

Introduction to MSSM

Shrihari Gopalakrishna



Institute of Mathematical Sciences (IMSc), Chennai

SUSY & DM 2013

IISc, Bangalore

October 2013

Talk Outline

- Supersymmetry (SUSY) Basics
 - Superfield formalism
 - Constructing SUSY invariant theory
 - SUSY breaking
- Minimal Supersymmetric Standard Model (MSSM)
 - SUSY preserving Lagrangian and soft-breaking terms
 - R-parity
 - Superpartner Mixing
- Implications
 - Dark Matter
 - 125 GeV Higgs



Supersymmetry (SUSY)

Symmetry: Fermions \leftrightarrow Bosons

$$Q |\Phi\rangle = |\Psi\rangle \quad ; \quad Q |\Psi\rangle = |\Phi\rangle$$

Q_α is a spinorial charge

Reviews: [Wess & Bagger]

SUSY algebra:

$$\{Q_\alpha, \bar{Q}_{\dot{\beta}}\} = 2\sigma^\mu_{\alpha\dot{\beta}} P_\mu$$

$$\{Q_\alpha, Q_\beta\} = \{\bar{Q}_{\dot{\alpha}}, \bar{Q}_{\dot{\beta}}\} = 0$$

$$[P^\mu, Q_\alpha] = [P^\mu, \bar{Q}_{\dot{\alpha}}] = 0$$



Superfield

SUSY algebra realized through a **Superfield** $\Phi(x_\mu, \theta, \bar{\theta})$

θ are fermionic (Grassmann) coordinates that anticommute: θ_α , $\alpha = \{1, 2\}$

$$\{\theta, \theta\} = \{\theta, \bar{\theta}\} = \{\bar{\theta}, \bar{\theta}\} = 0 \quad \Rightarrow \quad (\theta_1)^2 = (\theta_2)^2 = 0$$

$$\text{Define } \theta\theta \equiv \theta^\alpha\theta_\alpha; \quad \bar{\theta}\bar{\theta} \equiv \bar{\theta}_{\dot{\alpha}}\bar{\theta}^{\dot{\alpha}}; \quad \theta^\alpha \equiv \epsilon^{\alpha\beta}\theta_\beta; \quad \bar{\theta}_{\dot{\alpha}} \equiv \epsilon_{\dot{\alpha}\dot{\beta}}\bar{\theta}^{\dot{\beta}}; \quad \epsilon^{12} = \epsilon_{21} = +1$$

● Chiral Superfield

- $\bar{D}\Phi_L = 0$; $D\Phi_R = 0$
- $\Phi_L = \phi(y) + \sqrt{2}\theta\psi_L(y) + \theta\theta F(y)$
 F : auxiliary field

$$\delta_\xi \phi = \sqrt{2}\xi\psi$$

$$\delta_\xi \psi = i\sqrt{2}\sigma^m \bar{\xi} \partial_m \phi + \sqrt{2}\xi F$$

$$\delta_\xi F = i\sqrt{2}\bar{\xi} \bar{\sigma}^m \partial_m \psi$$

● Vector Superfield

- $V = V^\dagger$
- $V(x, \theta, \bar{\theta}) = -\theta\sigma_\mu\bar{\theta}A^\mu(x) + i\theta\theta\bar{\theta}\bar{\lambda}(x) - \bar{\theta}\bar{\theta}\theta\lambda(x) + \frac{1}{2}\theta\theta\bar{\theta}\bar{\theta}D(x)$
 D : auxiliary field

$$\delta_\xi A_{mn} = i[(\xi\sigma^n\partial_m\bar{\lambda} + \bar{\xi}\bar{\sigma}^n\partial_m\lambda) - (n \leftrightarrow m)]$$

$$\delta_\xi \lambda = i\xi D + \sigma^{mn}\xi A_{mn}$$

$$\delta_\xi D = \bar{\xi}\bar{\sigma}^m\partial_m\lambda - \xi\sigma^m\partial_m\bar{\lambda}$$



