LHC Direct Limits

LHC Indirect Probes

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Non-SUSY BSM and Dark Matter

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IIT Madras, Dec 2018

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# Talk Outline

- Non-SUSY BSM:
  - Composite-Higgs, Little Higgs Models of EWSB (4D/5D Duals)
  - LHC Limits on new vector-bosons, vector-like fermions (VLF), scalars
    - $G'_{\mu\nu}, g'_{\mu}, W'_{\mu}, Z'_{\mu}$
    - $t'_{(2/3)}$ ,  $b'_{(-1/3)}$ ,  $\chi_{(5/3)}$  (Top-partners)
    - Singlet scalar, 2HDM
- Dark Matter (DM)
  - WIMP basics
  - Some DM candidates
    - Higgs-portal
  - Direct, Indirect detection limts, LHC Limits

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Composite-Higgs/Warped Extra-dimension

# General Idea of Composite Higgs

[Georgi, Kaplan 1984]

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- Sector with global symmetry  $\mathcal{G}$ 
  - $\Sigma$  transforms under  ${\cal G}$
- $\langle \Sigma \rangle \neq 0$  such that  $\mathcal{G}$  broken to  $\mathcal{H}$ 
  - (massless) Goldstone Bosons (GB) in coset  $\mathcal{G}/\mathcal{H}$  :  $\pi^a$ 
    - $\pi^a$  are  $\{\phi^{1,2,3}, H, ...\}$  $(\phi^{1,2,3}$  become  $W_{longi}^{\pm}, Z_{longi}$  after EWSB)
    - Note: physical Higgs also a GB
- Gauging  $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$  subgroup & writing Yukawa terms

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# General Idea of Composite Higgs

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    - Note: physical Higgs also a GB
- Gauging  $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$  subgroup & writing Yukawa terms
  - explicitly breaks  $\mathcal{G}$   $\implies$  Higgs gets a mass (at
    - $\implies$  riggs gets a mass (at loop level): Pseudo-GD (PG
  - analogy: (light) Pions are PGB of  $SU(3)_L \otimes SU(3)_R \rightarrow SU(3)$

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#### Composite-Higgs/Warped Extra-dimension

#### 4D model $\leftrightarrow$ Warped model Duality

[Maldacena, 1997] [Arkani-Hamed, Porrati, Randall, 2000; Rattazzi, Zaffaroni, 2001]

#### Dual of Randall-Sundrum model RS1 (SM on IR Brane)

- Planck brane → UV Cutoff; Dynamical gravity in the 4D CFT
- TeV (IR) brane —> IR Cutoff; Conformal invariance broken below a TeV
  - All SM fields are composites of the CFT

#### Dual of Warped Models with Bulk SM

- UV localized fields are elementary
- IR localized fields (Higgs) are composite
  - 4D dual is Composite Higgs model
  - Shares many features with Walking Extended Technicolor
- Partial Compositeness
  - AdS dual is weakly coupled and hence calculable!
- KK states are dual to composite resonances

[Georgi, Kaplan 1984]

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Composite-Higgs/Warped Extra-dimension

# AdS dual of MCHM and Higgs mass

- AdS/CFT Corrsp :  ${\mathcal G}$  global symm of CFT  $\leftrightarrow$  AdS gauge symm
  - Bulk gauge group :  $SO(5) \otimes U(1)_X \qquad A_M = (A_\mu, A_5)$
  - Impose boundary condition (BC) to keep/break a symm:
- $A_5^{\hat{a}}(++)$  dual of PNGB  $\pi^a = \{\phi^{1,2,3}, h\}$  [Contino, Nomura, Pomarol 2003]
  - Gauge symmetry forbids tree-level mass
  - Mass at loop-level from gauge and top loops
     [Hosotani 1983]

#### Minimal Composite Higgs Model (MCHM) dual is [Agashe, Contino, Pomarol, 2004]

Global Symm:  $[SO(5) \otimes U(1)_X]/[SO(4) \otimes U(1)_X]$ Gauge & Yukawa interactions break SO(5)

- 1-loop effective potential gives:  $m_h^2 \sim \frac{2N_c}{N} y_t^2 v^2$  [Coleman, Weinberg 1988]
- Resonances  $m_{res} \sim g_{res} f$   $m_{res} \leftrightarrow m_{KK}$

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# Generating SM mass hierarchy

