

# Asymptotic Expansion of $\mathcal{N} = 4$ Dyon Degeneracy

**Nabamita Banerjee**



**Harish-Chandra Research Institute, Allahabad, India**

**Collaborators: D. Jatkar, A.Sen**

**References: (1) arXiv:0807.1314 [hep-th]**

**(2) arXiv:0810.3472 [hep-th]**

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## Plan of the talk

- 1 Introduction
- 2 Setup
- 3 Macroscopic Understanding
- 4 Result

## Motivation

- Black Holes are solutions of Einstein-Maxwell theory (low energy limit of string theory). They carry certain charges and quantum mechanically behave as thermodynamic objects.
  - They can also be described in terms of specific configuration of states in the full string theory, carrying similar set of charges.
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- We want to understand the statistical origin of Black Hole entropy, as the logarithm of the degeneracy of these states.

- Major success: for extremal Black Holes,

$$S_{BH} = S_{stat} \equiv \ln d(\vec{Q}), \quad \text{in large charge}(\vec{Q}) \text{ limit.}$$

- Can we go beyond large charge limit?

$$i.e. \quad S_{BH} \stackrel{?}{=} \ln d(\vec{Q}),$$

for large but finite charges.

To do this, we need to take two way approach to the problem.

## Gravity Side

- On gravity side, we need to consider all **higher derivative and quantum corrections**.
  - Higher derivative terms: Entropy function technology can be used.
  - Quantum corrections: Quantum entropy function can be used.
- Entropy corrections come as an expansion in inverse power of charges.

## String Theory Side

- We need to compute the degeneracy of states more accurately.

## GOAL

To understand these corrections to the entropy in the statistical side by doing systematic asymptotic expansion of the degeneracy function.

## Setup

### Theory

- We consider  $\mathcal{N} = 4$  superstring theory with rank  $r$  gauge group.
- At a generic point in the moduli space, the unbroken gauge group is  $U(1)^r$ .
- The low energy SUGRA theory has a continuous  $SO(6, r - 6) \times SL(2, R)$  symmetry.
- We denote the  $SO(6, r - 6)$  invariant metric by  $L$ . All inner products are defined with respect to  $L$ .

## Two Descriptions

### First Description

- Type IIB string theory on  $K3 \times S^1 \times \tilde{S}^1 / \mathbb{Z}_N$ .
- A dyonic state in this theory is a particular brane configuration.

### Second Description

- Equivalently heterotic string theory on  $T^4 \times S^1 \times \hat{S}^1 / \mathbb{Z}_N$
- A dyonic state in this theory is described by a state carrying some electric and magnetic charges.

## Duality

- The two descriptions of the theory are related by a chain of duality transformations as follows:

$$\left( \begin{array}{c} \text{IIB} \\ \mathbf{S}^1 \times \mathbf{S}^1 \end{array} \right) \xrightarrow{\mathbf{S}} \left( \begin{array}{c} \text{IIB} \\ \mathbf{S}^1 \times \mathbf{S}^1 \end{array} \right) \xrightarrow{\mathbf{T}} \left( \begin{array}{c} \text{IIA} \\ \mathbf{S}^1 \times \mathbf{S}^1 \end{array} \right) \xrightarrow{\text{st-st}} \left( \begin{array}{c} \text{Heterotic} \\ \mathbf{T}^6 \end{array} \right)$$

## Charge Vectors

- In general, any given state is characterized by  $r$  dimensional electric and magnetic charge vectors,  $\vec{Q}$  and  $\vec{P}$ .
- The T-duality invariants are,

$$Q^2 = Q^T L Q \quad P^2 = P^T L P \quad Q \cdot P = Q^T L P,$$

where  $L$ , T-duality matrix.

## Macroscopic Side

- We will consider quarter BPS dyonic Black Holes in the Heterotic theory.
- Restricting to Supergravity approximation, we can find the entropy carried by these Black Holes .

## Microscopic Side

- By duality, we can also regard the states associated with this Black Holes as states of some particular quarter BPS D-brane configuration in the type IIB theory.
- Considering the dynamics of various fields in the D-brane configuration, the complete degeneracy function has been evaluated (JHEP 0611:072,2006).

