

Towards 2

QUANTUM THEORY OF SPACETIME AND FIELDS

S. Doplicher

Meeting on "Noncommutative
Geometry and QFT"

I M S, CHENNAI 10-15/01/03

JOINT WORK WITH

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REFERENCES:

K. FREDENHAGEN, J. ROBERTS, S.D.:

PLB '94, CMP '95

D. BAHNS, K. FREDENHAGEN, ^{{ hep-th/}
G. PIACITELLI, S.D., PLB '02 ⁰²⁰¹²²²

---, IN PREPARATION

SEE ALSO CONF. AND SCHOOL MEETINGS:

S.D. PROC. ICNP '94 (10 '95),

- AIHP '96

- hep-th/0105251 (KARFACE)

K. FREDENHAGEN PROC. HERSENBURG
MEETING 1999

- SUMMER SCHOOL IN CRETA, 2002

LARGE SCALE STRUCTURE OF SPACETIME:

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Pseudo Riemannian Manifold
locally modelled on the
MINKOWSKI SPACE

SMALL SCALE STRUCTURE?

LOCALIZING AN EVENT WITHIN

$$\Delta x \sim \Delta y \sim \Delta z \sim a$$

requires (HEISENBERG):

$$E \sim \frac{1}{a} \quad (\hbar = c = G = 1)$$

SCHWARZSCHILD RADIUS

$$R \sim E \sim \frac{1}{a}$$

$$a \gtrsim R \Rightarrow a \gtrsim 1$$

$$(a \gtrsim \lambda_p = 1.6 \cdot 10^{-33} \text{ cm})$$

BUT : $\Delta x \sim a$

$$\Delta y \sim \Delta z \sim L$$

$$E \sim \frac{1}{a} \text{ SPREAD OVER}$$

DISC OF RADIUS L GENERATES

$$\varphi_L(x) \xrightarrow{L \rightarrow \infty} 0 \text{ at any } x!$$

SPACE TIME UNCERTAINTY
RELATIONS:

ADMISSIBLE $\Delta x_0, \dots, \Delta x_3$:
IF THE STATE ω DESCRIBES
OUTCOME OF LOCALIZATION OF
AN EVENT WITH UNCERTAINTIES
 $\Delta x_0, \dots, \Delta x_3, E, E$.

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 8\pi \omega(T_{\mu\nu})$$

SHOULD GIVE RISE TO NO
TRAPPED SURFACE HIDING THE
EVENT

LEADS TO (DPR, '84, '95):

$$\left| \begin{aligned} \Delta q_0 (\Delta q_1 + \Delta q_2 + \Delta q_3) &\gtrsim \lambda_p^2 \\ \Delta q_2 \Delta q_3 + \Delta q_3 \Delta q_1 + \Delta q_1 \Delta q_2 &\gtrsim \lambda_p^2 \end{aligned} \right|$$

Weak form; stronger consequences:

$$\left(\begin{aligned} a &\equiv \inf \Delta q_\mu, \mu=0,1,2,3; \\ b &= \sup \Delta q_\nu, \nu=1,2,3 \end{aligned} \right)$$

$$a b \gtrsim \lambda_p^2$$

(Gys), or in presence of background $g_{\mu\nu}$

$$g_{00} a b \gtrsim \lambda_p^2.$$

Will concentrate on the weak form
 leading to a BASIC MODEL with
 MINIMAL QUANTUM DEVIATIONS FROM
 MINKOWSKI SPACE.

BETTER PICTURE
OF SPACETIME IN
THE SMALL:

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$$[q_\mu, q_\nu] = i Q_{\mu\nu}$$

$$\Rightarrow \Delta_\omega q_\mu \cdot \Delta_\omega q_\nu \gtrsim \frac{1}{2} |\omega(Q_{\mu\nu})|$$

