

India-based Neutrino Observatory (INO)

Status Report

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For the INO Collaboration

(<http://www.imsc.res.in/~ino>)

Outline of talk

- Brief overview of the current status of neutrino physics

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- The India-based Neutrino Observatory
 - Location(s)
 - The ICAL Detector: RPC's and magnet design
 - Physics possibilities at ICAL: atmospheric and long-baseline physics

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 - Physics possibilities at ICAL: atmospheric and long-baseline physics
- Other physics studies possible at INO



Neutrinos: A (Very) Brief Overview

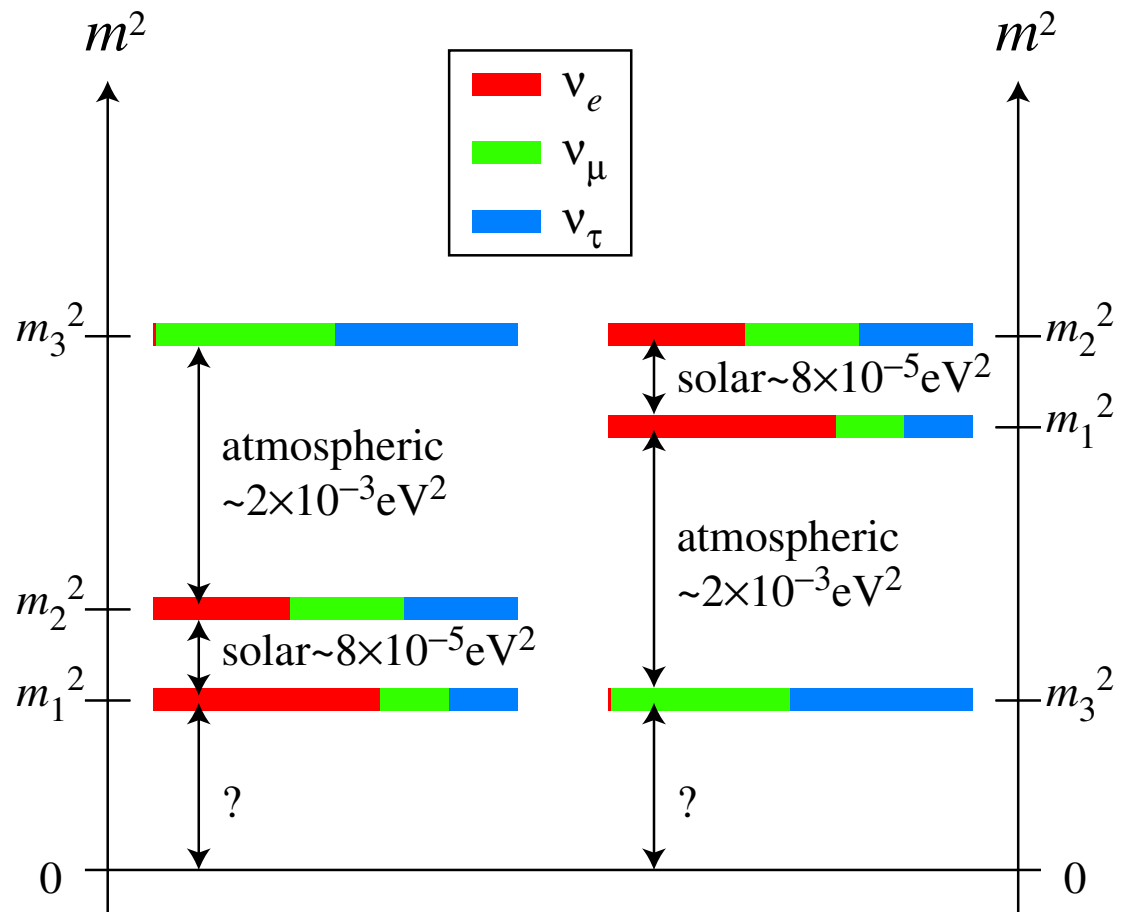
From: www.bnl.gov/

A Schematic of Neutrino Properties

Neutrino masses are not well-known. Oscillation studies only determine the **mass-squared differences**: $\Delta m_{ij}^2 = m_i^2 - m_j^2$ and the **mixing angles** θ_{ij} .

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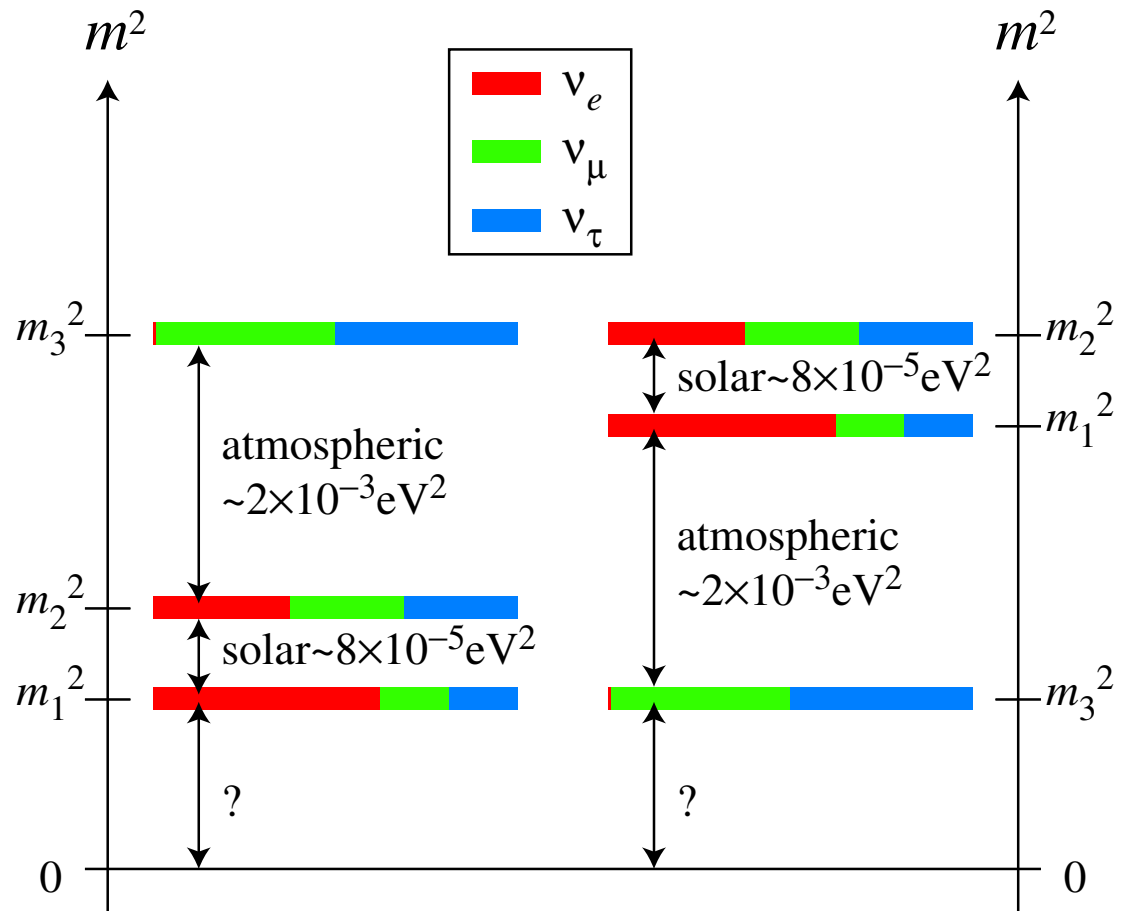
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$$\Delta m_{21}^2 \sim 0.8 \times 10^{-4} \text{ eV}^2 ;$$

$$|\Delta m_{32}^2| \sim 2.0 \times 10^{-3} \text{ eV}^2 ;$$

$$\sum_i m_i < 0.7\text{--}2 \text{ eV}.$$



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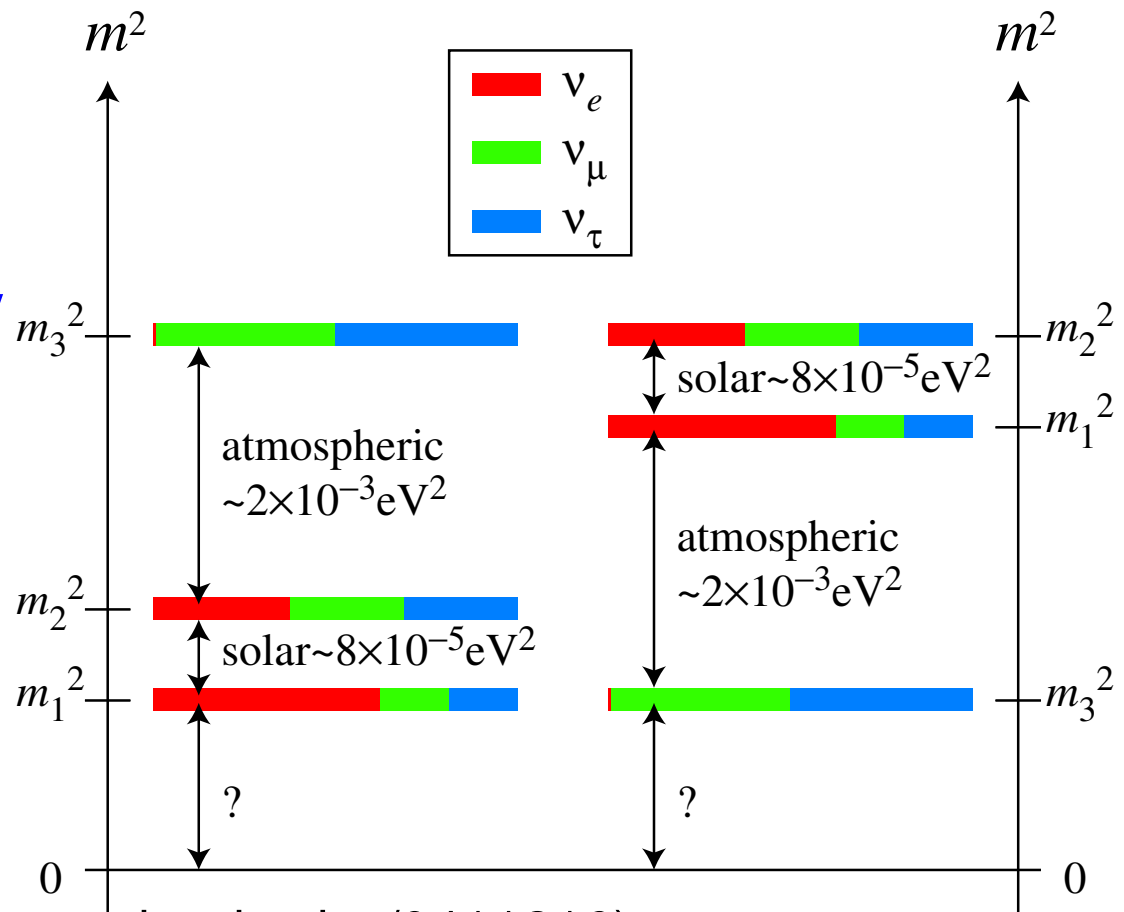
$$|\Delta m_{32}^2| \sim 2.0 \times 10^{-3} \text{ eV}^2 ;$$