Bulk Fermions explain SM mass hierarchy [Gherghetta, Pomarol 00][Grossman, Neubert '00]

$$\mathcal{L}_{Yuk}^{(5)} \supset \sqrt{|g|} \left\{ \frac{c_L k \, \bar{\psi}_L \psi_L + \frac{c_R k \, \bar{\psi}_R \psi_R}{c_L k \, \bar{\psi}_R \psi_L H} + h.c. \right\}$$



FCNC largely under control, but still strong constraints [Agashe, Perez, Soni, 2004] LHC Direct Limits

Composite-Higgs/Warped Extra-dimension

### Precision Electroweak Constraints

Precision Electroweak Constraints (S, T,  $Zb\bar{b}$ ) (perturbatively calculable on the warped side)



- Bulk gauge symm  $SU(2)_L imes U(1)$  (SM  $\psi$ , H on TeV Brane)
- T parameter  $\sim (rac{v}{M_{KK}})^2 (k\pi R)$  [Csaki, Erlich, Terning 02]
  - S parameter also  $(k\pi R)$  enhanced
- AdS bulk gauge symm  $SU(2)_R \Leftrightarrow$  CFT Custodial Symm

[Agashe, Delgado, May, Sundrum 03]

- T parameter Protected; S parameter  $\frac{1}{k\pi R}$  for light bulk fermions
- Implies heavy vector bosons:  $W'_{\mu}$ ,  $Z'_{\mu}$ , ...
- Problem: *Zbb* shifted
- 3rd gen quarks (2,2)

[Agashe, Contino, DaRold, Pomarol 06]

- Zbb coupling Protected
- Precision EW constraints  $\Rightarrow$   $M_{KK}\gtrsim$  1.5 2.5 TeV
- Implies top partners: t', b',  $\chi$ , ...

[Carena, Ponton, Santiago, Wagner 06,07]

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### Warped Fermions

- SM fermions : (+,+) BC  $\rightarrow$  zero-mode
- "Exotic" fermions : (-,+) BC  $\rightarrow$  No zero-mode
  - 1<sup>st</sup> KK vectorlike fermion



[Atre et al, '09, '11] [Aguilar-Saavedra, '09] [Mrazek, Wulzer, '09] [SG, Moreau, Singh, '10] [SG, Mandal, Mitra, Tibrewala, '11] [SG, Mandal, Mitra, Moreau : '13]

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The LSS Little-Higgs Model

# LSS Little Higgs Model

Implement collective symmetry breaking

- Higgs is a pseudo Goldstone boson
- no  $\Lambda^2$  divergent contribution at 1-loop
  - Gauge sector & Yukawa couplings specially constructed

A case study: Low, Skiba, Smith (LSS), 2002 : SU(6)/Sp(6)

- Coset:  $35 21 = 14 \rightarrow PNGB$  (Higgs included)
- Gauge sector:  $SU(2)_1 \otimes SU(2)_2 \otimes U(1)_1 \otimes U(1)_2 \rightarrow SU(2)_L \otimes U(1)_Y$
- Vector-like fermions

• SU(2) doublet: 
$$Q' = \begin{pmatrix} t' \\ b' \end{pmatrix}$$
; SU(2) singlets:  $t''$ ,  $b''$ 

Higgs sector is a 2HDM

$$V \supset m_1^2 \left|\phi_1\right|^2 + m_2^2 \left|\phi_2\right|^2 + \left(b^2 \phi_1^T \cdot \phi_2 + h.c.\right) + \lambda_5' \left|\phi_1^T \cdot \phi_2\right|^2$$

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The LSS Little-Higgs Model

# SU(6)/Sp(6) Little Higgs with 2HDM structure

Scalars are: h, H, A,  $H^{\pm}$ ; We focus on neutral scalars

- Seek hWW, hZZ SM like
  - Alignment limit:  $(\beta \alpha) \approx \frac{\pi}{2}$ ; [Gunion, Haber, 2002]
    - *HWW*,  $HZZ \approx \text{zero!}$
- Also seek htt SM like
- AWW, AZZ are zero at tree-level :  $\mathcal{CP}$  inv

$$\langle \phi_1 \rangle = \frac{1}{\sqrt{2}} \left( \begin{array}{c} 0 \\ v_1 \end{array} \right); \quad \langle \phi_2 \rangle = \frac{1}{\sqrt{2}} \left( \begin{array}{c} v_2 \\ 0 \end{array} \right); \quad \tan\beta = \frac{v_1}{v_2}$$

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The LSS Little-Higgs Model

# VLF properties and fine-tuning

[SG, T. Mukherjee, S. Sadhukhan: PRD 94, 015034 (2016)]



General aspects

# Vector-like fermion (VLF) decoupling

- VLF has independent source of mass M (not given by  $m = \lambda v$ )
  - Can make *M* arbitrarily large
    - Yukawa coupling can remain perturbative
    - As M increases, fine-tuning increases
  - Nice decoupling behavior : S,T, U,  $h \rightarrow \gamma \gamma$ ,  $gg \rightarrow h$ , ...