## Degeneracy Formula

- The microscopic degeneracy is,

$$d(\vec{Q}, \vec{P}) = (-1)^{Q \cdot P + 1} A \int_{\mathcal{C}} d\rho d\sigma d\nu \frac{e^{-\pi i(\rho Q^2 + \sigma P^2 + 2\nu Q \cdot P)}}{\Phi(\rho, \sigma, \nu)}$$

- Contour  $\mathcal{C}$  is a three real dimensional subspace of the complex dimensional space labeled by  $(\rho, \sigma, \nu)$ .

- For  $N = 1$  theory, the function  $\Phi(\rho, \sigma, \nu)$  is a modular form of weight 10.
- The analogous modular forms are also known for many other models . (Shamik's Talk)

## Asymptotic Expansion

### A .

- For a given set of charges, there are single centered and multi centered Black Hole solutions.
- We are Interested in single centered Black Hole entropy.
- We organize the integral such that the result can pick up the contribution from single centered Black holes. This is done by choosing the integration contour  $\mathcal{C}$  in a specific way.
- In particular, we need to set the asymptotic values of the moduli fields equal to their attractor values.

**B .**

- We have to do three integrals, over  $(\rho, \sigma, \nu)$ . For this, we need the pole of the integrand.
- The function has a second order zero at,

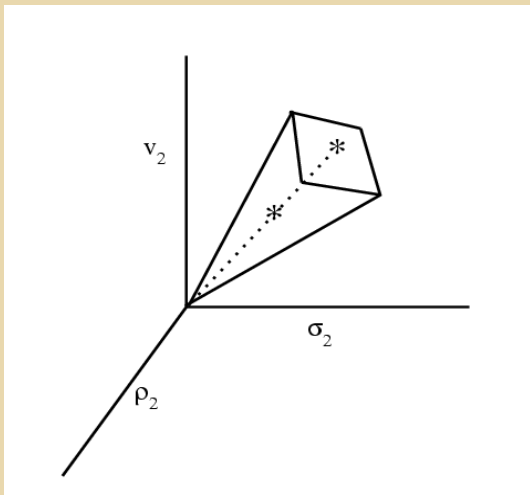
$$n_2(\sigma\rho - \nu^2) + j\nu + n_1\sigma - m_1\rho + m_2 = 0$$

for

$$m_1, n_1, m_2, n_2 \in \mathbb{Z}, j \in 2\mathbb{Z} + 1, \quad m_1 n_1 + m_2 n_2 + \frac{j^2}{4} = \frac{1}{4}$$

- We consider cases with  $n_2 \geq 1$ .

## Pole structure



**B.1**

- In saddle point approximation of the integral, the degeneracy is,

$$d(\vec{Q}, \vec{P}) = e^{\pi \sqrt{Q^2 P^2 - (Q \cdot P)^2} / n_2}$$

- For  $n_2 = 1$ , we get the maximum contribution to the degeneracy.
- For  $n_2 \geq 2$ , the degeneracy is exponentially suppressed compared to the leading one. Hence, to compute exponentially suppressed contribution, we need to look at these sub-leading poles.

## B.2

- For the leading pole, one integral can be done by residue method. The other two integrals are done by saddle point analysis.
- The  $v$  integral is done by residue method. Near the pole, the function  $\Phi$  behaves as.

$$\Phi(\rho, \sigma, v) \rightarrow v^2 g(\rho)g(\sigma) + O(v^4)$$

- The  $(\rho, \sigma)$  integral takes the form,

$$e^{S_{stat}(\vec{Q}, \vec{P})} = d(\vec{Q}, \vec{P}) \sim \int \frac{d^2\tau}{\tau_2^2} e^{F(\vec{\tau})}$$

where,  $\rho = \tau_1 + i\tau_2$  and  $\sigma = -\tau_1 + i\tau_2$

- The function  $F(\tau)$  can be easily computed after doing the  $v$  integral.

- This can be regarded as a zero dimensional field theory with fields  $(\tau_1, \tau_2)$  with action  $F(\vec{\tau}) - 2 \ln \tau_2$ .
- The result for statistical entropy  $S_{stat}$  can be obtained by computing the possible diagrams of this field theory up to any desired order in charges.
- This will produce all sub-leading correction in inverse power of charges.