$Q_{\mu\nu}$ must be s.t. this \Rightarrow STUR

BASIC MODEL

$$* [q_\mu, Q] = Q \cdot Q = 0$$

$$[q_0, q_1, q_2, q_3] \equiv \det \begin{pmatrix} q_0 & q_1 & q_2 & q_3 \\ q_0 & q_1 & q_2 & q_3 \\ q_0 & q_1 & q_2 & q_3 \\ q_0 & q_1 & q_2 & q_3 \end{pmatrix}$$

$$= \epsilon^{\mu\nu\rho\sigma} q_\mu q_\nu q_\rho q_\sigma = -\frac{1}{2} Q \cdot *Q$$

$$** \left(\frac{1}{2} Q \cdot *Q \right)^2 = I$$

$\vec{e}, \vec{m} \equiv$ el. resp. magn. parts of \mathcal{Q}

$$e^2 = m^2$$

$$(\vec{e} \cdot \vec{m})^2 = I$$

define manifold $\Sigma = \Sigma_+ \cup \Sigma_-$

$$\begin{aligned} \Sigma_{\pm} &\sim SL(2, \mathbb{C}) / \text{diag} \\ &\sim TS^2 \end{aligned}$$

DATA : $\vec{e} = \vec{m}, e^2 = m^2 = 1$

TANGENT VECTOR : SPEED, ORTH TO $\vec{e} = \vec{m}$, OF BOOST TAKING TO GENERIC PAIR

WEYL REL.

$$i\partial_4 \partial_3 \partial_2 \partial_1 \dots$$

Regular reps of $\mathcal{A} [\eta_\mu, q_\nu] = i Q_{\mu\nu}$

\equiv integrable to unitary reps of Weyl \mathcal{A} .

\sim non deg. reps of the enveloping

$$C^* \text{-alg } \Sigma \equiv$$

C^* completion of $\{ \mathcal{L}_0(\Sigma, L^1(\mathbb{R}^4)) \}$

with δ -trivial cov. dg ops at $\sigma \in \Sigma$

$$\equiv \mathcal{L}_0(\Sigma, \mathcal{K}) \sim \mathcal{L}_0(\Sigma) \otimes \mathcal{K}.$$

Generated by

$$f(Q) g(Q),$$

$$g \in \mathcal{L}_0(\Sigma);$$

$$f \in L^1(\mathbb{R}^4)^{\sim}, \quad f \equiv (\check{f})^{\sim}$$

$$f(Q) \equiv \int e^{i\alpha Q} \check{f}(\alpha) d^4\alpha$$

(... (Weyl - Wigner -oyal)

$$(a, \Lambda) \equiv L \in \mathcal{P}$$

$$q \rightarrow \Lambda q + a \cdot \mathbb{I}$$

risks autom τ_L of \mathcal{E}

$$L \rightarrow \tau_L \in \text{Aut } \mathcal{E}$$

action of full Poincaré' gp.

(Comes QP: Lukicak, Pnegg, Wey, ...)

Pure states of \mathcal{E} generalize points.

OPTIMALLY LOCALIZES STATES ω :

$$(\Delta_{\omega} q_0)^2 + \dots + (\Delta_{\omega} q_3)^2 = \min$$

$$\Leftrightarrow \omega(g(Q) e^{i\alpha q}) = e^{-\frac{1}{2} |\alpha|^2} \mu(y)$$

μ prob. measure on Σ_1 ;

$$\text{if } \sigma \in \Sigma_1, [q_{\mu}, q_{\nu}] = i \sigma_{\mu\nu} \cdot \mathbb{I}$$

\sim Schrödinger ops. q_1, q_2, p_1, p_2 .

QST seen by optimally loc. states

$\sim \sum_1 \times$ PHASE SPACE OF
2-d Q.M.

radius λ_p^2

effective cells of volume

λ_p^4 of FOLBY (HYS
J. MADORE,
A.R. BALACHANDRAN

CLASSICAL LIMIT:

$$\mathbb{R}^4 \times S^2 \times \{\pm\}$$

OPTIMALLY LOCALIZED STATES

$$\omega \equiv \mu \circ \eta$$

μ : state of $\mathcal{E}(\Sigma_1) \equiv \mathbb{Z}_1$

η : localisation map $\mathcal{E} \rightarrow \mathbb{Z}_1$

$$\eta(f(q)g(Q)) = \int \tilde{f}(\alpha) e^{-\frac{1}{2}|\alpha|^2} d^2\alpha \cdot g/5.$$

QFT on QST

On \mathbb{R}^4 :

$$f \in \mathcal{S} \rightarrow \phi(f) \approx \mathcal{O}$$

under op valued distribution

$$\phi \equiv \text{F.T. of } \check{\phi}$$

Def $\phi(q)$ as

$$\phi(q) = \int e^{ikq} \check{\phi}(k) d^4k$$

$\phi(q)$ affiliated to $\mathcal{L} \otimes \mathcal{O}$

$\omega \in \mathcal{I}(\mathcal{L})$,

$\langle \omega \times id, \phi(q) \rangle \equiv \phi(\omega) \in \mathcal{O}$
(affiliated to)

replaces "fields at a point."