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#### General aspects

### Generic signatures

#### New particles

- Vector resonances
- Fermion resonances
- Bigger coset  $\implies$  extra (pNBG) scalars (can be "light")

• Eg: 
$$SU(6)/Sp(6) \implies 2HDM$$

Observables

• Precision electroweak probes (tree-level!)

• 
$$S pprox 1.3 \, (4\pi) \, (v/m_{KK})^2 \quad \Longrightarrow \quad m_{KK} \gtrsim 2 \; TeV$$

- LHC signals
  - Direct

• Indirect Eg: *hVV* coupling shifts  $\frac{hVV_{BSM}}{hVV_{SM}} \approx (1 - \kappa v^2/f^2)$ 

FCNC probes

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General aspects

# Decay Modes of t', b', $\chi$

• VL Tree-level Decays

- b' 
  ightarrow tW , b' 
  ightarrow bZ , b' 
  ightarrow bh
- t' 
  ightarrow bW , t' 
  ightarrow tZ , t' 
  ightarrow th
- $\chi \to tW$

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#### Search Limits

### Vector-like fermion (t',b') search



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Search Limits

# graviton<sup>(1)</sup> $\rightarrow WW, ZZ$ search



[ATLAS: 1808.02380; PRD 2018]

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#### Search Limits

$$W_{R,L}^\prime o t ar b$$
 search



[ATLAS: 1807.10473; PLB 2018]

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#### Search Limits

# $gluon^{(1)} \rightarrow t\bar{t}$ search



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#### Search Limits

 $Z' \rightarrow t\bar{t}$  search



[CMS:1810.0590]

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#### Search Limits

$$\mathsf{Z}' 
ightarrow \ell^+ \ell^-$$



[Agashe, Davoudiasl, SG, Han, Huang, Perez, Si, Soni: 0709.0007 [hep-ph]]

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# VECTOR-LIKE FERMIONS IN ELECTROWEAK PRECISION & HIGGS OBSERVABLES

[S.Ellis, R.Godbole, SG, J.Wells; 1404.4398 [hep-ph], JHEP 09 (2014) 130]

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EW Precision & Higgs coupling probes

# EWPrecision + Higgs Observables

Precision electroweak observables (S, T, U)



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Modifications to hgg,  $h\gamma\gamma$  couplings: $\sigma(gg \rightarrow h)$  $\Gamma(h \rightarrow \gamma\gamma)$ 



We compute ratios  $\frac{\Gamma_{h \to gg}}{SM}$ ,  $\frac{\Gamma_{h \to \gamma\gamma}}{SM}$  using leading-order expressions QCD corrections to ratios small: [Furlan '11] [Gori, Low '13]

$$\mu_{\gamma\gamma}^{VBF} \approx \frac{\Gamma_{\gamma\gamma}}{\Gamma_{\gamma\gamma}^{SM}} ; \quad \mu_{ZZ}^{ggh} \approx \frac{\Gamma_{gg}}{\Gamma_{gg}^{SM}} ; \quad \mu_{\gamma\gamma}^{ggh} \approx \frac{\Gamma_{gg}}{\Gamma_{gg}^{SM}} \frac{\Gamma_{\gamma\gamma}}{\Gamma_{\gamma\gamma}^{SM}} ; \quad \frac{\mu_{\gamma\gamma}^{ggh}}{\mu_{ZZ}^{ggh}} \approx \frac{\Gamma_{\gamma\gamma}}{\Gamma_{\gamma\gamma}^{SM}} \approx \mu_{\gamma\gamma}^{VBF}$$







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 $\lambda_D=$  1,  $M_D=M_Q$ ,  $Y_Q=(1/6,-1/6)$  (solid, dashed)

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#### EW Precision & Higgs coupling probes

### Q + U model



[Q+U model from MVQD model with  $Y_{\chi} = -1/6$ ]

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EW Precision & Higgs coupling probes

### LHC constraints on Higgs couplings



[ATLAS-CONF-2018-31] [CMS-HIG-17-031]