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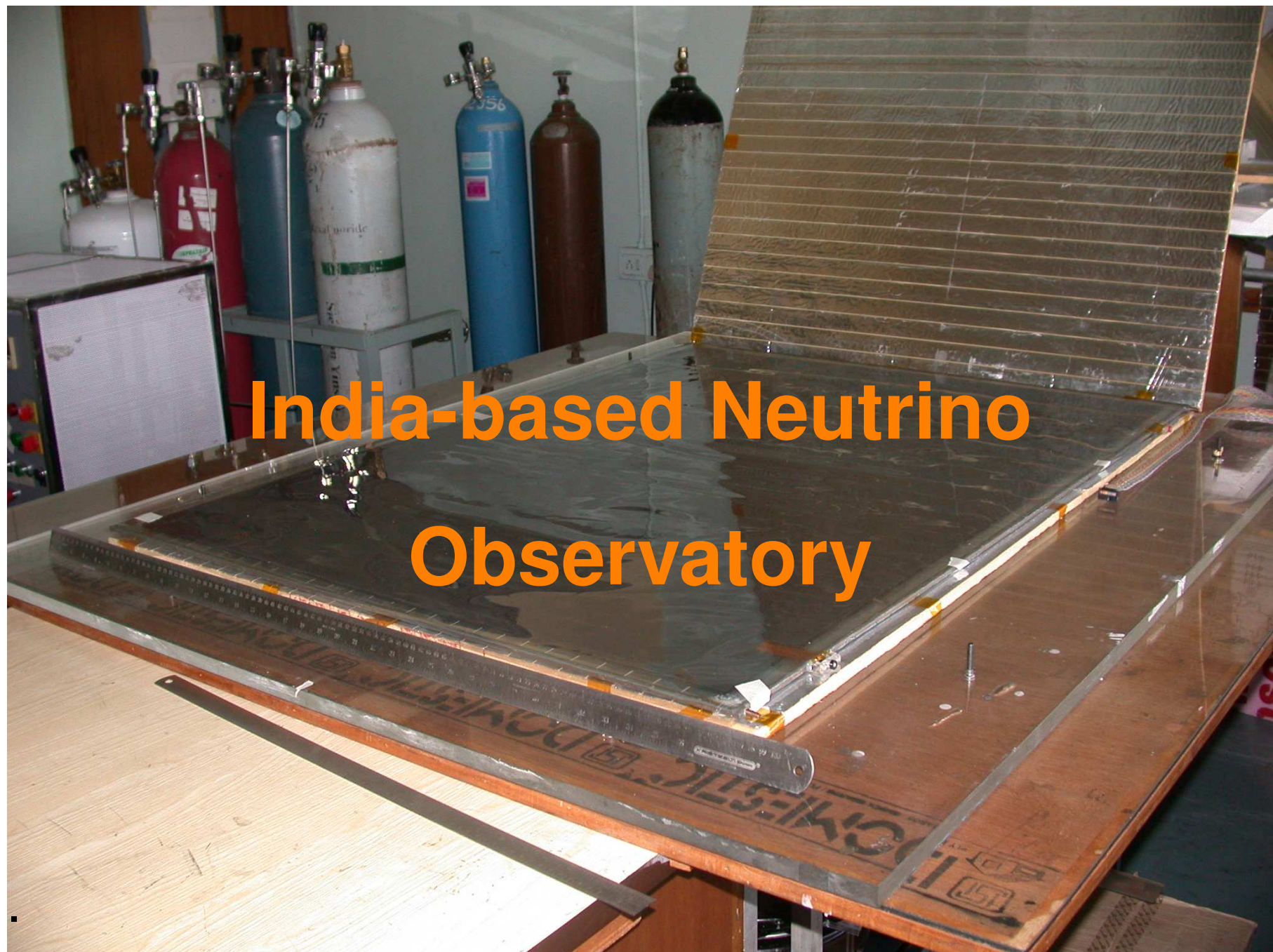
- $m_1 \sim m_2 \sim m_3 \sim 0.2 \text{ eV}$
(Degenerate hierarchy)

- $m_1 < m_2 \ll m_3$
(Normal hierarchy)

- $m_3 \ll m_1 < m_2$
Inverted hierarchy



(APS multi-divisional neutrino study, physics/0411216)



India-based Neutrino Observatory

The INO Collaboration

- Stage I : Study of atmospheric neutrinos
 - Site Survey
 - Detector R & D, including construction of a prototype (latter in progress)
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- Other detectors/physics like neutrinoless double beta decay?

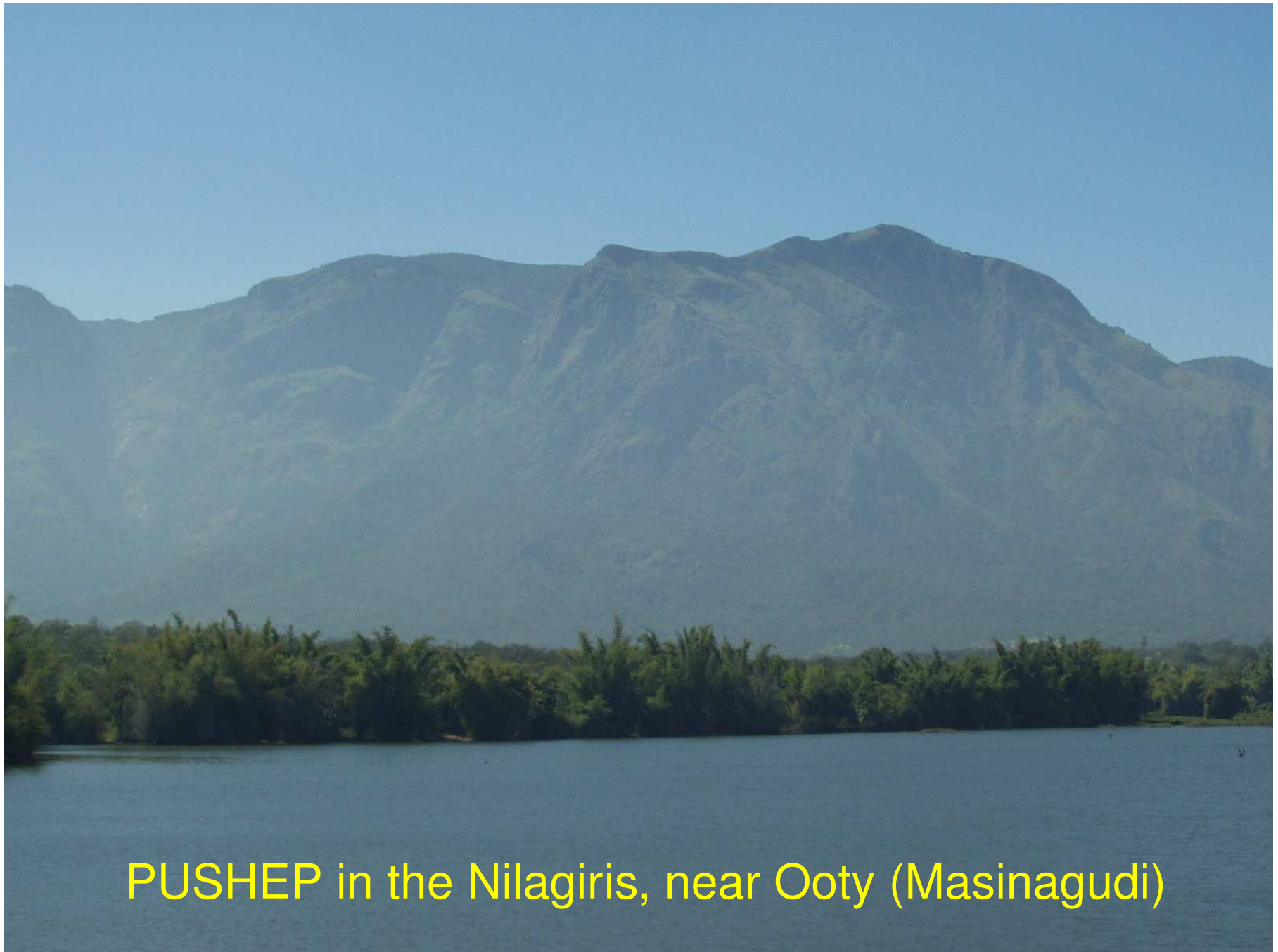
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The INO Collaboration

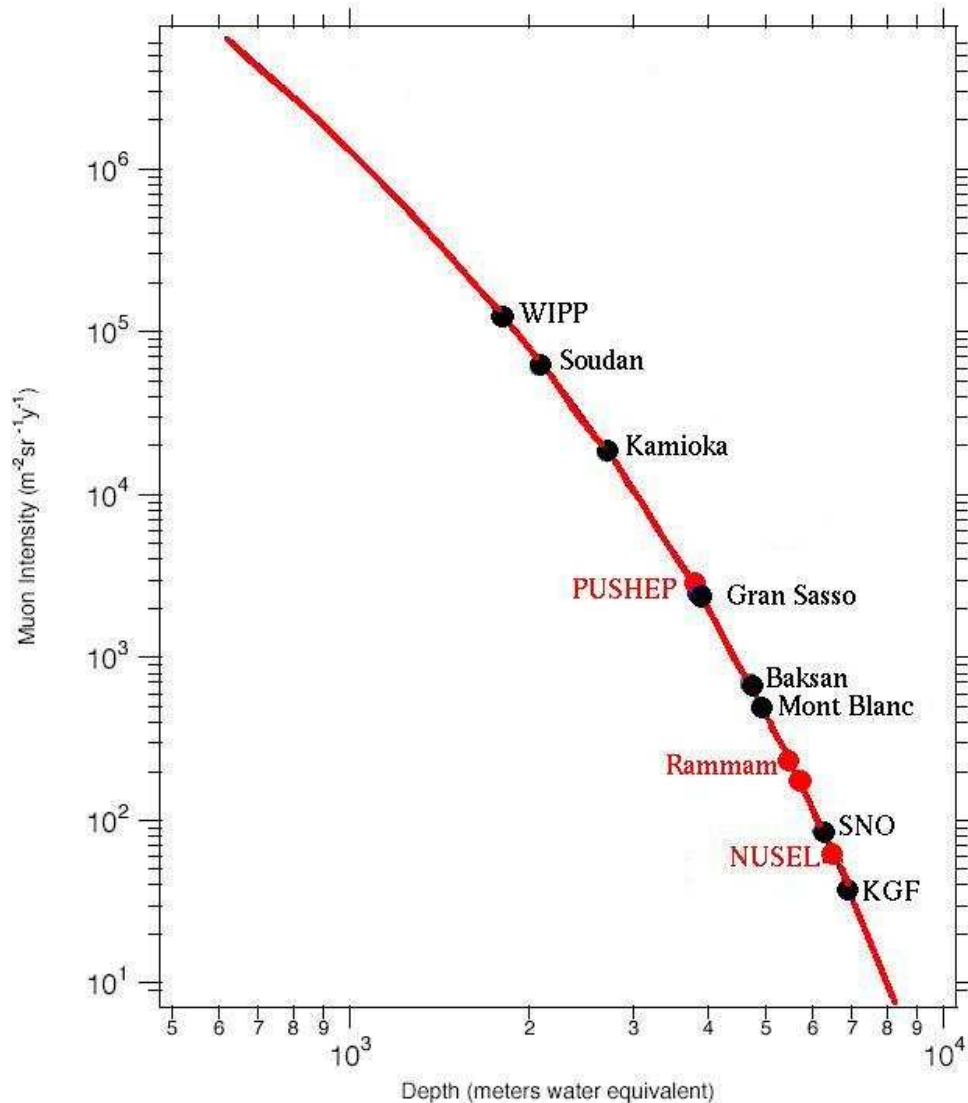
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- Stage II : Study of long-baseline neutrinos, from a neutrino factory.
- Other detectors/physics like neutrinoless double beta decay?
- Should be an international facility
- The technical review of the INO proposal is complete and is favourable. It has now been submitted to the funding agencies for approval.

Site survey: PUSHEP



PUSHEP in the Nilagiris, near Ooty (Masinagudi)

More on the site



- 2.1 km long access tunnel into mountain; cavern beneath the peak
- Experimental hall I: $25\text{m} \times 130\text{m} \times 30\text{m}$ (height) built to accommodate 50 kton + 50 kton modules (future expansion)
- Experimental Hall II: about half the size, to accommodate other, smaller experiment(s).