Constructing a SUSY theory

Under a SUSY transformation, F and D transform into total derivatives

- So they can be used for constructing SUSY invariant \mathcal{L}



Constructing a SUSY theory

Under a SUSY transformation, F and D transform into total derivatives

- So they can be used for constructing SUSY invariant \mathcal{L}

A SUSY gauge invariant theory

$$\mathcal{L} = \Phi_i^\dagger e^{2gV} \Phi_i \Big|_{\theta\theta\bar{\theta}\bar{\theta}} + (\mathcal{W}(\Phi_i) \Big|_{\theta\theta} + h.c.) + \frac{1}{32g^2} W_\alpha W^\alpha \Big|_{\theta\theta}$$

Superpotential $\mathcal{W}(\Phi)$, a Holomorphic function

Gauge Kinetic Function $W_\alpha = -\frac{1}{4} \bar{D}\bar{D}e^{-V} D_\alpha e^V$



Constructing a SUSY theory

Under a SUSY transformation, F and D transform into total derivatives

- So they can be used for constructing SUSY invariant \mathcal{L}

A SUSY gauge invariant theory

$$\mathcal{L} = \Phi_i^\dagger e^{2gV} \Phi_i \Big|_{\theta\theta\bar{\theta}\bar{\theta}} + (\mathcal{W}(\Phi_i)|_{\theta\theta} + h.c.) + \frac{1}{32g^2} W_\alpha W^\alpha \Big|_{\theta\theta}$$

Superpotential $\mathcal{W}(\Phi)$, a Holomorphic function

Gauge Kinetic Function $W_\alpha = -\frac{1}{4} \bar{D} \bar{D} e^{-V} D_\alpha e^V$

Eliminating the Auxiliary fields

$$\begin{aligned} \mathcal{L} = & |D_\mu \phi_i|^2 - i \bar{\psi}_i \sigma_\mu D^\mu \psi_i - g \sqrt{2} \left(\phi_i^* T^a \psi_i \lambda^a + \lambda^{a\dagger} \psi_i^\dagger T^a \phi_i \right) \\ & - \left(\frac{1}{2} \frac{\partial^2 \mathcal{W}(\phi_i)}{\partial \phi_j \partial \phi_k} \psi_j \psi_k + h.c. \right) - \left| \frac{\partial \mathcal{W}(\phi_i)}{\partial \phi_j} \right|^2 \\ & - \frac{1}{4} F_{\mu\nu}^a F_a^{\mu\nu} - \frac{1}{2} \sum_a \left| g \phi_i^* T_{ij}^a \phi_j \right|^2 - i \lambda^{a\dagger} \bar{\sigma}_\mu D^\mu \lambda_a \end{aligned}$$



Consequences

Solution to gauge hierarchy problem

$$\begin{array}{c}
 t_L, t_R \\
 \circlearrowleft \\
 h \text{ --- } \text{---} h \\
 -i \frac{y_t}{\sqrt{2}}
 \end{array}
 +
 \begin{array}{c}
 \tilde{t}_L, \tilde{t}_R \\
 \circlearrowleft \\
 h \text{ --- } \text{---} h \\
 -i \frac{y_{\tilde{t}}}{2}
 \end{array}
 = 0$$

Λ^2 divergence cancelled

[Romesh Kaul's talk]

(Similarly W^\pm , Z divergences cancelled by $\tilde{\lambda}$)

Consequences

Solution to gauge hierarchy problem

$$\begin{array}{c}
 \text{---} h \text{---} \bigcirc \text{---} h \text{---} \\
 \text{---} -i \frac{y_t}{\sqrt{2}} \text{---} \\
 t_L, t_R \\
 \end{array}
 +
 \begin{array}{c}
 \text{---} h \text{---} \bigcirc \text{---} h \text{---} \\
 \text{---} -i \frac{y_{\tilde{t}}}{2} \text{---} \\
 \tilde{t}_L, \tilde{t}_R \\
 \end{array}
 = 0$$

Λ^2 divergence cancelled

[Romesh Kaul's talk]

(Similarly W^\pm , Z divergences cancelled by $\tilde{\lambda}$)

- Lightest SUSY Particle (LSP) stable dark matter (if R_p conserved)
- Gauge Coupling Unification - SUSY $SO(10)$ GUT
Includes $\nu_R \Rightarrow$ Neutrino mass via seesaw