FREE FIELDS OK

(GAUSSIAN DECAY OF $[\phi(\omega), \phi(\omega_a)]$
IN SPACELIKE DIR. IF ω OPTIMALLY W.C.)

INTERACTING QFT:

PROBLEMS WITH:

- NON UNIQUENESS OF EXT. OF THE CLASSICAL PRESCRIPTIONS
- BREAKDOWN OF RELATIVISTIC INV.
- " OF CAUSALITY

ROUTE I (DFR '84, '85)

$\int d^3q f(q)$ has well defined
 $q_0 = t$ meaning as a pers. in
 map

but requires $\int d\sigma$ ^{independent}
 or rel. to Σ_1 :
 breaks Lorentz

$$H_I(t) = \int_{\Sigma_1} d\sigma \int d^3q h_I(q) d^3q$$

$$c_g = \int_{\Sigma_1} d\sigma \int_{x_0=t} d^3x (\phi * \dots * \phi)(x)$$

FORMALLY SELF-ADJOINT : NO

PROBLEM WITH UNITARITY

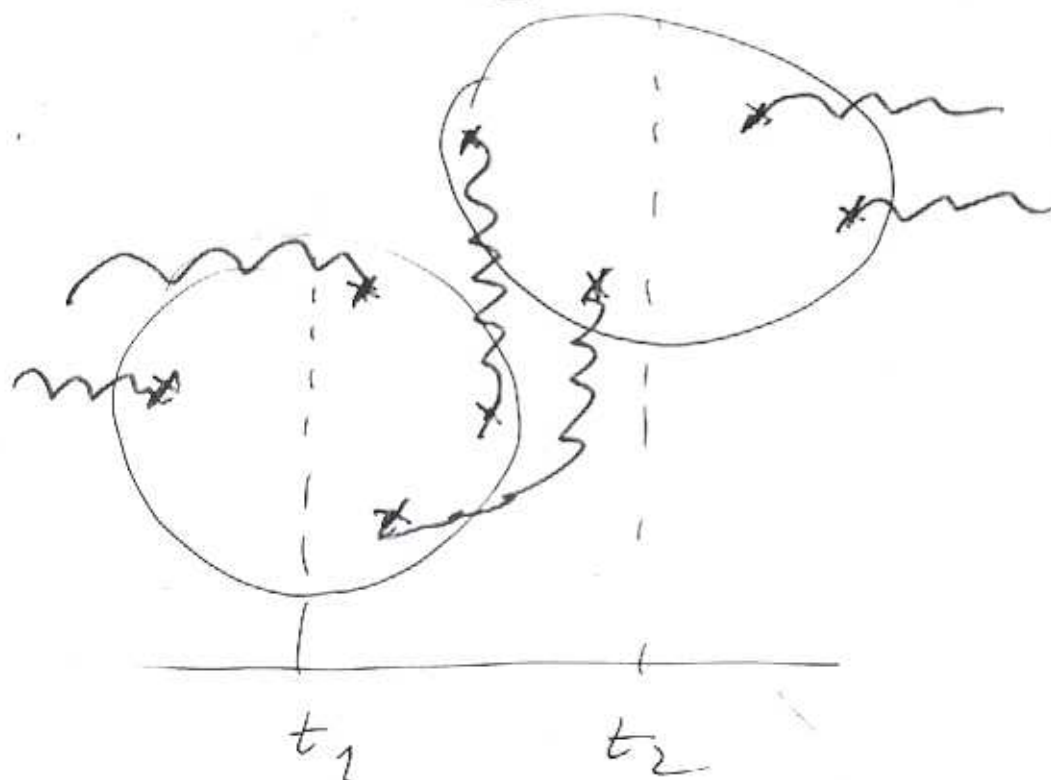
(cf BDFP for explicit consp.)