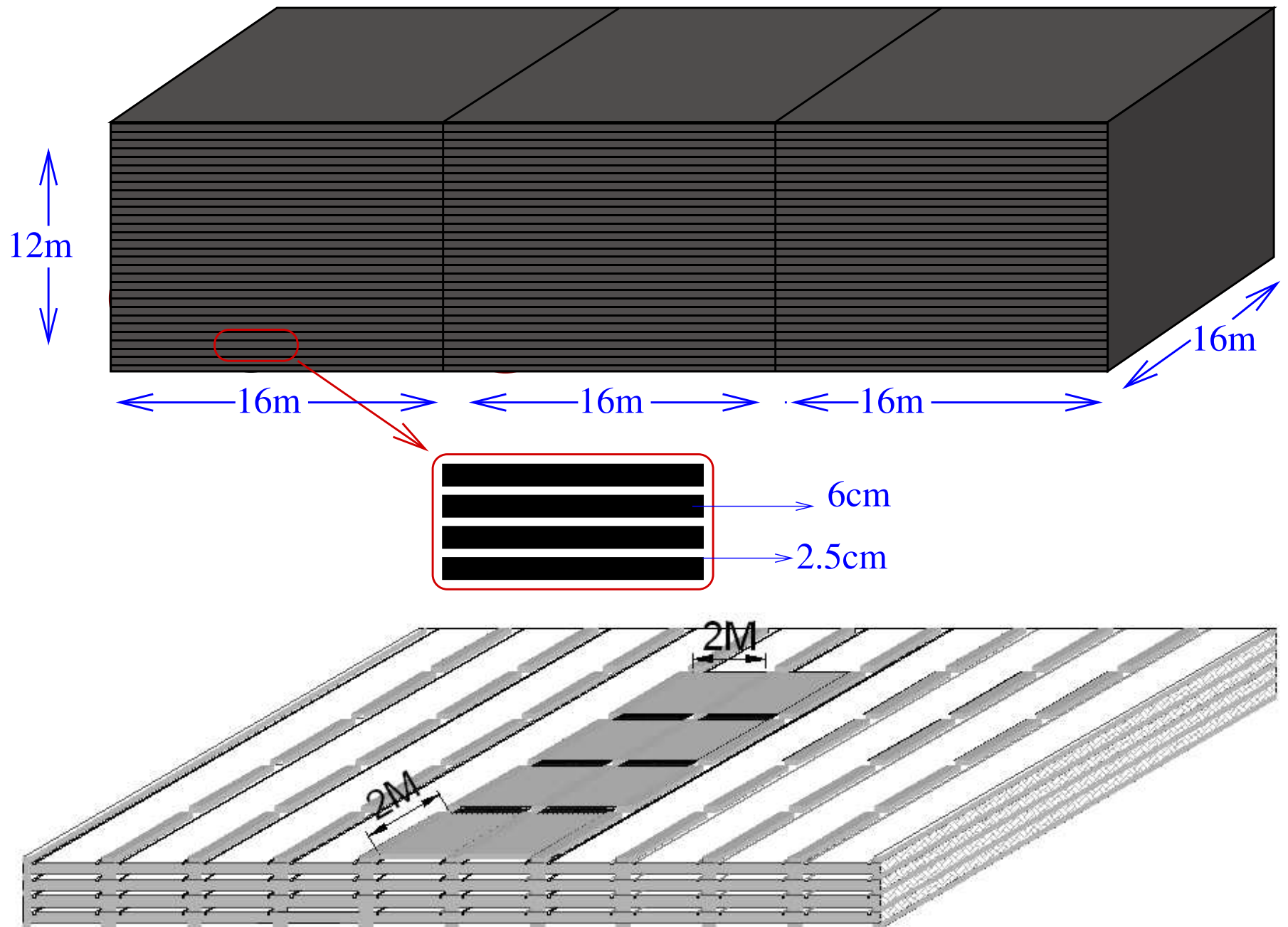
The choice of detector

- Large target mass: began with 30 kton; current design 50 kton
- Good tracking and energy resolution
- Nano-second time resolution for up/down discrimination; hence good directionality
- Good charge resolution
- Ease of construction (modular)

Use (magnetised) iron as target mass and RPC as active detector element. Similar to MONOLITH.

Note: Is sensitive to muons only, not electrons

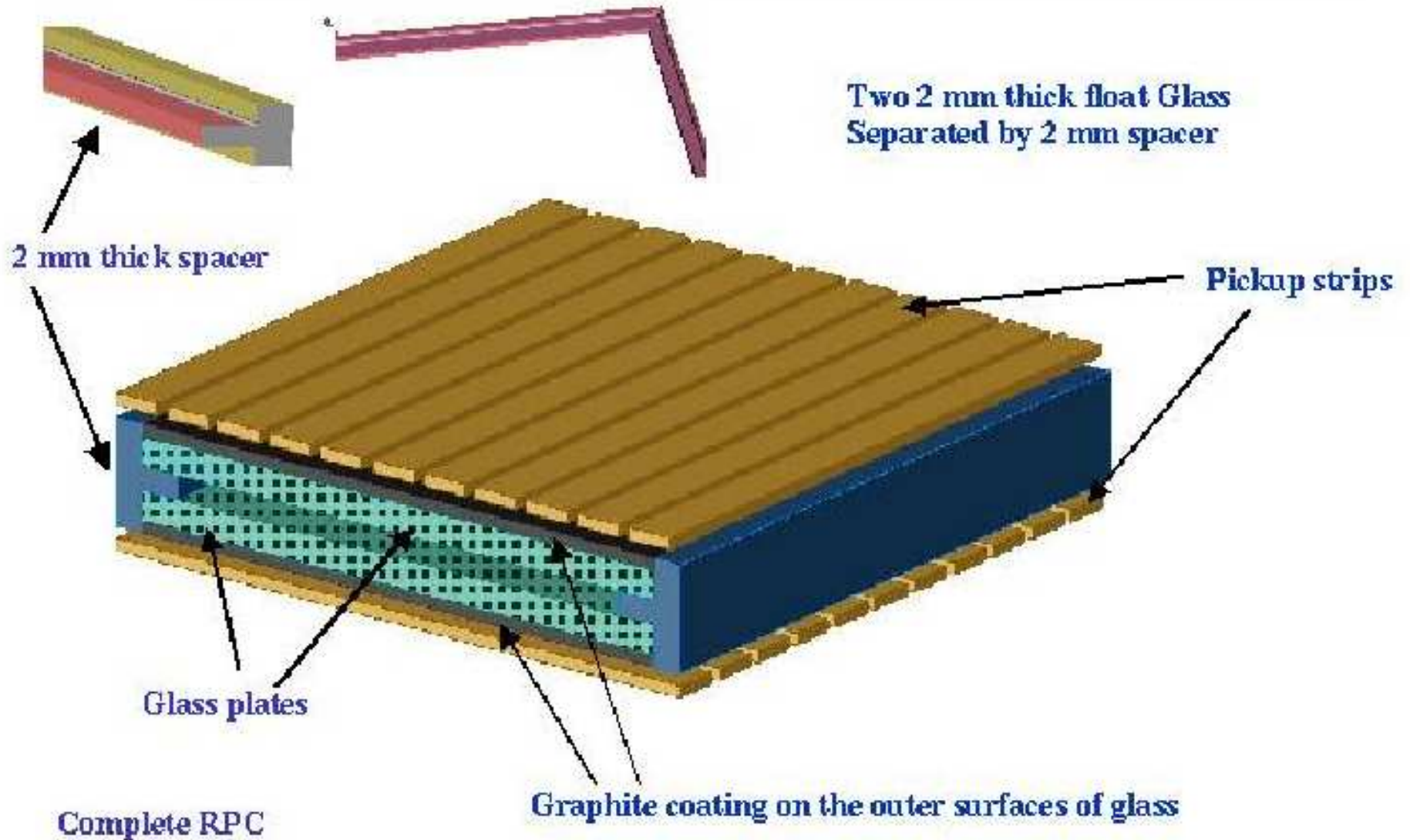
The ICAL detector



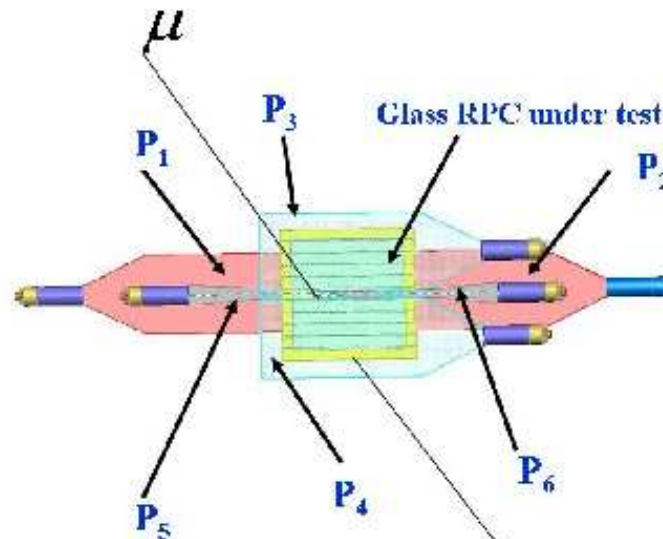
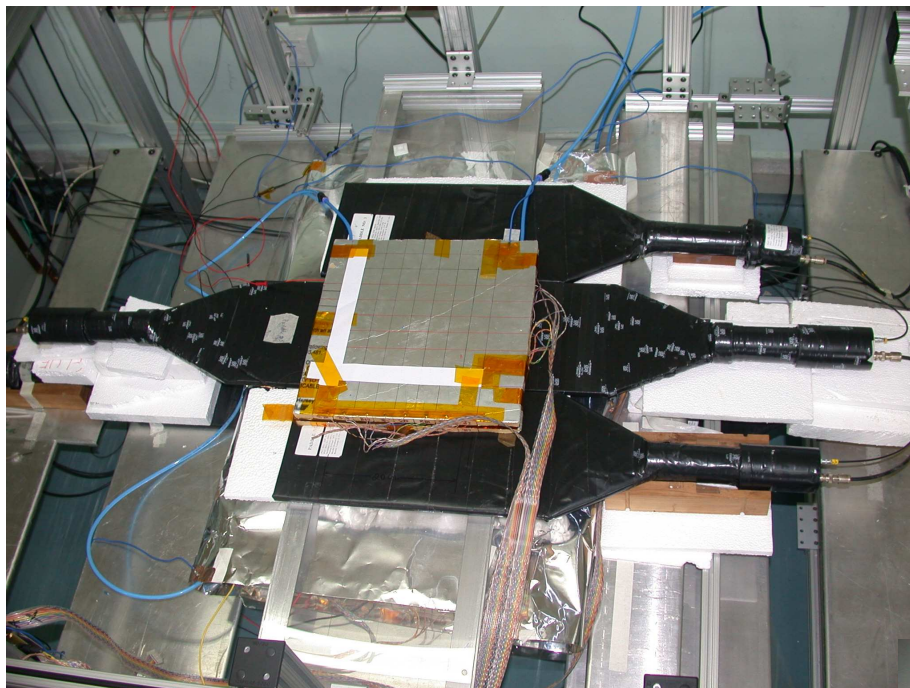
The active detector elements: RPC

RPC Construction:

Float glass, graphite, and spacers



Fabricating RPC's



Initially: 30 cm \times 30 cm

Currently: 1.0 m \times 0.9 m

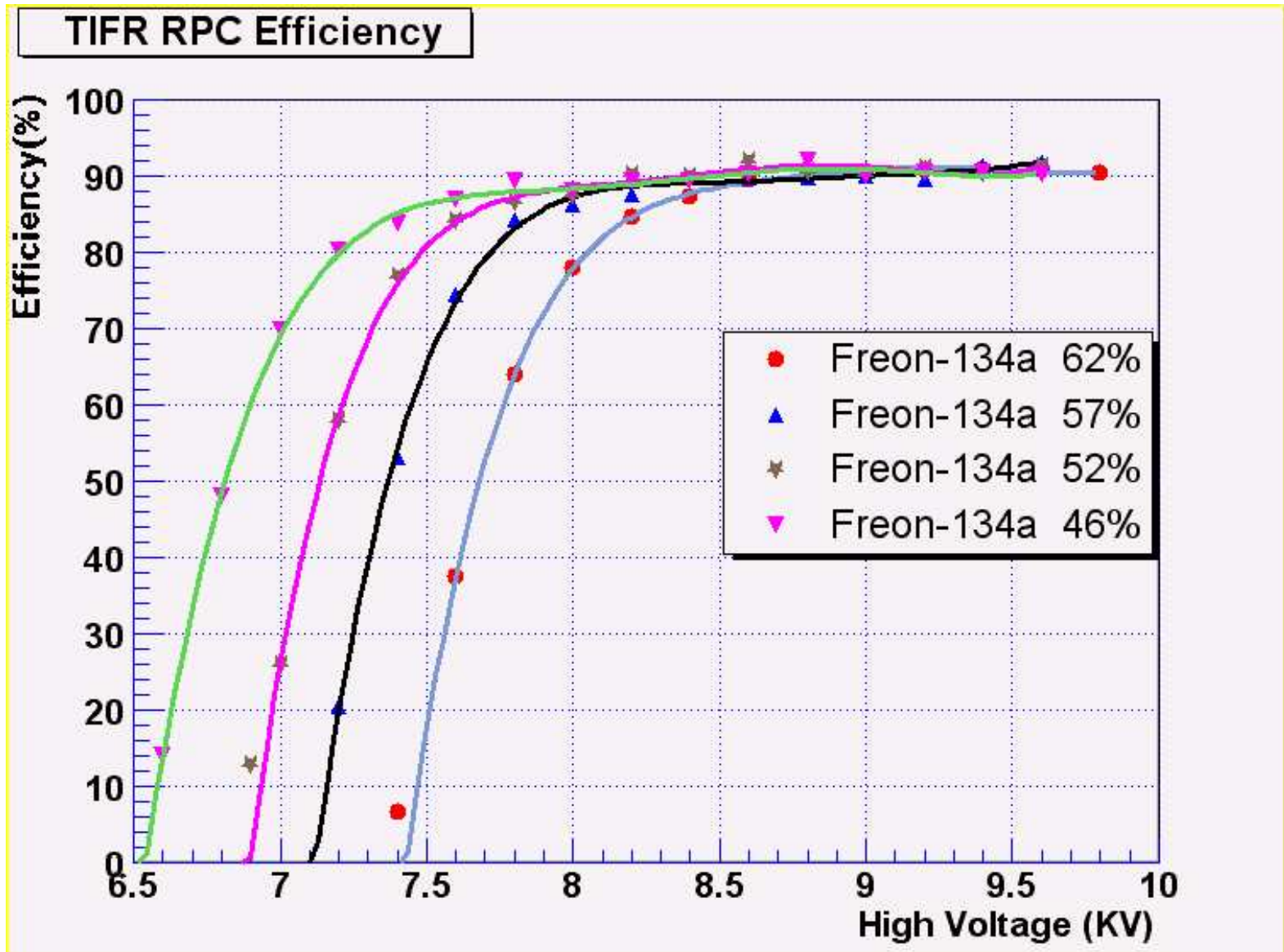


Specifications of the ICAL detector

ICAL	
No. of modules	3
Module dimension	16 m × 16 m × 12 m
Detector dimension	48 m × 16 m × 12 m
No. of layers	140
Iron plate thickness	~ 6 cm
Gap for RPC trays	2.5 cm
Magnetic field	1.3 Tesla
RPC	
RPC unit dimension	2 m × 2 m
Readout strip width	3 cm
No. of RPC units/Road/Layer	8
No. of Roads/Layer/Module	8
No. of RPC units/Layer	192
Total no. of RPC units	~ 27000
No. of electronic readout channels	3.6×10^6

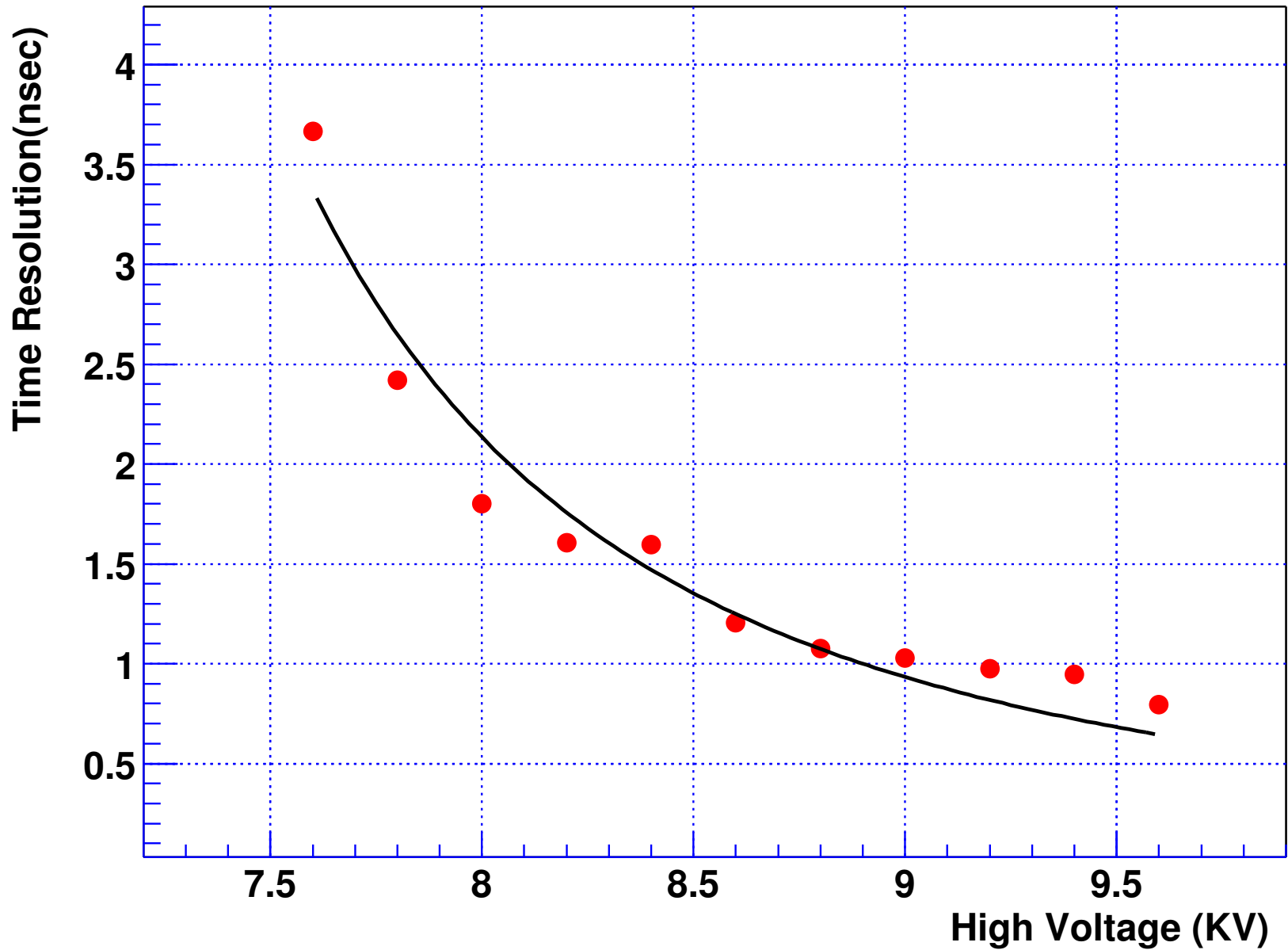
RPC Efficiency studies

Using different combinations of gas



RPC Time resolution

Time Resolution

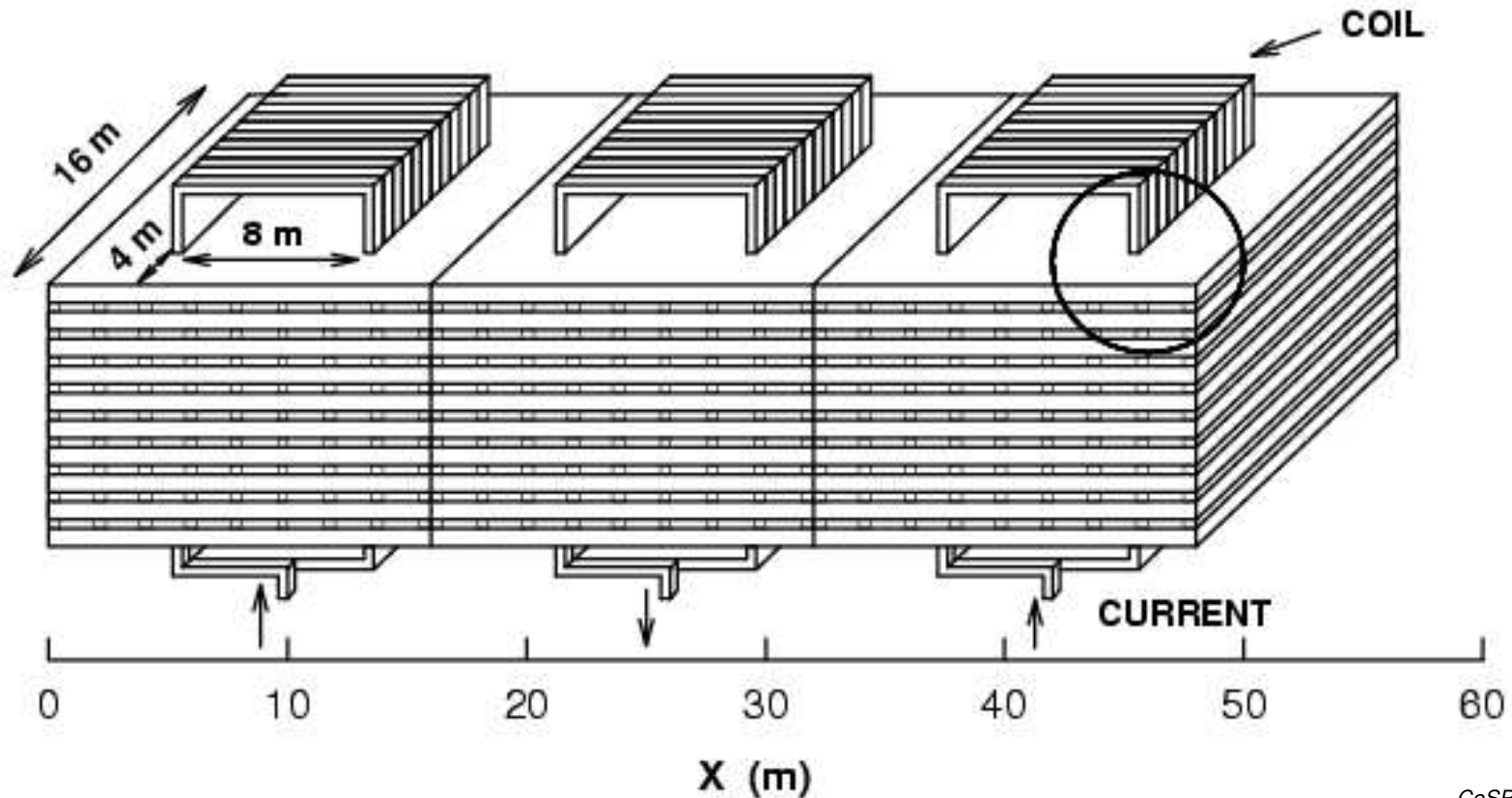


Magnet studies

Design criteria:

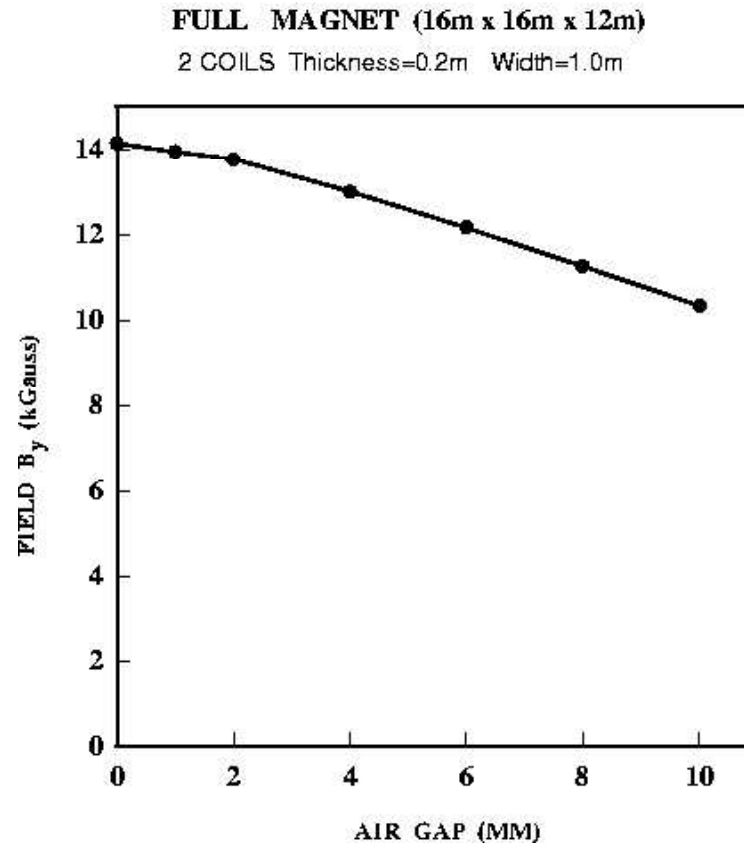
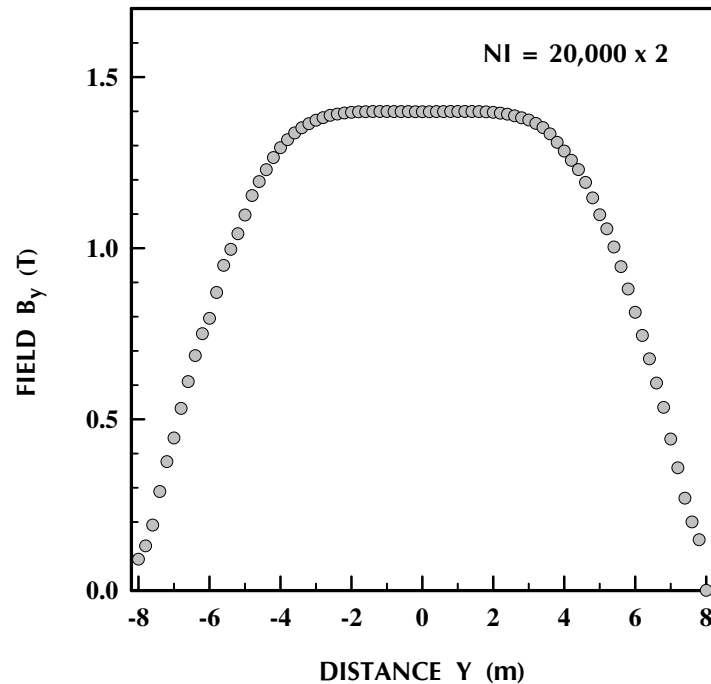
- Field uniformity
- Modularity
- Optimum copper-to-steel ratio
- Access for maintenance

Toroidal Magnet design



Magnetic field simulation

Field when there is a gap in a plate (dividing it along x-axis)

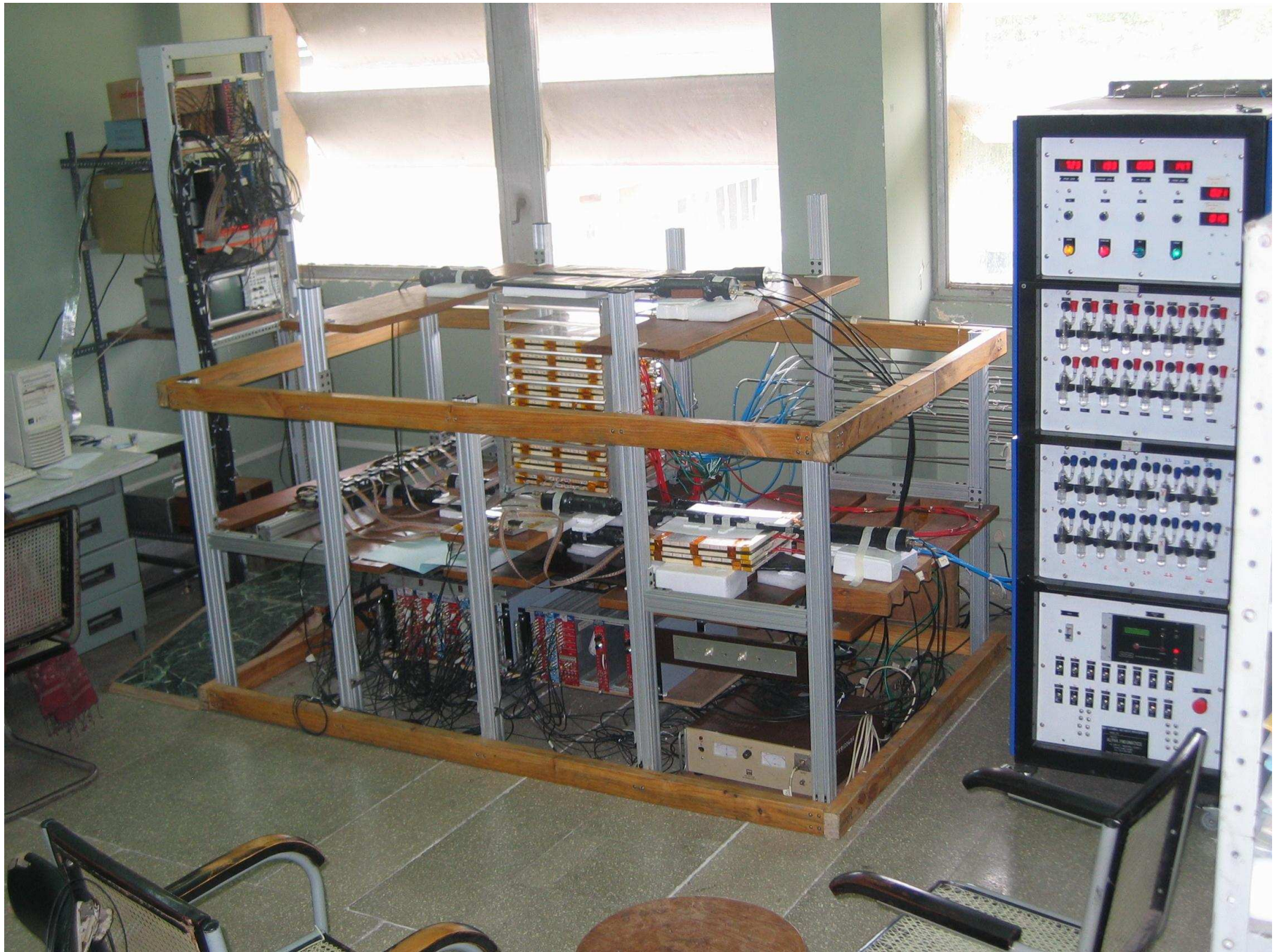


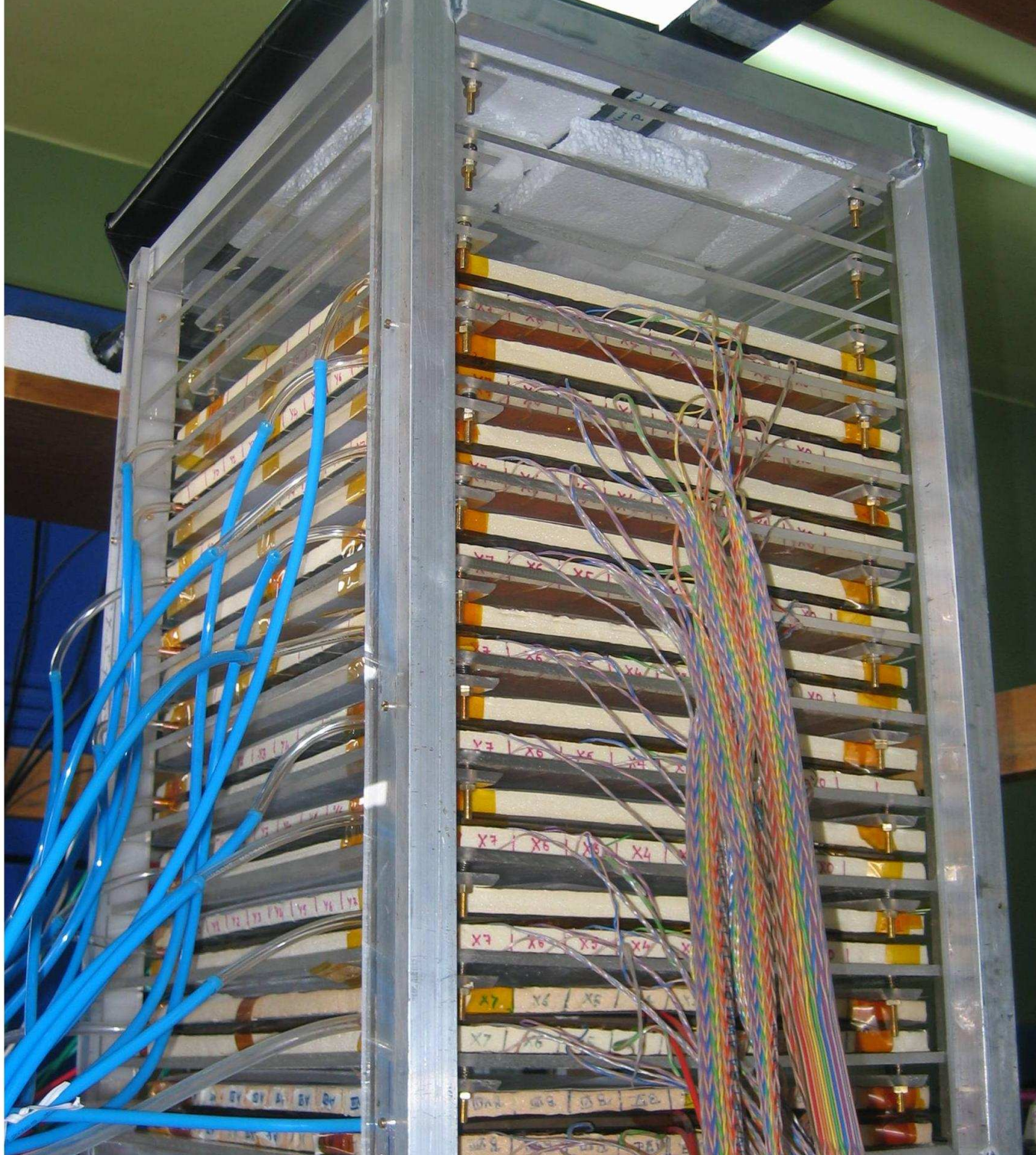
Field in x -direction uniform to within 0.25%.

Field in z -direction uniform except close to edges.