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Dark Matter ●	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection
Particle DM				
Particle	Dark Matter			

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- Thermal Relic
  - In thermal equilibrium in early universe
  - Details of its origin do NOT matter
    - So most studied
- Nonthermal Relic
  - Never in thermal equilibrium
  - Details of its origin matter

Dark Matter o	Particle Dark Matter ●00000	Direct Detection	Indirect detection	Collider detection
Some Candidates				
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### Some Particle DM Candidates

- LSP Lightest Supersymmetric Particle
  - $\tilde{\chi}_1^0$  Neutralino (SUSY partner of neutral gauge boson)
  - $\tilde{\nu}$  Sneutrino (SUSY partner of neutrino)
- Hidden sector DM
- LKP Lightest Kaluza-Klein Particle Extra space dimensions
- LTP Lightest T-odd Particle Little-higgs theory with Z<sub>2</sub>
- SuperWIMP Gravitino (SUSY partner of graviton)
- E-WIMP Right-handed sneutrino (partner of neutrino)
- WIMPzilla Extremely massive particle
- Your candidate here

Dark Matter o	Particle Dark Matter ●00000	Direct Detection	Indirect detection	Collider detection
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Some Candidates				

### Weakly Interacting Massive Particle (WIMP)

WIMP Cold dark matter - New particle

• Mass:  $M \sim 100 \, {
m GeV}$ 

"WIMP Miracle"

• Interaction strength:  $g \sim g_{EW}$ 

$$\Omega_0 \equiv \frac{n_0 M}{\rho_c} \approx 4 \times 10^{-10} \left( \frac{\mathrm{GeV}^{-2}}{\langle \sigma v \rangle} \right)$$

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• Precisely the scale being explored at colliders!

• DM at present colliders? (LHC connection)

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Some Candidates				
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Some Candidates						
Nontherm	al $ ilde{ u}_0$ DM					

[de Gouvea, SG, Porod 2006]

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If thermalization conditions not met  $\Rightarrow$  Nonthermal  $\tilde{\nu}_0$ 

Happens when :

- $Y_N \lesssim 10^{-6}$  i.e.,  $ilde{
  u}_0$  is almost pure right-handed
- Low Reheat temp  $T_{RH} <$  100 GeV ; Reheat into  $ilde{
  u}_0$  + SM

No Relic-from-decay of heavier SUSY particles No  $\tilde{\nu}_0$  thermalization from co-interaction with SUSY or Top

 $ilde{N}$  Relic density depends on Inflaton coupling to SM and  $ilde{N}$ 

Dark Matter 0	Particle Dark Matter 00●000	Direct Detection	Indirect detection	Collider detection		
Some Candidates						
Nontherm	al $ ilde{ u}_0$ DM					

[de Gouvea, SG, Porod 2006]

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Dark Matter 0	Particle Dark Matter 000●00	Direct Detection	Indirect detection	Collider detection
Some Candidates				
Hidden se	ector DM $\psi$			

[SG, Lee, Wells 2009]



Channels  $\psi\psi
ightarrow bar{b},W^+W^-\,,$  ZZ , hh ,  $tar{t}$ 

Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection
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Hidden s	ector DM $\psi$			

[SG, Lee, Wells 2009]

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Channels  $\psi \psi \rightarrow b \bar{b}, W^+ W^-, ZZ, hh, t\bar{t}$ 

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Some Candidat	es				
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#### Direct Detection of Hidden sector



Effective  $h\bar{N}N$  coupling  $pprox 2 imes 10^{-3}$  [Shifman, Vainshtein, Zakharov (1973)]

 $\psi\text{-}$  Nucleon c.s. :

$$\sigma\left(\psi N \to \psi N\right) \approx \frac{\kappa_{11}^2 s_h^2 c_h^2 \lambda_N^2}{8\pi v_{rel}} \frac{\left(\left|\mathbf{p}_{\psi}\right|^2 + m_N^2\right)}{(t - m_h^2)^2}$$

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Dark Matter	Particle Dark Matter	Direct Detection	Indirect detection	Collider detection
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Some Candidates				

# Singlet fermion DM with scalar mediator



For  $s_h \ll 1$ , loop suppressed coupling can give correct relic density direct-detection rate can be very small - experimentally very challenging

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Oark Matter	Particle Dark Matter 000000	Direct Detection ●○	Indirect detection	Collider detection
Direct Detection L	imits			
Direct D	etection Exper	rimental Limi	ts	



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Dark Matter	Particle Dark Matter	Direct Detection	Indirect detection	Collider detection
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#### Direct Detection Limits

### Direct Detection Spin Dependent Limits



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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection ●○○	Collider detection
AMS-02				
$e^+$ , $\bar{p}$ da	ita			



AMS02 e<sup>+</sup> [PRL113, 2014] [PRL117,2016];

PAMELA p [Nature 09]

Dark matter annihilations? .... or astro sources (pulsars) ? ... or other bkgnd?