SUSY breaking

- Exact SUSY $\implies M_{\psi} = M_{\phi}$; $M_A = M_{\tilde{\lambda}}$
 - So experiment \implies **SUSY must be broken**
- SUSY broken if and only if $\langle 0|H|0\rangle > 0$
 - Spontaneous SUSY breaking
 - O'Raifeartaigh F -term breaking
 - Fayet-Iliopoulos D -term breaking
 - $STr(M^2) = 0 \implies$ cannot break SUSY spontaneously using SM superfield
 - Hidden sector breaking $\xleftrightarrow{\text{Mediation}}$ Communicated to SM
Spectrum depends on Mediation type + RGE
- In effective low-energy theory
 - Explicit soft-breaking terms, i.e., with dimensionful parameters



MSSM

The Minimal Supersymmetric Standard Model (MSSM)

Reviews: [Martin] [Drees] [Drees,Godbole,Roy] [Baer,Tata]



MSSM fields

To every SM particle, add a **superpartner** (with spin differing by 1/2)

Matter fields (Chiral Superfields)

	$(SU(3), SU(2))_{U(1)}$	Components
Q	$(3, 2)_{1/6}$	$(\tilde{q}_L, q_L, F_Q) : \bar{q}_L = \begin{pmatrix} \tilde{u}_L \\ \tilde{d}_L \end{pmatrix}; q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}$
U^c	$(\bar{3}, 1)_{-2/3}$	$(\tilde{u}_R^*, u_R^c, F_U)$
D^c	$(\bar{3}, 1)_{1/3}$	$(\tilde{d}_R^*, d_R^c, F_D)$
L	$(1, 2)_{-1/2}$	$(\tilde{\ell}_L, \ell_L, F_L) : \tilde{\ell}_L = \begin{pmatrix} \tilde{\nu}_L \\ \tilde{e}_L \end{pmatrix}; \ell_L = \begin{pmatrix} \nu_L \\ e_L \end{pmatrix}$
E^c	$(1, 1)_1$	$(\tilde{e}_R^*, e_R^c, F_E)$
(N^c)	$(1, 1)_0$	$(\tilde{\nu}_R^*, \nu_R^c, F_N)$

Gauge fields
(Vector Superfields)

	Components
$SU(3)$	(g_μ, \tilde{g}, D_3)
$SU(2)$	(W_μ, \tilde{W}, D_2)
$U(1)$	(B_μ, \tilde{B}, D_1)

Higgs fields (Chiral Superfields)

	$(SU(3), SU(2))_{U(1)}$	Components
H_u	$(1, 2)_{1/2}$	$(h_u, \tilde{h}_u, F_{H_u}) : h_u = \begin{pmatrix} h_u^+ \\ h_u^0 \end{pmatrix}; \tilde{h}_u = \begin{pmatrix} \tilde{h}_u^+ \\ \tilde{h}_u^0 \end{pmatrix}$
H_d	$(1, 2)_{-1/2}$	$(h_d, \tilde{h}_d, F_{H_d}) : h_d = \begin{pmatrix} h_d^0 \\ h_d^- \end{pmatrix}; \tilde{h}_d = \begin{pmatrix} \tilde{h}_d^0 \\ \tilde{h}_d^- \end{pmatrix}$



MSSM Superpotential

Write most general \mathcal{W} consistent with $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$

- $\mathcal{W} = U^c y_u Q H_u - D^c y_d Q H_d - E^c y_e L H_d + \mu H_u H_d + (N^c y_n L H_u)$
- $\mathcal{W}_{\Delta L} = L H_u + L E^c L + Q D^c L ; \quad \mathcal{W}_{\Delta B} = U^c D^c D^c$
 - $\mathcal{W}_{\Delta L} + \mathcal{W}_{\Delta B}$ induce proton decay : $\tau_p \sim 10^{-10} s$ for $\tilde{m} \sim 1 \text{ TeV}$



MSSM Superpotential

Write most general \mathcal{W} consistent with $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$

- $\mathcal{W} = U^c y_u Q H_u - D^c y_d Q H_d - E^c y_e L H_d + \mu H_u H_d + (N^c y_n L H_u)$
- $\mathcal{W}_{\Delta L} = L H_u + L E^c L + Q D^c L$; $\mathcal{W}_{\Delta B} = U^c D^c D^c$
 - $\mathcal{W}_{\Delta L} + \mathcal{W}_{\Delta B}$ induce proton decay : $\tau_p \sim 10^{-10} s$ for $\tilde{m} \sim 1 \text{ TeV}$