LIAO, SIBOLD

FEYNMAN RULES REPLACED BY:

$$S = T \exp i \int_{-\infty}^{\infty} H_I(t) dt$$

$$H_I(t) = \int G(t, x_1, \dots, x_n) : \phi(x_1) \dots \phi(x_n) : dx_1 \dots dx_n$$



MUST TIME ORDER t_1, t_2, \dots

NOT THE TIME ARGUMENT OF

THE FIELD OPERATORS! THAT WOULD BE ALLOWED IF H_I WERE LOCAL

E.G. IF WE COULD WRITE

$$(\phi * \phi)(x) = e^{i/2 \int Q_{\mu\nu} \frac{\partial}{\partial x_\mu} \frac{\partial}{\partial x_\nu} \phi(x) \phi(y)}$$

NOT WITH TIME

NOTE (DFR '84, '85):

- IF $\mathcal{H}(x)$ IS THE FREE HAMILTONIAN DENSITY AND H THE FREE HAMILTONIAN,

$$H = \int_{x_0=t} \mathcal{H}(x_0, \vec{x}) d^3x$$

BUT ALSO

$$H = \int_{q_0=t} \mathcal{H}(q) d^3q$$

UNAFPECTED BY Q.S.T.

- EXPLICIT FORM OF NONLOCAL KERNEL IN CASE OF \mathbb{R}^3 : GIVEN BY:

$$\int_{q_0=t} f_1(q) f_2(q) f_3(q) d^3q = \int d^4a d^4b \int_{x_0=t+\lambda_p(b-a)_0} f_1(x+\lambda_p a) f_2(x+\lambda_p b) f_3(x) S(a,b) d^4x$$

$$S(a,b) = -\frac{1}{2\pi^4} \left(\frac{\sin \gamma_+(a,b)}{\gamma_+(a,b)} + \frac{\sin \gamma_-(a,b)}{\gamma_-(a,b)} \right),$$

$$\gamma_{\pm}(a,b) = 2 \left\| a_0 \vec{b} - b_0 \vec{a} \pm \vec{a} \times \vec{b} \right\|.$$

ROUTE II (BDFP to appear) ¹⁷

$$\phi(x_1) \dots \phi(x_m) \rightsquigarrow \phi(q) \phi(q') \dots \phi(q^{(n)})$$

q, q', \dots index variables:

$$q_m \otimes I \otimes I \otimes \dots = q_m^1$$

$$I \otimes q_m \otimes I \otimes \dots = q_m^2$$

$$I \otimes I \otimes \dots \otimes q_m = q_m^n$$

$$q^{ij} \equiv q^i - q^j$$

$$[q_m^{ij}, q_m^{kl}] \neq 0$$

IF WE IDENTIFY THE CENTRES,
SAME C.R. AS q_m !

$$\begin{aligned} \mathcal{E} \text{ is a } \mathbb{Z}\text{-bimodule, } \mathbb{Z} &= \\ &= \mathbb{Z}(M(\mathcal{E})) = \mathcal{C}_B(\mathcal{I}) : \end{aligned}$$

$$\mathcal{E} \otimes_{\mathbb{Z}} \dots \otimes_{\mathbb{Z}} \mathcal{E}$$

$$\bullet [q_m^{ij}, q_\nu^{ij}] = 2i Q_{\mu\nu}$$

$$\bar{q}_\mu \equiv \frac{1}{n} \sum_j q_\mu^j \quad \text{barycenter}$$

$$\bullet [\bar{q}_\mu, q_\nu^{ij}] = 0 \quad \text{strong sense}$$

von Neumann uniqueness and each $\sigma \in \Sigma$

$$\Rightarrow \bar{q}_\mu q_\nu^{ij} \stackrel{\sim}{=} \bar{q}_\mu \otimes_{\mathbb{Z}} q_\nu^{ij}$$

defines a natural map:

$$\bar{q} \longrightarrow \frac{1}{\sqrt{n}} q \otimes I \otimes I \dots$$

$$q^i - q^j \longrightarrow I \otimes (q^i - q^j)$$

$$\mathcal{E} \otimes_{\mathbb{Z}}^n \longrightarrow \mathcal{E} \otimes_{\mathbb{Z}}^{(n+1)} \quad (\text{into})$$