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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection ○●○	Collider detection
Fermi-LAT				
Fermi-L	AT $\gamma$			



[Fermi-LAT Galactic Excess: 1704.03910]

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Dark Matter	Particle Dark Matter	Direct Detection	Indirect detection	Collider detection
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Fermi-LAT				

### Fermi-LAT DSph Limits



Figure 9. Upper limits (95% confidence level) on the DM annihilation cross section derived from a combined analysis of the nominal target sample for the  $b\bar{b}$  (left) and  $\tau^{+}\tau^{-}$  (right) channels. Bands for the expected sensitivity are calculated by repeating the same analysis on 300 randomly selected sets of high-Galactic-latitude blank fields in the LAT data. The dashed line shows the median expected sensitivity while the bands represent the 68% and 95% quantiles. Spectroscopically measured J-factors are used when available; otherwise, J-factors are predicted photometrically with an uncertainty of 0.6 dex (solid red line). The solid black line shows the observed limit from the combined analysis of 15 dSphs from Ackermann et al. (2015b). The closed contours and marker show the best-fit regions (at 2 $\sigma$  confidence) in cross-section and mass from several DM interpretations of the GCE: green contour (Gordon & Macias 2013), red contour (Daylan et al. 2016), orange data point (Abazajian et al. 2014), purple contour (Calore et al. 2015). The dashed gray curve corresponds to the thermal relic cross section from Stegman et al. (2012).

[Fermi-LAT: 1611.03184]

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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection			
LHC Signals and Limits							
LHC DN	/ signatures						

DM leads to missing energy  $(\not\!\!\!E_T)$  events at the LHC

Look in  $pp \rightarrow \gamma + \not\!\!\! E_T$ ,  $j + \not\!\!\! E_T$ ,  $V_\mu + \not\!\!\! E_T$ ,  $h + \not\!\!\! E_T$ , ...





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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection
LHC Signals and Limi	ts			
Effective (	Operators			

[Goodman, Ibe, Rajaraman, Shepherd, Tait, Yu 1008.1783]

Name	Operator	Coefficient
D1	$\bar{\chi}\chi\bar{q}q$	$m_q/M_*^3$
D2	$\bar{\chi}\gamma^5\chi\bar{q}q$	$im_q/M_*^3$
D3	$\bar{\chi}\chi\bar{q}\gamma^5q$	$im_q/M_*^3$
D4	$\bar{\chi}\gamma^5\chi\bar{q}\gamma^5q$	$m_q/M_*^3$
D5	$\bar{\chi}\gamma^{\mu}\chi\bar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D6	$\bar{\chi}\gamma^{\mu}\gamma^{5}\chi\bar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D7	$\bar{\chi}\gamma^{\mu}\chi\bar{q}\gamma_{\mu}\gamma^{5}q$	$1/M_{*}^{2}$
D8	$\bar{\chi}\gamma^{\mu}\gamma^{5}\chi\bar{q}\gamma_{\mu}\gamma^{5}q$	$1/M_{*}^{2}$
D9	$\bar{\chi}\sigma^{\mu\nu}\chi\bar{q}\sigma_{\mu\nu}q$	$1/M_{*}^{2}$
D10	$\bar{\chi}\sigma_{\mu\nu}\gamma^5\chi\bar{q}\sigma_{lphaeta}$ q	i/M <sup>2</sup> *
D11	$\bar{\chi}\chi G_{\mu\nu}G^{\mu\nu}$	$\alpha_s/4M_*^3$
D12	$\bar{\chi}\gamma^5\chi G_{\mu\nu}G^{\mu\nu}$	$i\alpha_s/4M_*^3$
D13	$\bar{\chi}\chi G_{\mu\nu}\tilde{G}^{\mu\nu}$	$i\alpha_s/4M_*^3$
D14	$\bar{\chi}\gamma^5\chi G_{\mu\nu}\tilde{G}^{\mu\nu}$	$\alpha_s/4M_*^3$