So impose Matter Parity $R_M = (-1)^{3(B-L)}$ to forbid ΔL and ΔB terms

For components \implies **R-parity** $R_p = (-1)^{3(B-L)+2s}$

$R_p(\text{particle}) = +1$, $R_p(\text{sparticle}) = -1$

Consequence : The **Lightest SUSY Particle (LSP) is stable**

- Cosmologically stable Dark Matter
- Missing Energy at Colliders



Soft SUSY breaking

Effective parametrization with explicit soft-SUSY-breaking terms

$$\begin{aligned}
 \mathcal{L}_{SUSY Br}^{soft} \supset & -\tilde{Q}^\dagger \tilde{m}_Q^2 \tilde{Q} - \tilde{u}_R^\dagger \tilde{m}_u^2 \tilde{u}_R - \tilde{d}_R^\dagger \tilde{m}_d^2 \tilde{d}_R - \tilde{L}^\dagger \tilde{m}_L^2 \tilde{L} - \tilde{e}_R^\dagger \tilde{m}_e^2 \tilde{e}_R - (\tilde{\nu}_R^\dagger \tilde{m}_\nu^2 \tilde{\nu}_R) \\
 & - \frac{1}{2} M_1 \tilde{B} \tilde{B} - \frac{1}{2} M_2 \tilde{W} \tilde{W} - \frac{1}{2} M_3 \tilde{g} \tilde{g} + h.c. \\
 & - \tilde{u}^c A_u \tilde{Q} H_u + \tilde{d}^c A_d \tilde{Q} H_d + \tilde{e}^c A_e \tilde{L} H_d - (\tilde{\nu}^c A_\nu \tilde{L} H_u) + h.c. \\
 & - m_{H_u}^2 H_u^\dagger H_u - m_{H_d}^2 H_d^\dagger H_d - (B\mu H_u H_d + h.c.)
 \end{aligned}$$



Soft SUSY breaking

Effective parametrization with explicit soft-SUSY-breaking terms

$$\begin{aligned}
 \mathcal{L}_{SUSY Br}^{soft} \supset & -\tilde{Q}^\dagger \tilde{m}_Q^2 \tilde{Q} - \tilde{u}_R^\dagger \tilde{m}_u^2 \tilde{u}_R - \tilde{d}_R^\dagger \tilde{m}_d^2 \tilde{d}_R - \tilde{L}^\dagger \tilde{m}_L^2 \tilde{L} - \tilde{e}_R^\dagger \tilde{m}_e^2 \tilde{e}_R - (\tilde{\nu}_R^\dagger \tilde{m}_\nu^2 \tilde{\nu}_R) \\
 & - \frac{1}{2} M_1 \tilde{B} \tilde{B} - \frac{1}{2} M_2 \tilde{W} \tilde{W} - \frac{1}{2} M_3 \tilde{g} \tilde{g} + h.c. \\
 & - \tilde{u}^c A_u \tilde{Q} H_u + \tilde{d}^c A_d \tilde{Q} H_d + \tilde{e}^c A_e \tilde{L} H_d - (\tilde{\nu}^c A_\nu \tilde{L} H_u) + h.c. \\
 & - m_{H_u}^2 H_u^\dagger H_u - m_{H_d}^2 H_d^\dagger H_d - (B\mu H_u H_d + h.c.)
 \end{aligned}$$

UV SUSY breaking and mediation dynamics will set these parameters

- Eg: Gravity Mediation (MSUGRA, CMSSM)
 - Inputs \tilde{m}_0 , $M_{1/2}$, A_0 , $\tan\beta$, $\text{sign}(\mu)$ at GUT scale
 - TeV scale values determined by RGE



Electroweak symmetry breaking (EWSB)

$$\langle H_u \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_u \end{pmatrix}; \quad \langle H_d \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_d \\ 0 \end{pmatrix}; \quad v^2 = v_u^2 + v_d^2; \quad \tan \beta \equiv \frac{v_u}{v_d};$$

Physical Higgses: h^0 , H^0 , A^0 , H^\pm



Electroweak symmetry breaking (EWSB)

$$\langle H_u \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_u \end{pmatrix}; \quad \langle H_d \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_d \\ 0 \end{pmatrix}; \quad v^2 = v_u^2 + v_d^2; \quad \tan \beta \equiv \frac{v_u}{v_d};$$