COINCIDENT PTS: $q^{ij} = 0$ VIOLATES C.R.

$$\left[\frac{1}{\sqrt{2}} q_\mu^{ij}, \frac{1}{\sqrt{2}} q_\nu^{ij} \right] = i Q_{\mu\nu}$$

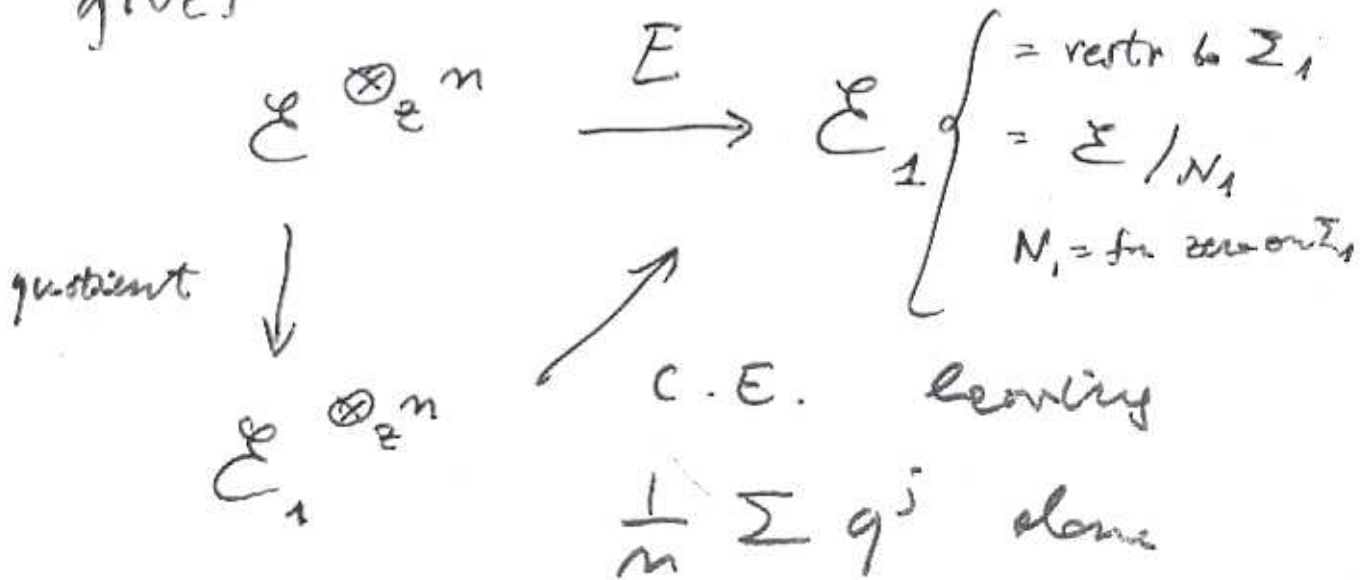
COMPARE THE MAP ABOVE

$$\mathcal{E}^{\otimes_{\mathbb{Z}} n} \rightarrow \mathcal{E}^{\otimes_{\mathbb{Z}} (n+1)}$$

WITH $\text{id} \otimes \eta \otimes \dots \otimes \eta$

$\eta : \mathcal{E} \rightarrow \mathcal{Z}_1$ optimal l.c. map

gives



evaluating on each $f(\frac{1}{\sqrt{2}}(q^i - q^j))$

the optimally localizing maps η

$$\begin{aligned} & \left(\begin{matrix} \vdots \\ \phi(q_1) \\ \vdots \\ \phi(q_n) \\ \vdots \end{matrix} \right)_{\mathcal{Q}} \equiv \\ & \equiv E \left(\left(\begin{matrix} \vdots \\ \phi(q_1) \\ \vdots \\ \phi(q_n) \\ \vdots \end{matrix} \right) \right) \end{aligned}$$

$$\eta: \mathcal{E} \rightarrow \mathbb{Z}_2$$

is characterized by

$$\eta(q_m) = 0$$

$$\eta\left(\sum_m q_m^2\right) = 2 \cdot I$$

IMMEDIATELY CHECKED FOR

$$\eta \otimes \eta, \quad \frac{1}{\sqrt{2}}(q \otimes I - I \otimes q)$$

INCIDENTALLY NOTE THAT THE (EUCL., 4-D)