Name	Operator	Coefficient
C1	$\chi^{\dagger}\chi\bar{q}q$	$m_q/M_*^2$
C2	$\chi^{\dagger} \chi \bar{q} \gamma^{5} q$	$im_q/M_*^2$
C3	$\chi^{\dagger} \partial_{\mu} \chi \bar{q} \gamma^{\mu} q$	$1/M_{*}^{2}$
C4	$\chi^{\dagger}\partial_{\mu}\chi\bar{q}\gamma^{\mu}\gamma^{5}q$	$1/M_{*}^{2}$
C5	$\chi^{\dagger}\chi G_{\mu u}G^{\mu u}$	$\alpha_s/4M_*^2$
C6	$\chi^{\dagger}\chi G_{\mu u}\tilde{G}^{\mu u}$	$i\alpha_s/4M_*^2$
R1	$\chi^2 \bar{q} q$	$m_q/2M_*^2$
R2	$\chi^2 \bar{q} \gamma^5 q$	$im_q/2M_*^2$
R3	$\chi^2 G_{\mu\nu} G^{\mu\nu}$	$\alpha_s/8M_*^2$
R4	$\chi^2 G_{\mu\nu} \tilde{G}^{\mu\nu}$	$i\alpha_s/8M_*^2$

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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection
LHC Signals and	Limits			
ATLAS	DM Limits			



[ATLAS-CONF-2018-051]

Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection
LHC Signals and I	Limits			
CMS DI	M Limits			



[CMS: PRD 97, 2018]

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Dark Matter 0	Particle Dark Matter 000000	Direct Detection	Indirect detection	Collider detection			
LHC Signals and Limit	LHC Signals and Limits						
Conclusions							

- Many compelling arguments why standard model is not the full story
  - possible new physics that stabilize the Higgs sector include composite-Higgs, Little Higgs models
  - so far no definitive signals of BSM at LHC
    - direct searches probing interesting region now!
    - hgg,  $h\gamma\gamma$  couplings may show deviations
- Strong evidence for Dark Matter. But what is it?
  - Many Candidates, Mass range  $10^{-5}$ eV  $10^{12}$ GeV
    - Vigorous search underway Direct, Indirect, Collider

• We may discover DM particle soon

# **BACKUP SLIDES**

#### **BACKUP SLIDES**

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# Fermion rep : $Zb\bar{b}$ not protected (DT model)

[Agashe, Delgado, May, Sundrum '03]

• Complete *SU*(2)<sub>*R*</sub> multiplet

• 
$$Q_L \equiv (2, 1)_{1/6} = (t_L, b_L)$$
  
 $\psi_{t_R} \equiv (1, 2)_{1/6} = (t_R, b')$   
 $\psi_{b_R} \equiv (1, 2)_{1/6} = (T, b_R)$ 

- "Project-out" b', T zero-modes by (-, +) B.C.
- New  $\psi_{VL}$  : b', T
- $b \leftrightarrow b'$  mixing
  - Zbb coupling shifted
    - So LEP constraint quite severe

# Fermion rep : $Zb\bar{b}$ protected (ST & TT models)

• 
$$Q_L = (2,2)_{2/3} = \begin{pmatrix} t_L & \chi \\ b_L & T \end{pmatrix}$$

[Agashe, Contino, DaRold, Pomarol '06]

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•  $Zb_L\overline{b_L}$  protected by custodial  $SU(2)_{L+R} \otimes P_{LR}$  invariance  $Wt_Lb_L$ ,  $Zt_Lt_L$  not protected, so shifts

Two t<sub>R</sub> possibilities:

I Singlet  $t_R$  (ST Model) :  $(1,1)_{2/3} = t_R$  New  $\psi_{VL}$  :  $\chi$ , T

2 Triplet  $t_R$  (TT Model) :

$$(1,3)_{2/3} \oplus (3,1)_{2/3} = \psi_{t_R}' \oplus \psi_{t_R}'' = \begin{pmatrix} \frac{t_R}{\sqrt{2}} & \chi' \\ b' & -\frac{t_R}{\sqrt{2}} \end{pmatrix} \oplus \begin{pmatrix} \frac{t}{\sqrt{2}} & \chi'' \\ b'' & -\frac{t'}{\sqrt{2}} \end{pmatrix}$$
  