Up to  $1/\text{charge}^2$  corrections

## Statistical Entropy Function

- Leading result:

$$\Gamma_0(\vec{\tau}_B) = -\frac{\pi}{2\tau_{B2}} |\mathbf{Q} - \tau_B \mathbf{P}|^2$$

- The  $1/\text{charge}$  correction:

$$\Gamma_1(\vec{\tau}_B) = \ln g(\tau_B) + \ln g(-\bar{\tau}_B) + (k+2) \ln(2\tau_{B2})$$

The  $1/\text{charge}^2$  correction :

$$\Gamma_2(\vec{\tau}_B) = \ln d_2(\vec{Q}, \vec{P}) = -\frac{\tau_{2B}}{\pi |\mathbf{Q} - \tau_B \mathbf{P}|^2} (k+2)$$

Up to  $1/\text{charge}^2$  corrections

## Statistical Entropy

- Leading entropy:  $S^{(0)} = \pi \sqrt{Q^2 P^2 - (Q \cdot P)^2}$
- The  $1/\text{charge}$  correction:

$$S^{(1)} = -\ln g(\tau_{(0)}) - \ln g(-\bar{\tau}_{(0)}) - (k+2) \ln(2\tau_{(0)_2})$$

- The  $1/\text{charge}^2$  correction :

$$S^{(2)} = \frac{2+k}{2\pi \sqrt{Q^2 P^2 - (Q \cdot P)^2}} + \frac{4\tau_{(0)_2^3}}{\pi |Q - \tau_{(0)} P|^2}$$

$$\left[ \left( \frac{g'(\tau_{(0)})}{g(\tau_{(0)})} + \frac{k+2}{\tau_{(0)} - \bar{\tau}_{(0)}} \right) \left( \frac{g'(-\bar{\tau}_{(0)})}{g(-\bar{\tau}_{(0)})} + \frac{k+2}{\tau_{(0)} - \bar{\tau}_{(0)}} \right) \right]$$

## Exponential Suppressed Correction

- The zeros of the function  $\Phi$ ,

$$n_2(\sigma\rho - v^2) + jv + n_1\sigma - m_1\rho + m_2 = 0$$

- We get such corrections from the sub-leading poles,  
 $n_2 \geq 2$ .

- For this we define

$$\Omega = \begin{bmatrix} \rho & v \\ v & \sigma \end{bmatrix}$$

- We look for a symplectic transformation of the form:

$$\begin{bmatrix} \hat{\rho} & \hat{\nu} \\ \hat{\nu} & \hat{\sigma} \end{bmatrix} \equiv \hat{\Omega} = (A \Omega + B)(C \Omega + D)^{-1},$$

such that

$$\hat{\nu} = \frac{n_2(\sigma\rho - \nu^2) + j\nu + n_1\sigma - m_1\rho + m_2}{\det(C\Omega + D)}$$

- The behavior of  $\Phi_k$  near the zero is,

$$\Phi_k(\rho, \sigma, \nu) = -\{\det(C \Omega + D)\}^{-k} 4\pi^2 \hat{\nu}^2 g(\hat{\rho}) g(\hat{\sigma}) + O(\nu^4)$$

## Degeneracy Formula

- The degeneracy formula for any sub-leading pole is,

$$\frac{\exp\left(\frac{\pi\sqrt{Q^2P^2 - (Q \cdot P)^2}}{n_2}\right)}{n_2} \left[ \det(C\Omega + D)^{k+2} g(\rho)^{-1} g(\sigma)^{-1} \right]_{\text{saddle}}$$
$$(-1)^{Q \cdot P} \exp\left[ i\pi(n_1P^2 - m_1Q^2 + jQ \cdot P)/n_2 \right]$$

- To evaluate  $\det(C\Omega + D)^{k+2} g(\rho)^{-1} g(\sigma)^{-1}$ , we actually need the transformation matrix  $A, B, C, D$ .
- We will now compare the entropy with the leading pole results.

## Comparison

| $Q^2, P^2$ | $Q.P$ | $d(Q, P)$          | $S_{stat}$ | $S_{stat}^{(0)}$ | $S_{stat}^{(1)}$ | $S_{stat}^{(2)}$ | $\Delta d$ |
|------------|-------|--------------------|------------|------------------|------------------|------------------|------------|
| 2          | 0     | $5 \times 10^4$    | 10.82      | 6.28             | 10.62            | 11.58            | 34.6       |
| 4          | 0     | $3 \times 10^7$    | 17.31      | 12.57            | 16.90            | 17.38            | 480.6      |
| 6          | 0     | $1 \times 10^{10}$ | 23.51      | 18.85            | 23.19            | 23.51            | 18573      |
| 6          | 2     | $4 \times 10^9$    | 22.15      | 17.77            | 21.94            | 22.20            | 27652      |
| 6          | -2    | $2 \times 10^9$    | 21.77      | 17.77            | 21.94            | 22.20            | -          |

## Macroscopic Understanding

### Power suppressed correction

- Power suppressed corrections are identified to the  $\alpha'/g_{string}$  correction to Black Hole macroscopic entropy.
- This Entropy function is just the value of the corresponding six derivative term in the Black Hole action computed on the  $AdS_2 \times S^2$  background.
- We do not know a candidate for this term in the action. Our Analysis tells us that the term has to be duality invariant and puts a strong constraint to the possible terms.
- We have also been able to eliminate terms like  $R^3$  as they gives zero result.

