(\phi_0^i)$$

- Real time **retarded** Green's function for **our case** can be obtained by choosing **ingoing** boundary condition at the horizon.

$$G_R(\omega) = T_0(r) \left. f_{-\omega}(r) \partial_r f_{\omega}(r) \right|_{\text{Boundary}} \quad (7)$$

Langevin Dynamics From Holography

Computing $G_R(\omega)$: 4 Steps

- 1 **Solve** the EOM for the string in that non-trivial background

$$f_\omega(r) = C_1 f_\omega^{(1)}(r) + C_2 f_\omega^{(2)}(r) \quad (8)$$

- 2 Impose **ingoing wave** boundary condition at the horizon

$$f_\omega^R(r) := C_1 f_\omega^{(1)}(r) + \cancel{C_2 f_\omega^{(2)}(r)} \quad (9)$$

- 3 Properly **normalize** it at the boundary

$$F_\omega^R(r) := \frac{f_\omega^R(r)}{f_\omega^R(r_B)} \quad (10)$$

- 4 Compute the **boundary action** and take functional derivative w.r.t **"source"** to obtain the retarded Green's function $G_R^0(\omega)$.

Langevin Dynamics From Holography

An Example : Brownian Motion in 1+1 Dimensions

String in BTZ

- String **EOM**

$$f''_{\omega}(r) + \frac{2(2r^2 - 4\pi^2 T^2 L^4)}{r(r^2 - 4\pi^2 T^2 L^4)} f'_{\omega}(r) + \frac{L^4 \omega^2}{(r^2 - 4\pi^2 T^2 L^4)^2} f_{\omega}(r) = 0 \quad (11)$$

- Two **independent solutions**

$$f_{\omega}(r) = C_1 \frac{P_1^{\frac{i\omega}{2\pi T}}\left(\frac{r}{2\pi TL^2}\right)}{r} + C_2 \frac{Q_1^{\frac{i\omega}{2\pi T}}\left(\frac{r}{2\pi TL^2}\right)}{r} \quad (12)$$

- **Ingoing & normalized** solution

$$F_{\omega}^R(r) = \frac{\frac{P_1^{\frac{i\omega}{2\pi T}}\left(\frac{r}{2\pi TL^2}\right)}{r}}{\frac{P_1^{\frac{i\omega}{2\pi T}}\left(\frac{r_B}{2\pi TL^2}\right)}{r_B}} \quad (13)$$

Langevin Dynamics From Holography

An Example : Brownian Motion in 1+1 Dimensions

String in BTZ (cont.)

- ▶ Retarded propagator can be read off from the on-shell action

$$\begin{aligned} G_R(\omega) &\equiv G_R^0(\omega) + \frac{\mu \omega^2}{2\pi} \\ &= \frac{\mu \omega}{2\pi} \frac{(\omega^2 + 4\pi^2 T^2)}{(\omega + i \frac{\mu}{\sqrt{\lambda}})} \end{aligned} \quad PB \& B. Sathiapalan (2014)$$

- ▶ Viscous drag

$$\gamma = 2\sqrt{\lambda}\pi T^2$$

- ▶ Higher order “dissipation coefficient”

$$\rho = \frac{\sqrt{\lambda}}{2\pi} - \frac{2(\sqrt{\lambda})^3 \pi T^2}{\mu^2} \xrightarrow{T \rightarrow 0} \frac{\sqrt{\lambda}}{2\pi}$$

Dissipation

Zero temperature dissipation is physical

- ✓ It is **finite** and therefore no need to renormalize by adding counterterms.
- ✓ It cannot be renormalized away in the boundary theory by any **hermitian** counter term.
- ✓ Quark moving at **constant velocity** doesn't feel any drag at $T = 0$.
- ▶ **Explanation** : This zero temperature dissipation is due to **radiation of accelerated charged** particle. Remarkably the dissipation in this highly non-linear boundary theory is given by simple “**Abraham-Lorentz**”-like formula for radiation reaction in classical electrodynamics!