New  $\psi_{VL} : \chi, T, \chi', b', \chi'', t'', b'$ 

# Fermion rep : $Zb\overline{b}$ protected (ST & TT models)

• 
$$Q_L = (2,2)_{2/3} = \begin{pmatrix} t_L & \chi \\ b_L & T \end{pmatrix}$$
 [Agashe, Contino, DaRold, Pomarol '06]

•  $Zb_L\overline{b_L}$  protected by custodial  $SU(2)_{L+R} \otimes P_{LR}$  invariance  $Wt_Lb_L$ ,  $Zt_Lt_L$  not protected, so shifts

#### Two $t_R$ possibilities:

- Singlet  $t_R$  (ST Model) :  $(1,1)_{2/3} = t_R$  New  $\psi_{VL}$  :  $\chi$ , T
- Triplet t<sub>R</sub> (TT Model) :

$$(1,3)_{2/3} \oplus (3,1)_{2/3} = \psi_{t_R}' \oplus \psi_{t_R}'' = \begin{pmatrix} \frac{t_R}{\sqrt{2}} & \chi' \\ b' & -\frac{t_R}{\sqrt{2}} \end{pmatrix} \oplus \begin{pmatrix} \frac{t'}{\sqrt{2}} & \chi'' \\ b'' & -\frac{t'}{\sqrt{2}} \end{pmatrix}$$
  
New  $\psi_{VL} : \chi, T, \chi', b', \chi'', t'', b'$ 

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### EWSB region

[Contino, Da Rold, Pomarol 0612048]



Figure 1: Contour plots of  $\epsilon$  in the plane  $(c_q, c_u)$  with  $\tilde{m}_u = 1$ ,  $\tilde{M}_u = -2$ , N = 8 for the  $MCHM_{10}$ (left plot), and in the plane  $(\tilde{m}_u, \tilde{M}_u)$  with  $c_q = 0.35$ ,  $c_u = 0.45$ , N = 8 for the  $MCHM_5$  (right plot). The two gray areas correspond to the region with EWSB and non-zero fermion masses,  $0 < \epsilon < 1$ . The lighter gray area is excluded when the bound  $S \lesssim 0.3$  is imposed. The dashed black line represents the curve with  $m_t^{\overline{MS}}(2 \text{ TeV}) = 150 \text{ GeV}$ , equivalent to  $m_t^{pole} = 173 \text{ GeV}$ .

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### Warped model t' parameters

[SG, T.Mandal, S.Mitra, R.Tibrewala, arXiv:1107.4306] [SG, T.Mandal, S.Mitra, G.Moreau : arXiv:1306.2656]



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# Warped model t' BR



# Custodial protected MCHM (5,10 reps)

- $\alpha$  : gauge loops,  $q_L$  or  $q_R$
- $\beta$  :  $q_L$  and  $q_R$

 $\langle s_h 
angle = \epsilon = \sqrt{rac{eta - lpha}{2eta}}$  So  $\epsilon$  is a measure of fine-tuning

$$m_h^2 \approx \frac{N_c m_t^2 \epsilon^2 \Lambda^2}{\pi^2 v^2}$$

$$S=rac{3}{8}rac{N}{\pi}\epsilon^2$$
  $S\lesssim 0.3$   $\Longrightarrow$   $f\gtrsim v\sqrt{rac{10}{8}rac{N}{\pi}}$ 

 $m_h^2 \gtrsim rac{N_c}{\pi^2} rac{m_t^2}{f^2} m_Q^2 \implies m_Q \lesssim 1 \, TeV\left(rac{m_h}{125 \, GeV}
ight) \left(rac{160 \, GeV}{m_t}
ight) \left(rac{f}{750 \, GeV}
ight)$ 

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$$\Sigma = \exp\left(\frac{i\pi^{a}X^{a}}{f}\right) \langle \Sigma \rangle ; \qquad \langle \Sigma \rangle = \begin{pmatrix} 0 & -\mathbb{1} \\ \mathbb{1} & 0 \end{pmatrix}$$
$$\pi^{a}X^{a} \supset \begin{pmatrix} 0 & 0 & -1 \\ 0 & 0 & -1 \\ 0 & 0 & -1 \\ \phi_{2}^{\dagger} & 0 & -\phi_{1}^{T} & 0 \\ \phi_{2}^{\dagger} & 0 & -\phi_{1}^{T} & 0 \\ 0 & -s^{*} & -\phi_{1}^{*} & 0 & 0 \\ s^{*} & 0 & -\phi_{1}^{*} & 0 & 0 \\ \phi_{1}^{\dagger} & 0 & \phi_{2}^{T} & 0 \end{pmatrix}$$