DISTANCE IN QST IS THE
OPERATOR d : $d^2 =$

$$= |q - q'|^2 \equiv \sum_{m=0}^3 (q_m \otimes I - I \otimes q_m)^2$$

SINCE $\frac{1}{\sqrt{2}}(q - q') \approx q,$

$$|q - q'|^2 \geq 4, \quad d \geq 2\lambda_p,$$

WHAT ABOUT

$$\inf |\vec{q} - \vec{q}'| \quad ?$$

$$\frac{1}{\sqrt{2}} (\vec{q} - \vec{q}') \approx \vec{q}$$

$$(\inf |\vec{q}|)^2 = \inf_{\omega} \omega (q_1^2 + q_2^2 + q_3^2)$$

$$\geq \omega (q_2^2) + 2 \Delta q_1 - \Delta q_3$$

$$\geq \omega (q_2^2) + |m_2|$$

since choice of axis 2 was arbitrary,

$$(\inf |\vec{q}|)^2 \geq |\vec{m}| = \|\delta\|$$

and inf has to occur for $\delta \in \Sigma_{\perp}$.

Choose $\vec{n}_2 = \vec{m}$; $[q_2, q_1] = [q_2, q_3] = 0$,

$[q_3, q_1] = -iI$; can make choice of ω

s.t. $\Delta q_1 = \Delta q_3 = \frac{1}{\sqrt{2}}$, $\Delta q_2 < \epsilon$:

\Rightarrow

$$\boxed{\inf |\vec{q} - \vec{q}'| = \sqrt{2} \cdot \lambda_F}$$

$$H_I(t) \equiv \int_{q_0=t}^1 d^3q \lambda : \phi(q) :^n$$

ϵ -priori on op. valued fn on Σ_1 ,
is constant on Σ_1 ; no ad-hoc
integration on 0 is needed,

$$\lambda \rightarrow \lambda(t) \text{ (e.g. } \lambda e^{-\epsilon t^2} \text{)}$$

$$S_\lambda \equiv T \exp i \int_{-\infty}^{\infty} \lambda(t) H_I(t) dt$$

Then the GELL'MANN-LOW EXPANSION

$$S_\lambda \sim \langle S_\lambda \rangle_0$$

if finite at each order of p.e.,

i.e. NO UV DIVERGENCE.

BUT: $i\epsilon$ does not leave the
FREE HAMILTONIAN UNCHANGED.

$$\int_{q_0=t} d^3 q \lambda : \phi^n(q) :_{\mathcal{Q}} =$$

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$$= \int_{x_0=t} d^3 x H_{\text{eff}}^{\text{I}}(x);$$

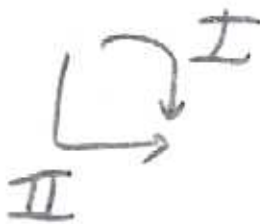
$$H_{\text{eff}}^{\text{I}}(x) =$$

$$= \int_{\mathbb{R}^{4n}} d^4 x_1 \dots d^4 x_n K_n(x-x_1, \dots, x-x_n) \lambda : \phi(x_1) \dots \phi(x_n) : ,$$

$$K_n(y_1, \dots, y_n) =$$

$$= C_n e^{-\frac{1}{2} \sum_{\mu, j} (y_j^\mu)^2} \delta^{(4)}\left(\sum_{j=1}^n y_j\right).$$

(BDFF, to appear soon)

However: both 

- VIOLATE LORENTZ

- NON LOCAL

III YANG FELDMAN EQ:

$$\phi(q) = \phi^{\text{in}}(q) + \lambda \Delta^{\text{ret}} \times \phi^3(q);$$

PARTIAL WICK PRODUCTS, LORENTZ COV. PRES.