Physical Higgses: h^0 , H^0 , A^0 , H^\pm

$$\mathcal{V} = (|\mu|^2 + m_{H_u}^2)|h_u|^2 + (|\mu|^2 + m_{H_d}^2)|h_d|^2 - (b_\mu h_u h_d + h.c.) + \frac{1}{8}(g^2 + g'^2)(|h_u|^2 - |h_d|^2)^2$$

Minimization and EWSB

$$\sin(2\beta) = \frac{2b_\mu}{m_{H_u}^2 + m_{H_d}^2 + 2|\mu|^2} \quad \text{and} \quad m_Z^2 = \frac{|m_{H_d}^2 - m_{H_u}^2|}{\sqrt{1 - \sin^2(2\beta)}} - m_{H_u}^2 - m_{H_d}^2 - 2|\mu|^2$$



Electroweak symmetry breaking (EWSB)

$$\langle H_u \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_u \end{pmatrix}; \quad \langle H_d \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_d \\ 0 \end{pmatrix}; \quad v^2 = v_u^2 + v_d^2; \quad \tan \beta \equiv \frac{v_u}{v_d};$$

Physical Higgses: h^0 , H^0 , A^0 , H^\pm

$$\mathcal{V} = (|\mu|^2 + m_{H_u}^2)|h_u|^2 + (|\mu|^2 + m_{H_d}^2)|h_d|^2 - (b_\mu h_u h_d + h.c.) + \frac{1}{8}(g^2 + g'^2)(|h_u|^2 - |h_d|^2)^2$$

Minimization and EWSB

$$\sin(2\beta) = \frac{2b_\mu}{m_{H_u}^2 + m_{H_d}^2 + 2|\mu|^2} \quad \text{and} \quad m_Z^2 = \frac{|m_{H_d}^2 - m_{H_u}^2|}{\sqrt{1 - \sin^2(2\beta)}} - m_{H_u}^2 - m_{H_d}^2 - 2|\mu|^2$$

- So we need μ^2 (SUSY preserving param) $\sim m^2$ (SUSY br param)! Why?
 - This is called the **μ -problem**



125 GeV Higgs

At 1-loop

$$m_h^2 \approx m_Z^2 \cos^2(2\beta) + \frac{3g_2^2 m_t^4}{8\pi^2 m_W^2} \left[\ln \left(\frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right) + \frac{X_t^2}{m_{\tilde{t}_1} m_{\tilde{t}_2}} \left(1 - \frac{X_t^2}{12m_{\tilde{t}_1} m_{\tilde{t}_2}} \right) \right]$$

[Haber, Hempfling, Hoang ,1997] [Debtosh's thesis]

where $X_t = A_t - \mu \cot \beta$

- $m_h = 125$ GeV needs sizable loop contribution

- Hard! Needs large $m_{\tilde{t}_1} m_{\tilde{t}_2}$ or large X_t^2

- But $\delta m_{H_u}^2 \approx \frac{3g_2^2 m_t^4}{8\pi^2 m_W^2} \left[\ln \left(\frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right) + \frac{X_t^2}{2m_{\tilde{t}_1} m_{\tilde{t}_2}} \left(1 - \frac{X_t^2}{6m_{\tilde{t}_1} m_{\tilde{t}_2}} \right) \right]$

So fine-tuning necessary to keep m_Z^2 correct (cf previous EWSB relation)
 "Little hierarchy problem"



Neutralino mixing

Neutralino: Neutral EW gauginos ($\tilde{B}, \tilde{W}^3, \tilde{H}_u^0, \tilde{H}_d^0$): Majorana states

$\mathcal{L} \supset$

$$-\frac{1}{2} \begin{pmatrix} \tilde{B} & \tilde{W}^3 & \tilde{H}_u^0 & \tilde{H}_d^0 \end{pmatrix} \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \begin{pmatrix} \tilde{B} \\ \tilde{W}^3 \\ \tilde{H}_u^0 \\ \tilde{H}_d^0 \end{pmatrix}$$