$$\vec{F}_{rad} = \frac{\sqrt{\lambda}}{2\pi} \dot{\vec{a}}$$

Dissipation

Dissipation at $T = 0$

String in pure AdS_{d+1}

► String EOM

$$f''_{\omega}(r) + \frac{4}{r} f'_{\omega}(r) + \frac{L^4 \omega^2}{r^4} f_{\omega}(r) = 0 \quad (14)$$

► Ingoing & normalized solution

$$F_{\omega}^R(r) = \frac{r_B}{r} \frac{e^{+i\frac{L^2\omega}{r}}(r - iL^2\omega)}{e^{+i\frac{L^2\omega}{r_B}}(r_B - iL^2\omega)} \quad (15)$$

► Retarded Green's function

$$G_R(\omega) := G_R^0(\omega) + \frac{\mu\omega^2}{2\pi} = \frac{\mu\omega^3}{2\pi} \frac{1}{(\omega + i\frac{\mu}{\sqrt{\lambda}})} \quad (16)$$

► Higher order “dissipation coefficient”

$$\rho = \frac{\sqrt{\lambda}}{2\pi} \quad (\text{Identical!}) \quad (17)$$

Dissipation

Is the zero temperature dissipation universal?

Dissipation

Is the zero temperature dissipation universal?

String in AdS_5 -BH

- ▶ String EOM can *not* be solved exactly!

$$f''_{\omega}(r) + \frac{4r^3}{(r^4 - \pi^4 T^4 L^8)} f'_{\omega}(r) + \frac{\omega^2 L^4 r^4}{(r^4 - \pi^4 T^4 L^8)^2} f_{\omega}(r) = 0 \quad (18)$$

- ▶ Ingoing & normalized ansatz

$$F_{\omega}^R(r) = \left(1 - \frac{\pi^4 T^4 L^8}{r^4}\right)^{-i\frac{\Omega}{4}} (1 - i\Omega f_1(r) - \Omega^2 f_2(r) + i\Omega^3 f_3(r) + \dots) \quad (19)$$

- ▶ The limit we are interested in

$$\omega, T \rightarrow 0 \quad \text{and} \quad \Omega := \frac{\omega}{\pi T} = \text{fixed} \quad (20)$$

- ▶ Solving $f_1(r)$, $f_2(r)$ and $f_3(r)$ perturbatively and following the same procedure

$$G_R(\omega) \equiv G_R^0 + \frac{\mu\omega^2}{2\pi} = -i \frac{\sqrt{\lambda}}{2\pi} \left(\frac{\pi - \text{Log}4}{4} \right) \omega^3 \quad (21)$$

Dissipation

A phase transition at $T = 0$?

Dissipation as $T \rightarrow 0$

- ▶ Calculating $G_R(\omega)$ in AdS_5 -Schwarzschild black hole bulk geometry

$$\rho = \frac{(\pi - \log 4)}{4} \frac{\sqrt{\lambda}}{2\pi}$$

Dissipation at $T = 0$

- ▶ Calculating $G_R(\omega)$ in pure AdS_5 bulk geometry

$$\rho = \frac{\sqrt{\lambda}}{2\pi}$$

Dissipation

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- ▶ Calculating $G_R(\omega)$ in pure AdS_5 bulk geometry

$$\rho = \frac{\sqrt{\lambda}}{2\pi}$$

- ▶ **Conclusion** : Possibly due to “deconfinement” transition at $T=0$.