Gauge group:  $SU(2)_1 \otimes SU(2)_2 \otimes U(1)_1 \otimes U(1)_2$  with  $Q_1^{\mathfrak{s}} = \begin{pmatrix} \sigma^{\mathfrak{s}} & 0 \\ 0 & 0 \end{pmatrix}$ ;  $Q_2^{\mathfrak{s}} = \begin{pmatrix} 0 & 0 \\ 0 & \sigma^{\mathfrak{s}} \end{pmatrix}$ BSM vector-like fermions: 1 VL doublet (Q') + 2 VL singlets (t'', b'')1-loop effective potential  $\implies$  potential for PNGB, EWSB

# Effects from VLF on A, H



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### Computing Thermal DM Relic Density: Boltzmann Eqn

$$\begin{array}{l} \frac{d}{dt}n = -3Hn - \langle \sigma v \rangle_{SI} \left( n^2 - n_{eq}^2 \right) - \langle \sigma v \rangle_{CI} \left( nn_{\phi} - n_{eq}n_{\phi \ eq} \right) + C_{\Gamma} \\ \text{Thermal equilibrium if : } \langle \sigma v \rangle_{SI} n_{\tilde{\nu}_0} > 3H \quad \text{OR} \quad \langle \sigma v \rangle_{CI} n_{\phi} > 3H \end{array}$$

$$n_{eq} = \begin{cases} (\zeta(3)/\pi^2)gT^3 & T \gg M \\ g \left(\frac{MT}{2\pi}\right)^{3/2} e^{-(M-\mu)/T} & T \lesssim M \end{cases}$$



[Kolb & Turner, Early Universe]

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# Thermal (Mixed) $\tilde{\nu}_0$ DM



$$\Omega_0 h^2 = \frac{10^{-4}}{s_1^4} \left\{ \left[ g^2 \left( \frac{100 \text{ GeV}}{M_{\tilde{W}}} \right) + g'^2 \left( \frac{100 \text{ GeV}}{M_{\tilde{B}}} \right) \right]^2 \right\}^{-1}$$

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 $\therefore$   $s_1 \approx 0.2$  results in observed relic density

# Thermal avg. c.s. $\langle \sigma v \rangle$



Co-annihilation and/or decays into DM may also be important

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# Thermal avg. c.s. $\langle \sigma v \rangle$



Co-annihilation and/or decays into DM may also be important

### $\psi$ Relic Density + Direct Detection



[SG, S.Lee, J.Wells 2009]

 $M_{\psi} = 250 \, GeV, m_h = 120 \, GeV, \kappa_{11} = 2.0, s_h = 0.25, \kappa_{3\phi} = 1, m_H = 1 \, TeV$ Shaded:

 $\sigma_{Dir}\gtrsim 10^{-43}~{
m cm}^2$  (dark);  $\gtrsim 10^{-44}~{
m cm}^2$  (medium);  $\gtrsim 10^{-45}~{
m cm}^2$  (light)

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### Higgs decay to DM

- Higgs decay and BR
  - If  $m_h > 2M_{\psi}$  :  $h \to \psi \overline{\psi}$  Invisible Decay!
    - Decay channels:  $h \rightarrow \psi \bar{\psi} , \ b \bar{b} , \ WW , \ ZZ , \ t \bar{t}$



 $M_{\psi} \approx 59 \, GeV, s_h = 0.25, \kappa_{11} = 2.0, \kappa_{3\phi} = 1.0, m_H = 1 \, TeV$ 

NB: Relic density not enforced

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# Hidden sector DM @ 14 TeV LHC



[O. J. P. Eboli and D. Zeppenfeld, 2000]

$$\begin{split} p_T^J &> 40 \ , \ |\eta_j| < 5.0 \ , \ |\eta_{j_1} - \eta_{j_2}| > 4.4 \ , \ \eta_{j_1} \cdot \eta_{j_2} < 0 \ , \\ \phi_T &> 100 \ {\rm GeV} \ , \ M_{ii} > 1200 \ {\rm GeV} \ , \ \phi_{ii} < 1 \ . \end{split}$$

For 
$$s_h = 0.25$$
,  $BR_{INV} = 0.25$ :

$m_h$ (GeV)	$\sigma_S BR_{inv}(fb)$	$\sigma_B(fb)$	$\mathcal{L}_{5\sigma}$ (fb <sup>-1</sup> )
120	22.7	167	8
200	18	167	12.8
300	13.2	167	23.7

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