BUT BREAKS DOWN ELSEWHERE: ASYMPT. STATES?
(BDFP WORK IN PROGRESS)

IV GAUGE THEORY ON

QFT (WESS, GOLDBERGER, ZUHNINO
(DF in preparation) ^{UNIVERSAL} _{coll. (i)})

"Classical" diff. calculus on \mathcal{E} :

$$d^d A = \frac{\partial}{\partial a_\mu} \tau_a(A) \Big|_{a=0} da_\mu$$

UNIVERSAL DIFF. CALC. :

$$dA = I \otimes A - A \otimes I$$

(cf A. CONNES "Noncommutative Geometry")

METRIC ?

CODIFFERENTIAL δ ?

ANSWER (FOR FLAT CASE

$$g = \text{diag}(+, -, -, -),$$

WHICH YIELDS BACK TO

CLASSICAL YANG-MILLS EQ

(NON COMMUTATIVE GAUGE :

$$U(1) \rightarrow \mathcal{U}(\mathcal{C}_0(\mathbb{R}) + \langle I \rangle)$$

ON $\mathcal{E} \Sigma$:

$$U(1) \rightarrow \mathcal{U}(\mathcal{E} + \langle I \rangle) \equiv G =$$

$$= \Pi \times \mathcal{G}$$

$\mathcal{G}(\sigma) = \{ \text{unitaries of the form } I + T, T \text{ compact} \},$
 $\sigma \in \Sigma ;$

ESSENTIAL INGREDIENT:

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$$\text{Ad} (Q^{-1} q)_\mu = \frac{\partial}{\partial x_\mu}$$

(Q^{-1} def by QUANTUM CONDITIONS).

GAUGE TR:

$$q \rightarrow U q U^{-1} \neq q!$$

J. WESS et al:

$$q_{\text{cov}} = q + \langle A, dq \rangle$$

allows to construct gauge inv
quantities.

PROB. OBSERVABLE TESTS?

(contact with work by

AMELINO-CAMELIA et al ?)

WESS et al

CHAICHIAN et al; MORITA

How COULD THESE RELATIONS
REDUCE POSSIBLY TO THE ALCWARD
REL.

$$[q_\mu, q_\nu] = i g_{00}^{-1} Q_{\mu\nu}$$

IN A QUASI STATIONARY, SEMICLASSICAL
APPROXIMATION?

AN ALCWARD ARGUMENT IN FAVOUR:

REPLACE THE NOMIN. g_{00}^{-1} BY THE
INVARIANT SCALAR CURVATURE R :

$$[q_\mu, q_\nu] = i \alpha R Q_{\mu\nu}$$

$Q_{\mu\nu}$ as in the basic model. $EE \Rightarrow$

$$[q_\mu, q_\nu] = -8\pi i \alpha g^{\lambda\rho} T_{\lambda\rho} Q_{\mu\nu}$$

APPROXIMATION: REPLACE R.H.S.

BY EXPECTATION IN A SHARPLY LOCALIZED
STATE ψ , AND $g^{\lambda\rho}$ BY CLASSICAL

$$[q_\mu, q_\nu] = -8\pi i \alpha \mathcal{Q}_{\mu\nu} g^{\lambda\rho} \omega(T_{\lambda\rho})$$

②
I SPREADING TO THE SHARP LOCALIZATION REGION

$$\omega(T_{\lambda\rho}) = \omega_0(T_{\lambda\rho})$$

②
I ASSUME

$$(\Omega, T_{\lambda\rho}, \Omega) = \Lambda \delta_{\lambda\rho} \delta_{\mu\nu}$$

STATIONARY APPROX FOR $g^{\lambda\rho}$:

$$g^{ij} = g^{j0} = 0 \quad j=1,2,3$$

THEN $g^{00} = (g_{00})^{-1}$

AND OUR APPROX. QUANTUM CONDITIONS BEFORE:

$$[q_\mu, q_\nu] = -8\pi i \alpha \Lambda g_{00}^{-1} \mathcal{Q}_{\mu\nu}$$

HOWEVER THE DEPENDENCE UPON
 $T_{\mu\nu}$ OF $[q_\mu, q_\nu]$ MIGHT HAVE
PHYSICALLY MEANINGFUL CONSEQUENCES:

- INFLATION ?
- COSMOLOGICAL CONSTANT $\neq 0$?
- DARK MATTER ?