Dissipation

Dissipation at finite density

- ▶ **Reissner-Nordström (RN)** black hole in asymptotically AdS space time

$$ds^2 = \frac{L_{d+1}^2}{z^2} (-f(z) dt^2 + d\vec{x}^2) + \frac{L_{d+1}^2}{z^2} \frac{dz^2}{f(z)} \quad (22)$$

where,

$$f(z) = 1 + Q^2 z^{2d-2} - Mz^d$$

$$A_t(z) = \mu \left(1 - \frac{z^{d-2}}{z_0^{d-2}} \right)$$

- ▶ We'll define a new length scale z_* where $Q := \sqrt{\frac{d}{d-2}} \frac{1}{z_*^{d-1}}$
- ▶ There are two possibilities :
 - ▶ Extremal BH ($T = 0$)
 - ▶ Non-extremal BH ($T \neq 0$)

Dissipation

Dissipation at finite density

Finite density and zero temperature

- ▶ The string EOM

$$x''_{\omega}(z) + \frac{\frac{d}{dz}\left(\frac{f(z)}{z^2}\right)}{\frac{f(z)}{z^2}} x'_{\omega}(z) + \frac{\omega^2}{[f(z)]^2} x_{\omega}(z) = 0 \quad (23)$$

- ▶ **Subtlety** : At zero temperature the $f(z)$ has a **double zero** at the horizon. Thus this singular term dominates at the horizon irrespective of however small ω we choose.
- ▶ **Matching technique** : Isolate the 'singular' near horizon region and treat ω perturbatively "outside".
 - ▶ Inner/IR region : $AdS_2 \times \mathbb{R}^{d-1}$
 - ▶ Outer/UV region : Full RN-AdS background

Hong Liu et al. (2009)

Dissipation

Dissipation at finite density

- ▶ Matching the solutions in two regions near $z = z_*$ we obtain

$$\begin{aligned} G_R(\omega) &:= \frac{b_+ + \mathcal{G}_R(\omega) z_* b_-}{a_+ + \mathcal{G}_R(\omega) z_* a_-} \\ &= \frac{(b_+^{(0)} + \omega^2 b_+^{(2)} + \dots) - i\omega(b_-^{(0)} + \omega^2 b_-^{(2)} + \dots) z_*}{(a_+^{(0)} + \omega^2 a_+^{(2)} + \dots) - i\omega(a_-^{(0)} + \omega^2 a_-^{(2)} + \dots) z_*} \end{aligned} \quad (24)$$

- ▶ The leading-order Green's function

$$G_R^{(0)}(\omega) = -\frac{i \omega z_*}{(1 + i \omega z_* a_-^{(0)})} \quad (25)$$

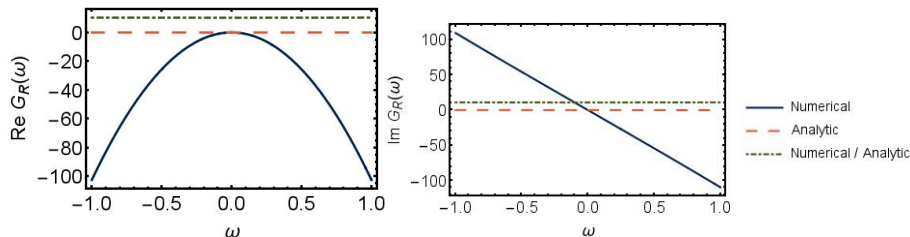
- ▶ For small frequency

$$G_R^{(0)}(\omega) \approx -i \omega z_* (1 - i \omega z_* a_-^{(0)}) \quad (26)$$

Dissipation

Dissipation at finite density

Choosing $a_-^{(0)} = 1$

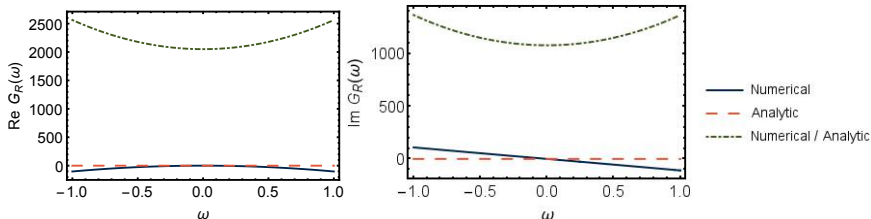


$\frac{\text{Numerical}}{\text{analytic}}$ is nice a straight line in each case. Therefore the Green's functions match well up to some overall normalization.