Diagonalizing this $\begin{pmatrix} \tilde{B} \\ \tilde{W}^3 \\ \tilde{H}_u^0 \\ \tilde{H}_d^0 \end{pmatrix} \rightarrow \begin{pmatrix} \tilde{\chi}_1^0 \\ \tilde{\chi}_2^0 \\ \tilde{\chi}_3^0 \\ \tilde{\chi}_4^0 \end{pmatrix}$: the mass eigenstates



Chargino mixing

Chargino: Charged EW gauginos ($\tilde{W}^\pm, \tilde{H}^\pm$), $\tilde{w}^\pm = \tilde{w}_1 \pm i\tilde{w}_2$

Form Dirac states $\tilde{W}^+ = \begin{pmatrix} \tilde{W}_\alpha^+ \\ \tilde{W}^{-\dot{\alpha}} \end{pmatrix}$; $\tilde{H}^+ = \begin{pmatrix} \tilde{H}_u^+ \\ \tilde{H}_d^+ \end{pmatrix}$

$\mathcal{L} \supset - \begin{pmatrix} \overline{\tilde{W}^+} & \overline{\tilde{H}^+} \end{pmatrix} (M_\chi P_L + M_\chi^\dagger P_R) \begin{pmatrix} \tilde{W}^+ \\ \tilde{H}^+ \end{pmatrix} + h.c.$

$$\text{where } M_\chi = \begin{pmatrix} M_2 & \sqrt{2} \sin \beta m_W \\ \sqrt{2} \cos \beta m_W & \mu \end{pmatrix}$$

Diagonalizing this $\begin{pmatrix} \tilde{W}^+ \\ \tilde{H}^+ \end{pmatrix} \rightarrow \begin{pmatrix} \tilde{\chi}_1^+ \\ \tilde{\chi}_2^+ \end{pmatrix}$: the mass eigenstates



Scalar mixing

Eg. stop sector $(\tilde{t}_L, \tilde{t}_R)$

$$\mathcal{L} \supset - \begin{pmatrix} \tilde{t}_L^* & \tilde{t}_R^* \end{pmatrix} \begin{pmatrix} \tilde{m}_{LL}^2 + m_t^2 + \Delta_L & (v_u A_t - \mu^* \cot \beta m_t)^* \\ (v_u A_t - \mu^* \cot \beta m_t) & \tilde{m}_{RR}^2 + m_t^2 + \Delta_R \end{pmatrix} \begin{pmatrix} \tilde{t}_L \\ \tilde{t}_R \end{pmatrix}$$

$$\text{where } \Delta = (T_3 - Qs_W^2) \cos(2\beta) m_Z^2$$

Diagonalizing this $\begin{pmatrix} \tilde{t}_L \\ \tilde{t}_R \end{pmatrix} \rightarrow \begin{pmatrix} \tilde{t}_1 \\ \tilde{t}_2 \end{pmatrix}$: the mass eigenstates



Flavor Issues

Flavor Problem

- In general \tilde{m}_{ij} can have arbitrary flavor structure and phases (MSSM has 9 new phases + 1 CKM phase)
 - FCNC & EDM experiments severely constrain these
 - some deeper reason (in SUSY br mediation)?
 - Minimal Flavor Violation (MFV)
 - Only CKM phase



Dark Matter (DM)

With R_p conserved, a superpartner (odd) cannot decay to SM (even)

∴ Lightest Supersymmetric Particle (LSP) is stable

LSP is in thermal equilibrium in the early universe, and left over today as Dark Matter

- $\tilde{\chi}_1^0$ LSP with ~ 100 GeV mass is a good DM candidate
 - Weakly Interacting Massive Particle (WIMP)



Dark Matter (DM)

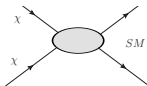
With R_p conserved, a superpartner (odd) cannot decay to SM (even)
 \therefore Lightest Supersymmetric Particle (LSP) is stable

LSP is in thermal equilibrium in the early universe, and left over today as Dark Matter

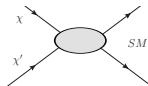
- $\tilde{\chi}_1^0$ LSP with ~ 100 GeV mass is a good DM candidate
 - Weakly Interacting Massive Particle (WIMP)