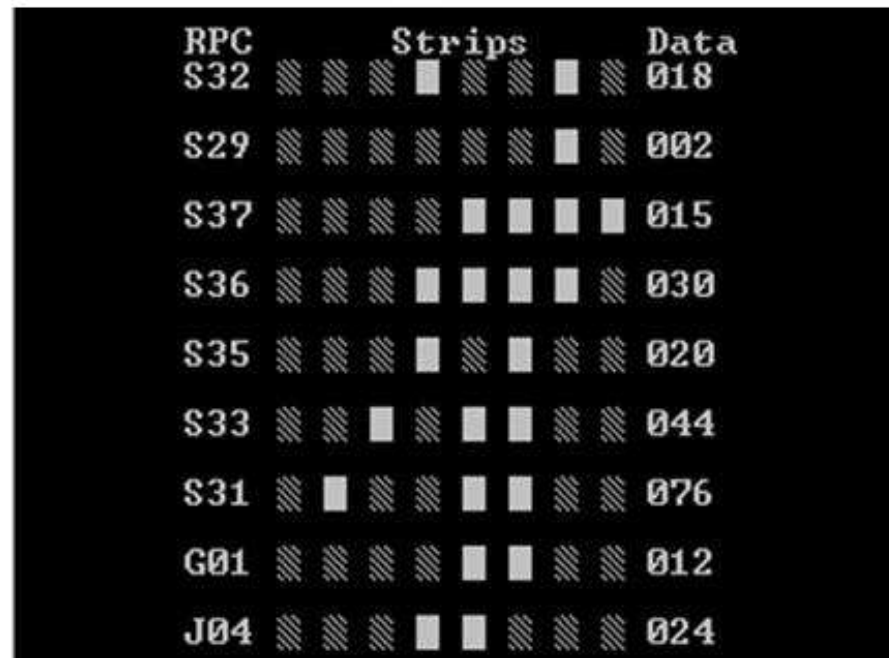
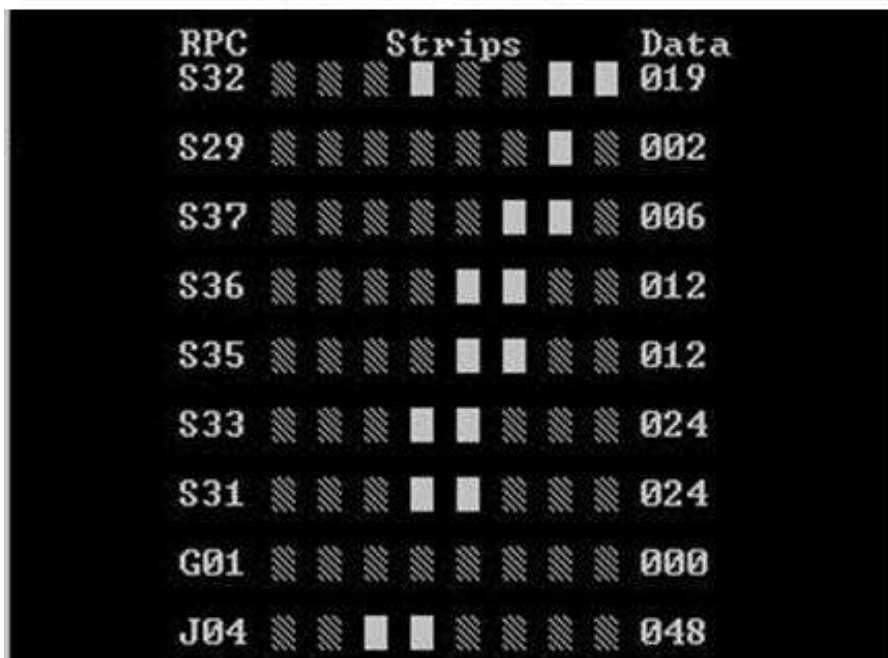
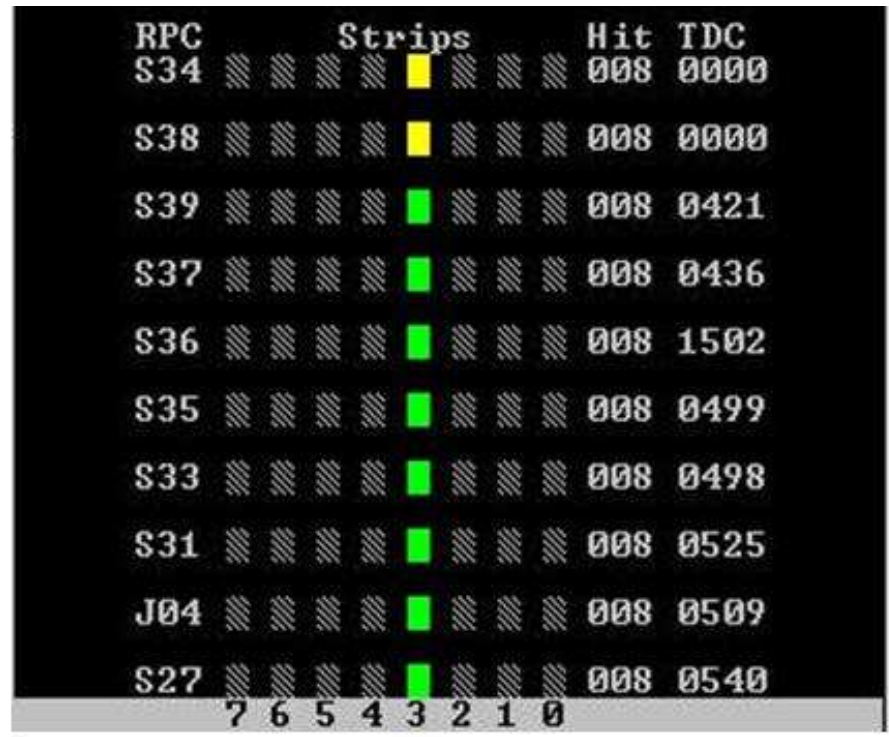
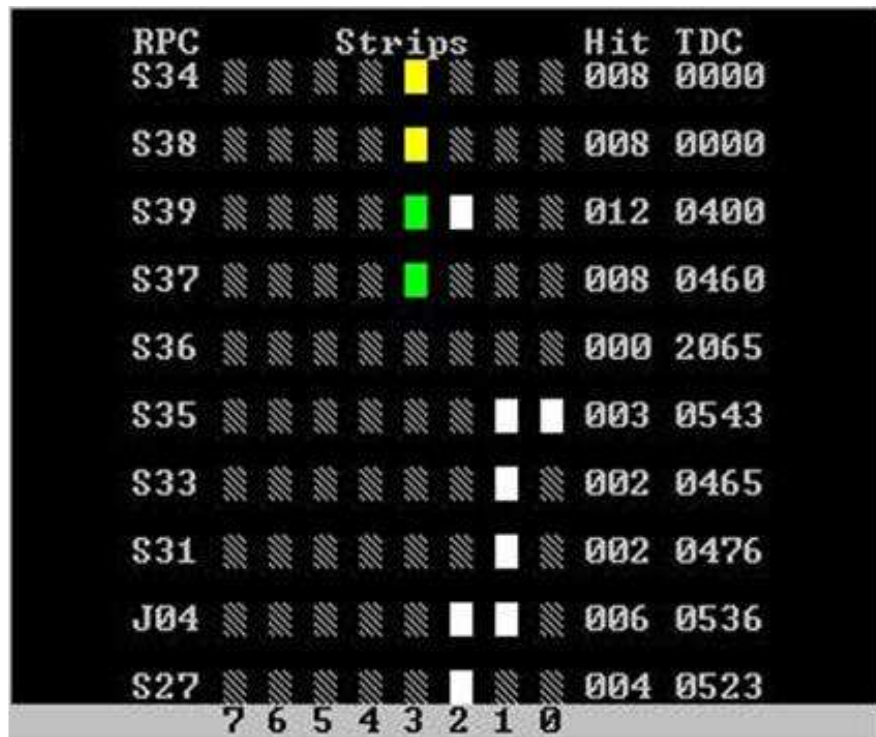
Cannot tolerate more than 4 mm gap in plate welding.

For the prototype, at TIFR ...