Dissipation

Dissipation at finite density

Choosing $a_-^{(0)} = 5$



$\frac{\text{Numerical}}{\text{analytic}}$ is *not* a straight line!

Dissipation

Dissipation at finite density

Finite density and small temperature

- ▶ Inner Region changes to **BH in $AdS_2 \times \mathbb{R}^{d-1}$**
- ▶ The story is same with two possible modifications
 - ▶ The $\mathcal{G}_R(\omega)$ may change and can be T -dependent.
 - ▶ a_{\pm}, b_{\pm} will be T -dependent.
- ▶ Actually the retarded Green's function at small T becomes

$$\begin{aligned} G_R^T(\omega) &= \frac{b_+(\omega, T) + \mathcal{G}_R(\omega, T) z_*}{a_+(\omega, T) + \mathcal{G}_R(\omega, T) z_*} \frac{b_-(\omega, T)}{a_-(\omega, T)} \\ &= \frac{b_+(\omega, T) - i\omega z_*}{a_+(\omega, T) - i\omega z_*} \frac{b_-(\omega, T)}{a_-(\omega, T)} \end{aligned} \quad (27)$$

- ▶ Leading order dissipation is same as zero temperature.

Conclusions

- ▶ The temperature independent dissipation is identical for all dimensions as long as the systems are in zero temperature (bulk is pure AdS).
- ▶ For higher dimensions $T \rightarrow 0$ and $T = 0$ (e.g, AdS_5 -BH and pure AdS_5 bulk, say) the coefficients don't match!
- ▶ Retarded Green's function at $T = 0$ is computed at finite density. Zero temperature dissipation shows up as leading term.
- ▶ The form of the Retarded Green's function at finite density and small (but finite) temperature is also obtained. The leading dissipative part remains the same.
- ▶ The leading order Green's function is “matched” (or rather *compared*) with numerical results up to some overall normalization.



Back up slides I

Perturbative solutions

Functional forms of $f_1(z)$, $f_2(z)$, $f_3(z)$

$$f_1(z) = \frac{1}{2} \tan^{-1}(\pi Tz) - \frac{1}{2} \text{Log}(1 + \pi Tz) + \frac{1}{4} \text{Log}(1 + \pi^2 T^2 z^2) \quad (28)$$

$$f_2(z) = \frac{1}{32} [4\{-4 + \tan^{-1}(\pi Tz) - \text{Log}(1 + \pi Tz)\}\{\tan^{-1}(\pi Tz) - \text{Log}(1 + \pi Tz)\} \\ - 4\{2 + \tan^{-1}(\pi Tz) - \text{Log}(1 + \pi Tz)\}\text{Log}(1 + \pi^2 T^2 z^2) + \text{Log}(1 + \pi^2 T^2 z^2)^2] \quad (29)$$

$$f_3(z) = \dots \quad (30)$$

Back up slides II

Black hole in $AdS_2 \times \mathbb{R}^{d-1}$

- ▶ The near horizon geometry for Near-extremal ($T \ll \mu$) RN Black hole

$$ds^2 = \frac{L_2^2}{\zeta^2} \left(-g(\zeta) dt^2 + \frac{d\zeta^2}{g(\zeta)} \right) + \mu_*^2 L_{d+1}^2 d\vec{x}^2 \quad (31)$$

$$A_t(\zeta) = \frac{1}{\sqrt{2d(d-1)}} \frac{1}{\zeta} \left(1 - \frac{\zeta}{\zeta_0} \right) \quad (32)$$

where $g(\zeta) := \left(1 - \frac{\zeta^2}{\zeta_0^2} \right)$, $\zeta_0 := \frac{z_*^2}{d(d-1)(z_* - z_0)}$

- ▶ The corresponding temperature, $T = \frac{1}{2\pi\zeta_0}$
- ▶ This nice structure breaks down for large temperature ($T \approx \mu$).