It's number density can be depleted by annihilating with another sparticle

Self-annihilation



Co-annihilation



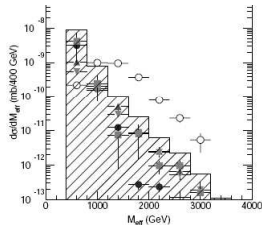
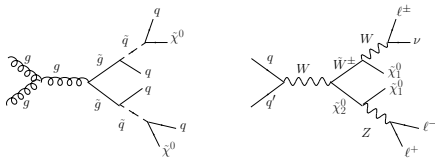
Could be detected in ongoing experiments

- Direct detection (on earth) $\psi N \rightarrow \psi N$
- Indirect detection (cosmic rays) $\psi\psi \rightarrow \gamma\gamma, e^+e^-, \dots$

Aside: Almost pure ν_R , if LSP, could be non-thermal DM

SUSY at LHC

- Cascade decays
- R_p conservation \implies Missing energy signals



[ATLAS Physics TDR]

- Can we determine the spin and couplings to show SUSY?
 - Angular distributions

Conclusions

- Supersymmetry solves the Hierarchy Problem
- MSSM is an effective theory at the TeV scale
 - SUSY breaking in hidden sector
 - Communicated to SM via mediation mechanism (which determine the MSSM parameters)
 - Gravity mediation, Gauge mediation, Anomaly mediation
- R_p conservation
 - Dark Matter Candidate
 - Missing Energy at Colliders
- 14 TeV LHC run crucial for SUSY



BACKUP SLIDES

BACKUP SLIDES

Standard Model (SM) problems

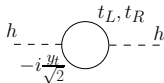
- Gauge hierarchy problem ($M_{EW} \ll M_{Pl}$)
Electroweak scale : $M_{EW} = 10^3$ GeV Gravity scale : $M_{Pl} = 10^{19}$ GeV
 - Higgs sector unstable (quadratic divergence)
- Fermion mass hierarchy problem ($m_e \ll m_t$)
 - Flavor symmetry?
- What is the dark matter
- Inadequate source of CP violation for observed baryon asymmetry
- Cosmological constant problem



Gauge hierarchy problem in detail

$$\mathcal{L}_{SM} \supset -\frac{1}{2}m_h^2 h^2 + \left(-\frac{y_t}{\sqrt{2}} h t_R \bar{t}_L + h.c.\right) + \dots$$

Higgs mass is not protected by any symmetry!



$$\frac{\delta m_h^2}{2} = -\frac{3y_t^2}{8\pi^2} \Lambda^2 \quad (\Lambda \text{ is momentum cut-off})$$

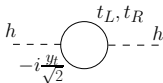
Quadratic divergence! \Rightarrow unnatural (fine-tuning) $m_h = 125 \text{ GeV}$, $\Lambda \sim M_{pl} \approx 10^{19} \text{ GeV}$



Gauge hierarchy problem in detail

$$\mathcal{L}_{SM} \supset -\frac{1}{2}m_h^2 h^2 + \left(-\frac{y_t}{\sqrt{2}} h t_R \bar{t}_L + h.c.\right) + \dots$$

Higgs mass is not protected by any symmetry!



$$\frac{\delta m_h^2}{2} = -\frac{3y_t^2}{8\pi^2} \Lambda^2 \quad (\Lambda \text{ is momentum cut-off})$$

Quadratic divergence! \Rightarrow unnatural (fine-tuning) $m_h = 125 \text{ GeV}$, $\Lambda \sim M_{pl} \approx 10^{19} \text{ GeV}$

New physics (BSM) restores naturalness?

Below what scale (Λ) should it appear?

$$\text{Fine-tuning measure: } f_T \equiv \frac{m_h^2}{\delta m_h^2}$$

$$f_T > 0.1 \implies \Lambda < 2 \text{ TeV} \quad (\text{for } m_h = 120 \text{ GeV})$$

So expect new physics below 2 TeV scale



SM problems cured by Physics Beyond SM (BSM) ?

Some BSM proposals

- Supersymmetry
- Strong dynamics (Technicolor, Composite Higgs)

- **Extra dimensions**

- Flat extra dims
- Warped (AdS space) extra dim

5-D gravity theory in AdS $\overset{\leftarrow}{\text{DUAL}} \overset{\rightarrow}{\text{DUAL}}$ 4-D conformal field theory (CFT)

AdS/CFT correspondence [Maldacena 97]

- Little Higgs



LHC Data (SUSY jets + MET)