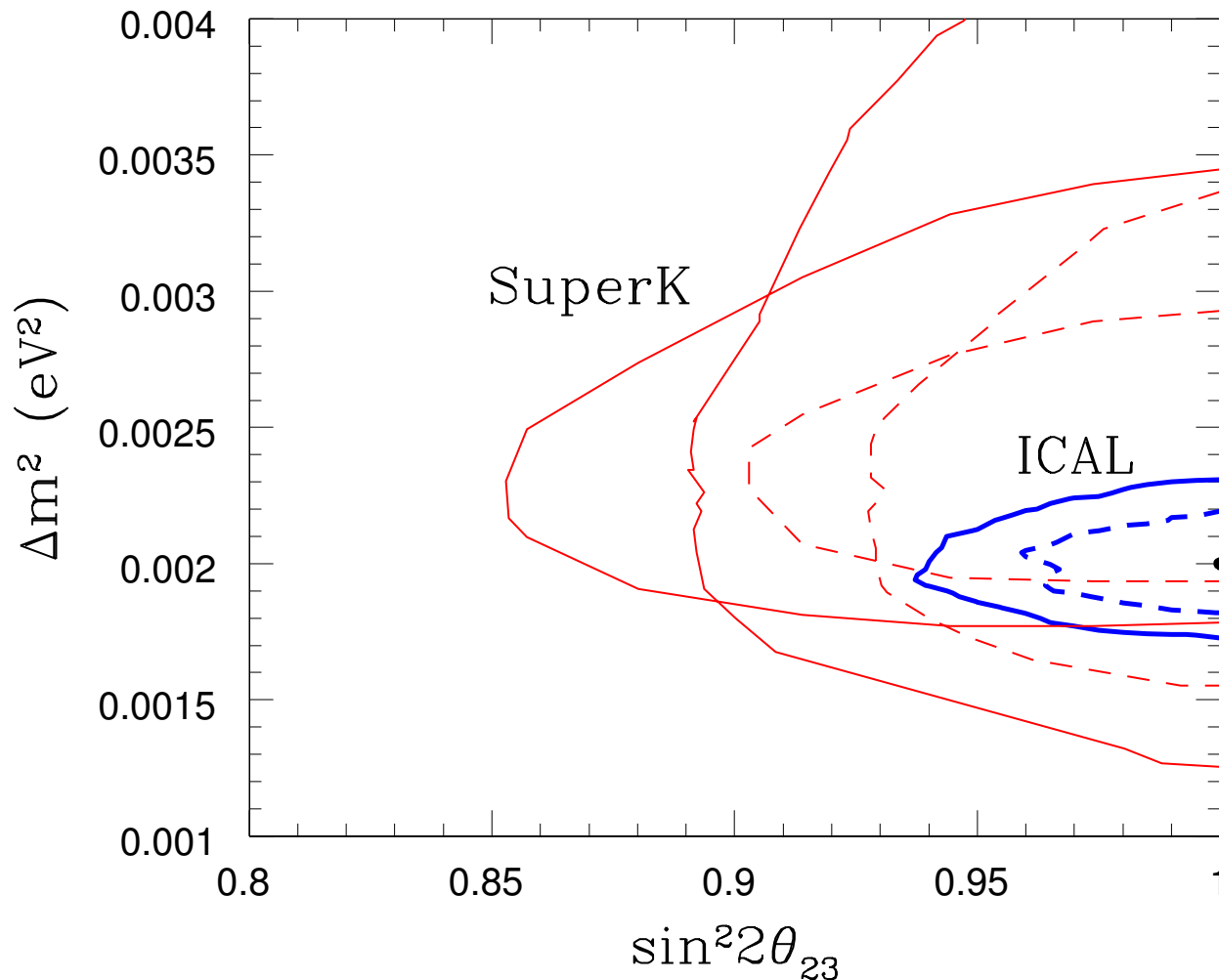


Tracks from atmospheric muons at TIFR



Event Simulations

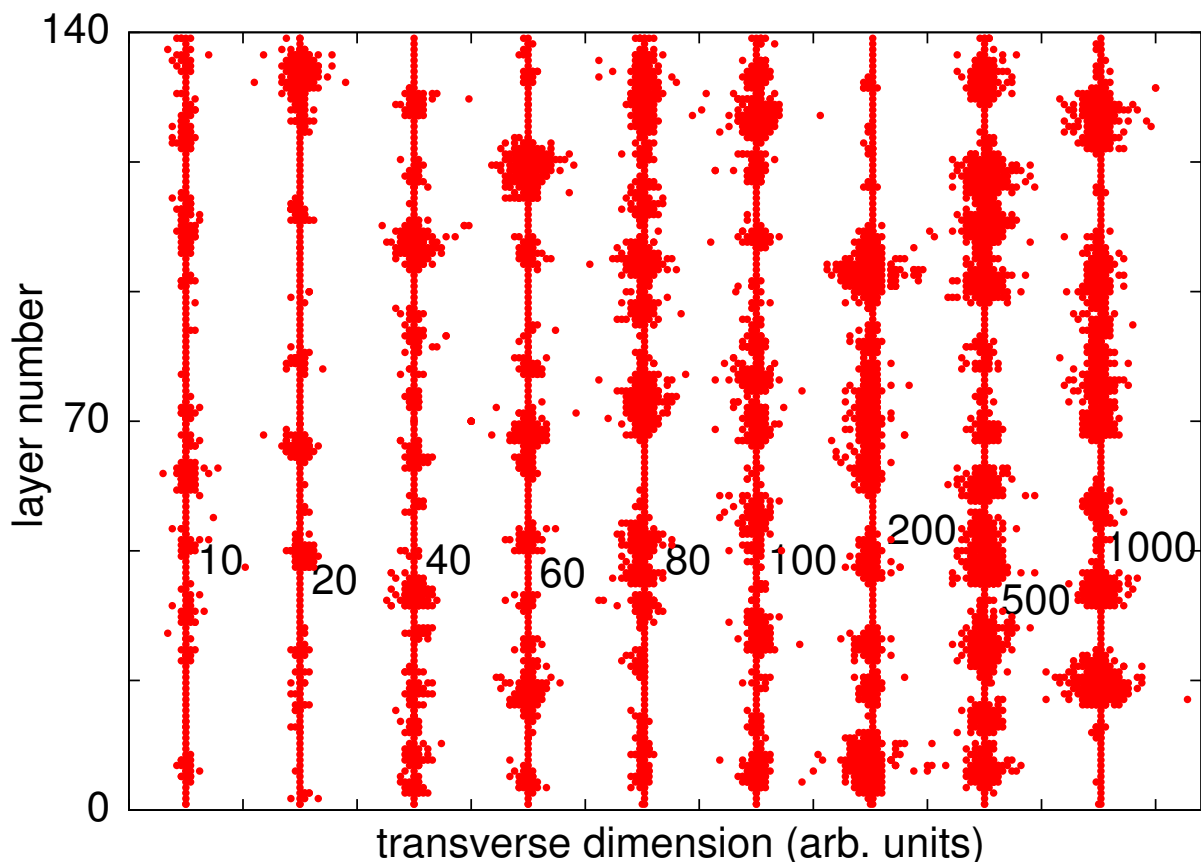
- Source: Atmospheric Neutrinos, 6 years' exposure, from Nuance neutrino generator.
- ICAL simulation with GEANT, $B_y = 1$ T.



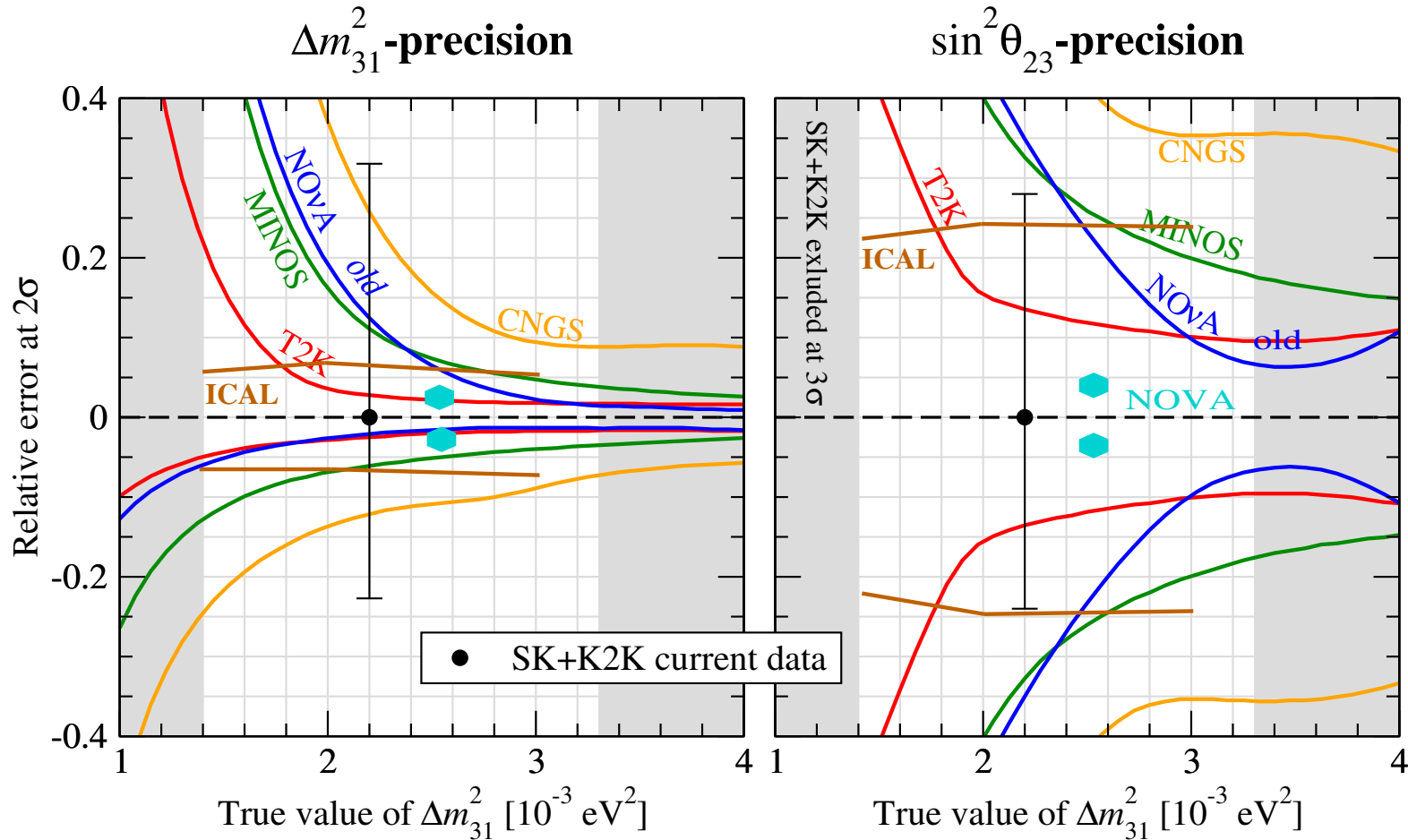
Shown are 90 and 99 CL contours in comparison with Super-K zenith angle as well as L/E results

Event Simulations II

- Source: Cosmic ray muons, both as background to neutrino events and high energy muons as events
- ICAL simulation of vertical upward TeV energy muons with GEANT, using 1 Tesla uniform magnetic field in the y -direction.



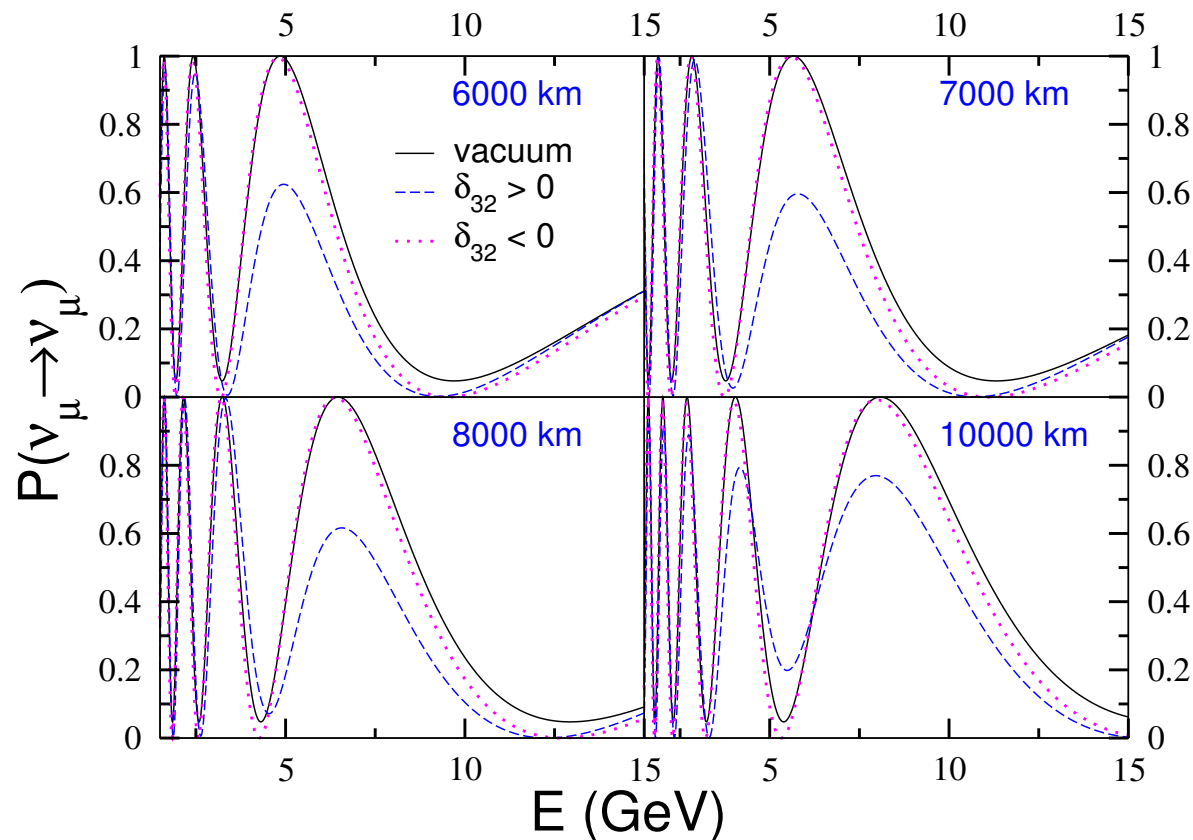
Physics Studies with ICAL



All experiments with 5 years' running; new NOVA 25kton, 6 years (6×10^{21} pot).