## Exponentially Suppressed Correction

- These corrections naturally come from Quantum Entropy Function.

## Quantum Entropy Function

- This is a proposal for computing the exact degeneracy of states of an extremal Black Holes (arXiv:0809.3304 [hep-th]).
- These Black Holes have the following near horizon geometry.

$$ds^2 = v \left( (r^2 - 1)d\theta^2 + \frac{dr^2}{r^2 - 1} \right), \quad F_{rt}^{(i)} = -ie_i, \quad \dots$$

$v$ ,  $e_i$  are constants, .... denotes near horizon values for other fields.

- The degeneracy is given as,

$$d(\vec{q}) = \left\langle \exp[-iq_i \oint d\theta A_\theta^{(i)}] \right\rangle_{AdS_2}^{finite}$$

- where  $\langle \rangle_{AdS_2}$  denotes the unnormalized path integral over various fields of string theory on euclidean global  $AdS_2$ .
- The superscript '*finite*' refers to the finite part of the amplitude. To get this, we need to put a cut of to regularize the AdS volume.

- Explicit computation shows that this proposal reproduces the right classical degeneracy of states for quarter BPS Black Holes in  $\mathcal{N} = 4$  theories as,

$$d(q) \simeq \exp \left( \pi \sqrt{Q^2 P^2 - (Q \cdot P)^2} \right).$$

## Possible Quantum Corrections

- There are two sources of quantum corrections.
  - From fluctuations of the string fields around  $AdS_2$  background.
  - There can be different classical solutions with similar asymptotic configuration.

## Fluctuation of the Background

- The degeneracy is actually given as the finite part of the amplitude in the  $AdS_2$  background, and hence can only get power law corrections from fluctuation modes.

## Different Solutions

- The different solution can come with a different action, and hence we can get a different exponential factor.
- Q. Can we identify such a different solution?

## Solution

- We consider a  $\mathbb{Z}_N$  quotient of the previous background by,

$$\theta \rightarrow \theta + \frac{2\pi}{N}, \quad \phi \rightarrow \phi - \frac{2\pi}{N} \quad (1)$$

- The new solution looks like,

$$ds^2 = v \left( (\tilde{r}^2 - 1) d\tilde{\theta}^2 + \frac{d\tilde{r}^2}{\tilde{r}^2 - 1} \right), \quad F_{\tilde{r}\tilde{\theta}}^{(i)} = -i e_i, \quad \dots,$$

$$\tilde{\theta} \equiv \tilde{\theta} + \frac{2\pi}{N}$$

- In a new coordinate  $r = \tilde{r}/N$ ,  $\theta = N\tilde{\theta}$ , the solution looks as,

$$ds^2 = v \left( (r^2 - N^{-2})d\theta^2 + \frac{dr^2}{r^2 - N^{-2}} \right), \quad F_{r\theta}^{(i)} = -i \mathbf{e}_i, \quad \dots$$

$$\theta \equiv \theta + 2\pi, \quad \phi \rightarrow \phi - \frac{2\pi}{N}$$

- This has the same asymptotic behavior as the original solution.

- The finite contribution to the quantum entropy function is,

$$d(\vec{Q}, \vec{P}) = \exp \left( \pi \sqrt{Q^2 P^2 - (Q \cdot P)^2 / N} \right)$$

- We recover the correct exponentially sub-leading correction identifying  $N = n_2$ .

## Results

- We have shown that the degeneracy formula is valid for a generic quarter BPS dyonic Black Hole state in  $\mathcal{N} = 4$  theory.
- We have explored possible power suppressed and exponentially suppressed corrections to the microscopic degeneracy formula.
- We have also identified the roots of these corrections in the Black Hole macroscopic entropy.
- The exponentially suppressed corrections to the black hole entropy is universal, does not depend on the particular kind of extremal Black Holes.

**THANK YOU**