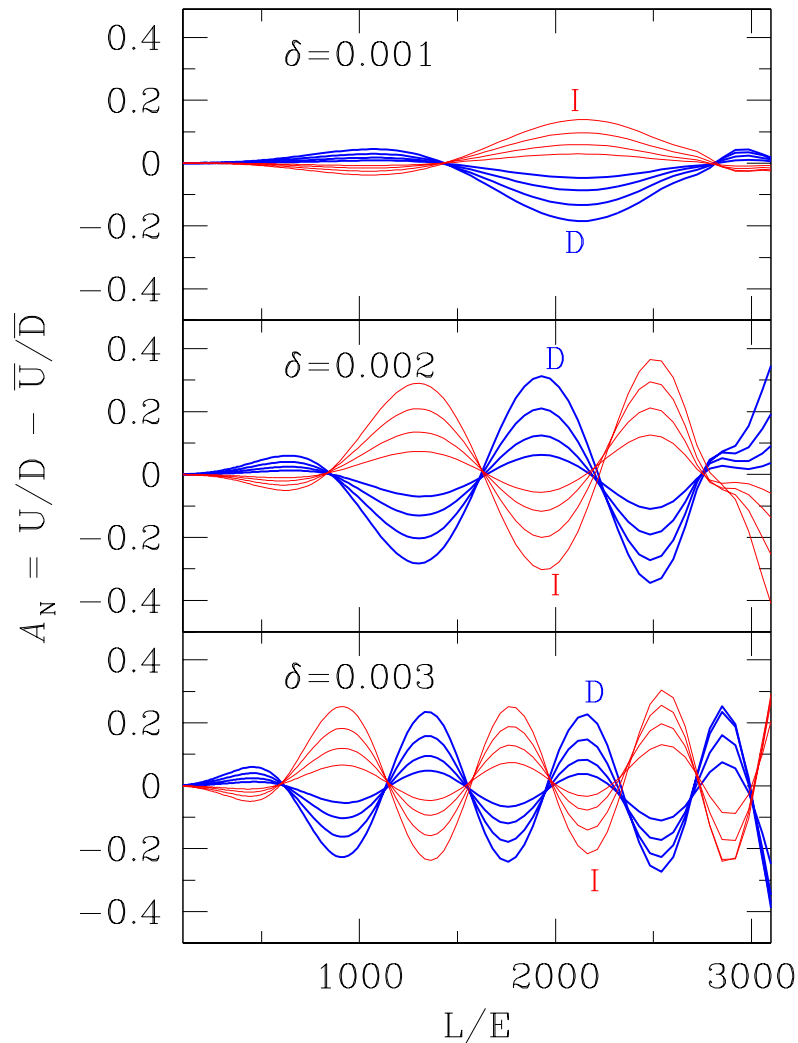
Adapted from: P. Huber, M. Lindner, M. Rolinec, T. Schwetz and W. Winter, hep-ph/0412133.

Matter effects with atmospheric neutrinos



- Matter effects involve the participation of all three (active) flavours; hence involves both $\sin \theta_{13}$ and the CP phase δ .
- Hence sensitive to the mass ordering of the 2–3 states, provided $\theta_{13} > 6^\circ$; however, needs large exposures

The difference asymmetry



D: Direct/normal; **I:** Inverted hierarchy

Sign of $\delta \equiv \Delta m_{32}^2$ for $\theta_{13} = 5, 7, 9, 11^\circ$

Hence sensitive to the mass ordering (red vs blue) of the 2–3 states

With exposures of 500 kton-years, can get a 90%CL result if

$$\sin^2 2\theta_{13} > 0.09 \text{ (10\% R)}$$

$$\sin^2 2\theta_{13} > 0.07 \text{ (5\% R)}$$

However, needs large exposures of about 800 kton-years for smaller

$$\sin^2 2\theta_{13} > 0.07 \text{ (10\% R)}$$

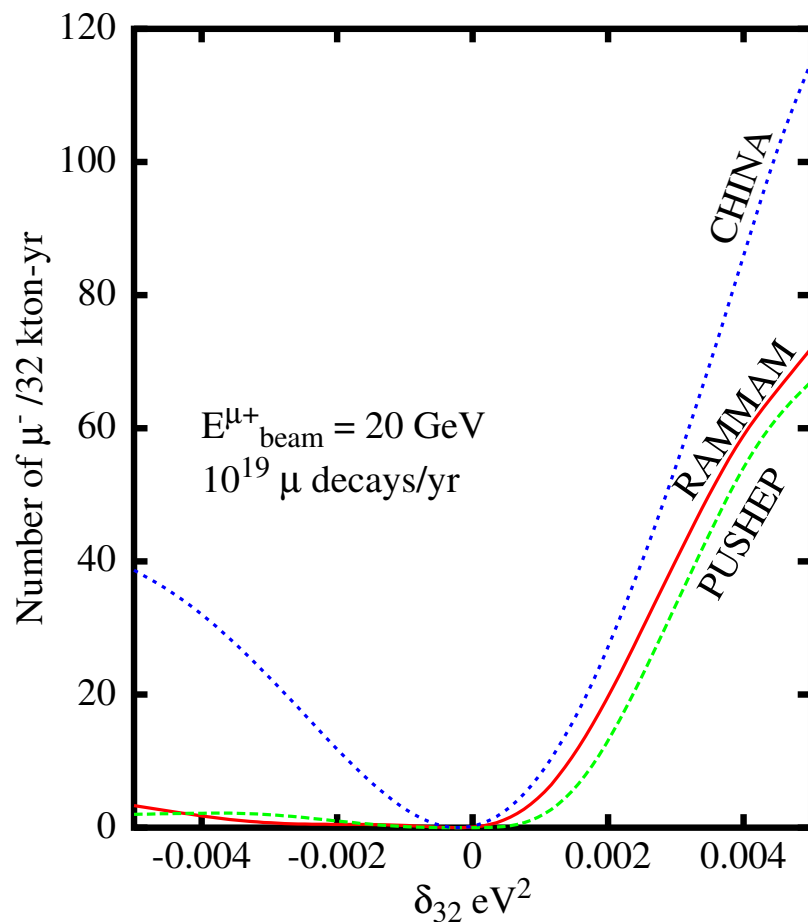
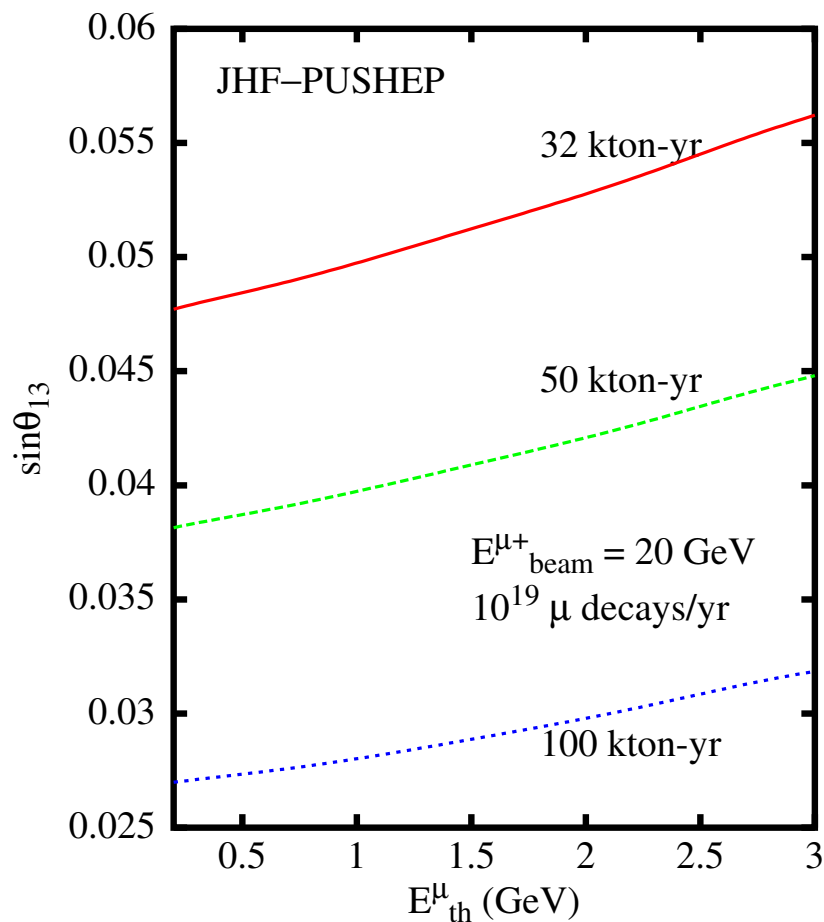
$$\sin^2 2\theta_{13} > 0.05 \text{ (5\% R)}$$

Other physics possibilities

... with atmospheric neutrinos

- **Discrimination of octant of θ_{23} provided $\theta_{13} > 7^\circ$** ($\sin^2 2\theta_{13} > 0.06$); harder than mass ordering
- **Probing CPT violation** from rates of neutrino- to rates of anti-neutrino events in the detector: sensitive to δb , which adds to $\Delta m_{32}^2/(2E)$ in oscillation probability expression.
- **Constraining long-range leptonic forces** by introducing a matter-dependent term in the oscillation probability even in the absence of U_{e3} , so that neutrinos and anti-neutrinos oscillate differently.
- **Discrimination between oscillation of ν_μ to active ν_τ and sterile ν_s** from up/down ratio in “muon-less” events?

Stage II: Neutrino factories and INO



θ_{13} reach and sign of Δm_{32}^2 vs wrong sign μ

Can also study CP violation: note, JHF-PUSHEP (6556 km) and CERN-PUSHEP (7145 km) are close to magic.

In short . . .

The outlook looks good! This is a massive project:

Looking for active collaboration both within India and abroad

The INO Collaboration¹

- **Aligarh Muslim University, Aligarh:**
M. Sajjad Athar, Rashid Hasan, S. K. Singh
- **Banaras Hindu University, Varanasi:**
B. K. Singh, C. P. Singh, V. Singh
- **Bhabha Atomic Research Centre (BARC), Mumbai:**
V. Arumugam, Anita Behere, M. S. Bhatia, V. B. Chandratre, R. K. Choudhury,
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A. W. Matkar, P. K. Mukhopadhyay, S. C. Ojha², L. M. Pant, K. Srinivas
- **Calcutta University (CU), Kolkata:**
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- **Delhi University (DU), Delhi:**
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- **Harish Chandra Research Institute (HRI), Allahabad:**
Sanjib K. Agarwalla, Sandhya Choubey, Anindya Datta, Raj Gandhi, Pomita Ghoshal,
Srubabati Goswami, Poonam Mehta, Sukanta Panda, S. Rakshit, Amitava Raychaudhuri
- **University of Hawaii (UHW), Hawaii:**
Sandip Pakvasa
- **Himachal Pradesh University (HPU), Shimla:**
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- **Indian Institute of Technology, Bombay (IITB), Mumbai:**
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- **The Institute of Mathematical Sciences (IMSc), Chennai:**
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- **Institute of Physics (IOP), Bhubaneswar:**
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- **North Bengal University (NBU), Siliguri:**
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- **Panjab University (PU), Chandigarh:**
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- **Physical Research Laboratory (PRL), Ahmedabad:**
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- **Saha Institute of Nuclear Physics (SINP), Kolkata:**
Sudeb Bhattacharya, Suwendu Bose, Sukalyan Chattopadhyay, Ambar Ghosal, Asimananda Goswami, Kamales Kar,
Debashish Majumdar, Palash B. Pal, Satyajit Saha, Abhijit Samanta,
Abhijit Sanyal, Sandip Sarkar, Swapan Sen, Manoj Sharan
- **Sikkim Manipal Institute of Technology, Sikkim:**
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- **Tata Institute of Fundamental Research (TIFR), Mumbai:**
B. S. Acharya, Sudeshna Banerjee, Sarika Bhide, Amol Dighe, S. R. Dugad, P. Ghosh,
K. S. Gothe, S. K. Gupta, S. D. Kalmani, N. Krishnan, Naba K. Mondal, P. Nagaraj,
B. K. Nagesh, Biswajit Paul, Shobha K. Rao, A. K. Ray, L. V. Reddy,
B. Satyanarayana, S. Upadhyaya, Piyush Verma
- **Variable Energy Cyclotron Centre (VECC), Kolkata:**
R. K. Bhandari, Subhasish Chattopadhyay, Premomay Ghosh, B. Mohanty,
G. S. N. Murthy, Tapan Nayak, S. K. Pal, P. R. Sarma, R. N. Singaraju, Y. P. Viyogi

Scientific Steering Committee

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Ramanath Cowsik, *Indian Institute of Astrophysics, Bangalore*
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Probir Roy, *Tata Institute of Fundamental Research, Mumbai*
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¹This is an open collaboration and experimentalists are especially encouraged to join.

²since retired

³Replacing Abdul Salam who was a member until March 5, 2005